

EVOLUTION OF MAGMATIC VOLATILES  
DURING DRILLING INTO A MAGMA BODY,  
KRAFLA ICELAND

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# Abstract

In 2009, the Iceland Deep Drilling Project IDDP-1 borehole unexpectedly intercepted a silicic magma body beneath Krafla volcano at a depth of approximately 2.1 km. Glass samples quenched in the drilling mud were directly sampled at ~15 minute intervals during drill circulation following magma interception at 15:17, providing a unique record of the changing state of the magmatic system over this period. The nature of the glass cuttings varied through time, with a decrease in the proportion of vesicular cuttings and an increase in dense glassy cuttings. The recovered glass cuttings provide a unique opportunity to investigate the magmatic response to a sudden decompression event, and in particular the response of magmatic volatiles over time. During decompression, volatiles diffuse through the magma and exsolve into growing bubbles. Conversely, pressure increases or temperature decreases can cause volatile species to diffuse back from bubbles into the surrounding melt. These processes create distinct diffusion gradients in magmatic glass surrounding the bubbles, and hence the time-dependent volatile response to magmatic decompression is uniquely preserved by the quenched IDDP-1 fragments.

For this project an Electron Probe Micro-analyser was used to determine the major element compositions of each sample, while Fourier transform infrared spectroscopy (FTIR) was used to generate maps of H<sub>2</sub>O and CO<sub>2</sub> absorbance across broad areas. Average H<sub>2</sub>O and CO<sub>2</sub> concentrations determined from the volatile maps are used to reconstruct the time-dependent evolution of magmatic H<sub>2</sub>O and CO<sub>2</sub> at the magma interface. In addition, transects around bubbles were extracted to examine whether H<sub>2</sub>O and/or CO<sub>2</sub> diffusion profiles developed, and if so how they responded with time.

Results from the major element analyses suggest that the clear glass represents the boundary layer between the magma body and its surrounding rock, some of which appears to have been assimilated at the magma-rock interface. The brown

glass cuttings (Melt-1b) appear to more closely represent the main magma body. Volatile results show that the concentration of total H<sub>2</sub>O did not change over time; however, OH/H<sub>2</sub>O<sub>m</sub> and CO<sub>2</sub> both increase over time after initial intersection. These results are consistent with an increase in temperature that is interpreted to reflect deeper hotter magma rising into the borehole. The volatile distributions around bubbles show both resorption and bubble growth characteristics. This differs from an expected simple decompression (i.e. producing bubble growth) and suggests that the magma maintained a somewhat heterogeneous condition prior to quenching. This may reflect the relatively small borehole size, and together with the fact that no eruption was triggered suggests that very small decompression events can be moderated by melt immediately surrounding the region of low pressure.

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# Chapter 1: Introduction

## 1.1 Project background

In 2009, the Iceland Deep Drilling Project (IDDP-1) borehole unexpectedly intersected a silicic (76.5 % SiO<sub>2</sub>) magma body at a depth of 2104 m below the surface (Elders et al. 2011; Zierenberg et al. 2013; Trewick et al. 2016). When IDDP-1 drilled into the magma body there was no eruption, although magma did apparently rise up the borehole. The lack of eruptions provides research opportunities for future geothermal energy development (Fridleifsson & Elders, 2014) and raise the possibility of drilling directly into magma bodies to efficiently generate cheap geothermal energy. However, engineering, safety and geological studies need to be fully explored. During the drilling the magma was sampled directly as glass fragments that quenched in the drilling mud. These were collected at regular intervals over the time that the magma was being drilled into. These time-series samples have now been made available for further study.

For this project, the evolution of magmatic volatiles (H<sub>2</sub>O and CO<sub>2</sub>) concentration over time have been measured and the volatile concentration around bubbles in the volcanic fragments have also been mapped. With addition of major element glass chemistry, the quench temperature, H<sub>2</sub>O-CO<sub>2</sub> saturation pressure and viscosity over time of the Krafla glass cuttings were calculated. The results from this project will help understand how rhyolitic magmas evolve in response to a sudden change in pressure. Volatile contents were calculated using synchrotron radiation Fourier-transform infrared spectroscopy (SR-FTIR) and Focal plane array detectors Fourier-transform infrared spectroscopy (FPA-FTIR), from the glass that quenched in the IDDP-1 borehole. Volatile distributions around bubbles in the quenched magma were mapped by SR-FTIR thereby allowing the H<sub>2</sub>O and CO<sub>2</sub> species distribution around bubbles over a time-series to be quantified at a high spatial resolution and their evolution to be assessed.

## 1.2 Project objectives

The main goals for this project were to understand how the magma responded to the sudden change of pressure and understand why IDDP-1 borehole at Krafla did not erupt. To reach these goals the following objectives were set out:

1. Determine any changes in magmatic volatile concentration through time.
2. Map the distribution of H<sub>2</sub>O species and CO<sub>2</sub> around bubbles in the samples from the time-series set of samples collected during drilling into the Krafla magma body.
3. Determine the magmatic degassing history of the Krafla magma body in response to a sudden loss of pressure as a result of intersection by the drilling.
4. Use the degassing history to make implications on quench timing and viscosity.
5. Investigate the implications of why the Krafla magma did not erupt while it was being drilled.

## 1.3 Thesis format

The thesis is organised as follows:

**Chapter 2** provides a detailed background on the geological setting, the IDDP-1 drilling project and the recovered glass cuttings. Previous studies on these glass cuttings will also be discussed.

**Chapter 3** provides a detailed background on the behaviour of H<sub>2</sub>O and CO<sub>2</sub> in silicate melts that are required to interpret the results, including H<sub>2</sub>O and CO<sub>2</sub> solubilities and diffusivities and the transition between silicate melt and glass.

**Chapter 4** presents the details on of sample preparation and analytical methodologies, including discussion of errors and limitations.

**Chapter 5** presents the results from synchrotron radiation Fourier-transform infrared spectroscopy (SR-FTIR) and electron probe microanalysis (EPMA).



**Chapter 6** discusses the results and interprets the evolution of magmatic volatiles in the intersected magma body over the time the samples were collected, why no eruption occurred during decompression. A framework for future investigation will also be presented in this chapter.

**Chapter 7** concludes a summary of the findings of the thesis.

# Chapter 2: Geothermal drilling and Krafla, Iceland

## 2.1 Introduction

This chapter provides background on geothermal drilling and an insight into the IDDP-1 drilling at Krafla central volcano. It is important to concisely summarise the components and process of drilling to interpret the effect of drilling on the H<sub>2</sub>O and CO<sub>2</sub> content of the magma. Details of the IDDP-1 drilling project and the events leading to the magma drilling are also presented in this chapter. Insight into the chemistry and glass water content of the obsidian chips recovered during drilling will also be presented in this chapter.

## 2.2 Geothermal drilling

Drilling is the most important and expensive part of geothermal exploration, development, and utilisation (Huenges 2010, Glassley 2015). The aim for geothermal drilling is to access the target reservoir in a safe and cost effective way. Drilling involves exerting forces on the rock to be drilled using a multi-toothed drill bit. Rotary drilling is the standard method used in geothermal exploration, in which a string of drill pipe is hung from a derrick and turned by a motor (DiPippo 2008).

### 2.2.1 Drilling equipment and techniques

#### 2.2.1.1 Drillstrings

The drillstring consist of the bottomhole assembly and the drillpipe. The drillstring serves several general purposes, including:

- Feeding drilling fluid to the drill bit.
- Imparting rotary motion to the drill bit.

- Providing and allowing down force to be set on the bit.
- Lowering and raising the bit.

### 2.2.1.2 Bottom hole assembly (BHA)

The bottomhole assembly is the lower portion of the drill-string, consisting of the following:

- The drills bits that are used to excavate the rock/formation.
- Rollercone bits which are used to crush rocks. These consist of three conical rollers, which are equipped with either steel teeth or with special inserts made of tungsten carbide.
- Drill collars (DC) are thick-walled pipes which generate the “weight on bit” (WOB) to assure the drill bit can drill into the formation.
- Jar, a device that produces high kinetic shock energy onto a stuck drill-string section below the jar.
- Shock sub, designed to dampen axial vibrations of the drill-string induced by the drill bit.
- Heavy weight drillpipe (HWDP) is an “intermediate” between the heavy DC and the drillpipe.
- Mud motor, which is a positive displacement drilling motor which uses hydraulic forces of the drilling fluid to drive the drill bit.

### 2.2.1.3 Drilling fluids

The drilling fluid serves many essential drilling functions including:

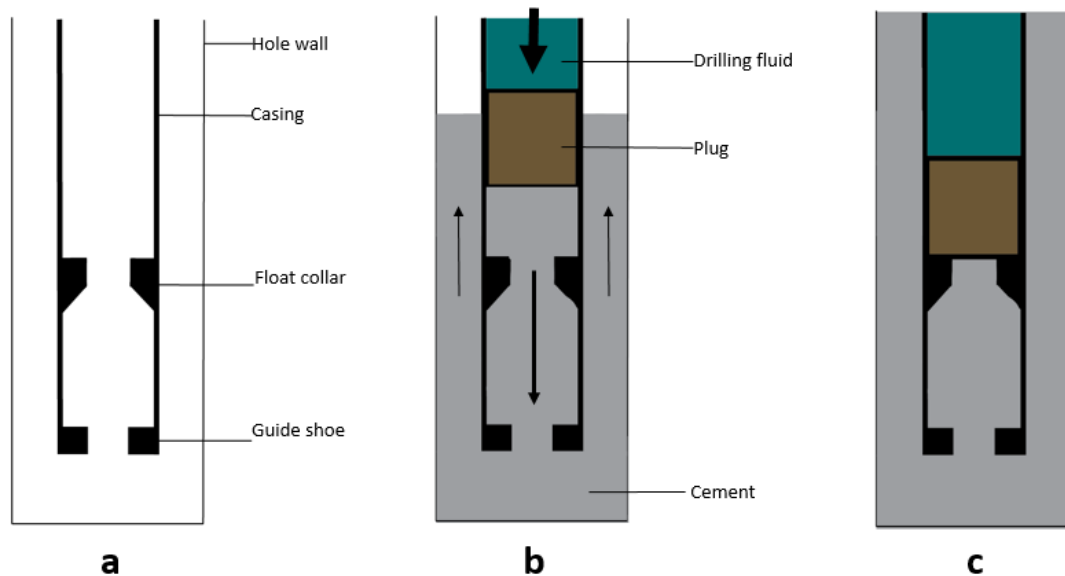
- Cooling the drill bit.
- Transporting cuttings to the surface.
- Preventing cuttings to settle when circulation has stopped.
- Reduce friction between the drill-string and borehole wall.
- Stabilisation of the borehole wall.
- Exertion of hydraulic pressure to prevent gas or fluids from entering the borehole.

- Transport information in the formation of gases, cuttings and fluid, drilled to the surface.

The exact nature of the drill fluid is designed to perform specific tasks. Clay particles are blended into a solution to achieve desired properties (DiPippo 2008). A drilling fluid is pumped down the centre of the drill pipe to the bit and passes through nozzles that accelerate and direct it onto the rock beneath the bit. The drilling fluid captures rock fragments formed by the drill bit and transports them to the surface, through the annulus between the well wall and the drill string (Huenges 2010). When the drilling fluid, which includes rock fragments reaches the surface, it flows over a screen and shaker to separate the rock fragments for examination. The drilling fluid is then returned to the bottom by a mud pump where the cycle is repeated.

#### 2.2.1.4 Casing and cementation

Casing strings are used to secure the walls of drilled hole and isolate fresh water zones to prevent contamination of groundwater during drilling. Fig 2.1 shows a standard method of cement casings. When the hole has been dug to a desired depth the casing is lowered into the hole (Fig 2.1a). A cement mix is then forced down between the inside of the hole and the outside of the casing with a plug, using the drilling fluid (Fig 2.1b). Direct displacement of the cement should lead to the complete filling of the annulus with cement by the time the plug comes to rest on the float collar (Fig 2.1c). Cementing the casing is successful when the excess cement returns to the surface.



**Figure 2.1: Schematic diagram of cementing casing pipes.**

## 2.2.2 Technical risk

### 2.2.2.1 Loss of circulation

The most common drilling problem in geothermal wells is loss of circulation which occurs when a bit encounters a highly permeable zone. The drilling fluid may be absorbed by the formation and therefore not return to the surface (Huenges 2010). “Total loss” is when the loss is great enough to prevent any return. Loss of circulation is desirable when the well is in the production zone as it indicates that the formation has high permeability (Glassley 2015). However, when loss of circulation occurs in shallow zones, it can be an issue as drilling fluids provide well cooling, lubrication, and removal of cuttings. One way to overcome loss of circulation is by pumping down loss circulating material (LCM) pills. LCM pills are made from fibrous, flaky, and granular materials which create a viscous layer onto the porous formation or cracks, enabling drilling fluids to return to the surface (Huenges 2010; Glassley 2015).

#### 2.2.2.2 Blowout prevention

Fluids in geothermal reservoirs can be very hot and under high pressure. To prevent the unintentional and rapid release of such fluids, blowout prevention

(BOP) equipment is installed. BOPs are a set of fast-acting ram-type valves attached to the surface casing that can be used to release the pressure in the hole.

### 2.2.2.3 Stuck pipes or tools

Stuck pipes occur when drill strings cannot be pulled out of the hole, however, circulation may still occur (Devereux 2012). Causes of stuck pipes down well can be associated to; 1) the geometry of the borehole, 2) solid particles which restrict or prevent mud circulation, and 3) mud overbalances over permeable formations which induce differential sticking of the drilling apparatus (Devereux, 2012). Stuck drill strings can be released through by either pumping down high viscosity pills for cleaning the borehole or adding friction-reducing additives to reduce friction coefficient. If wall cave-ins and borehole breakouts are the issue, the weight of the drilling fluid is increased to clean out the borehole (Huenges 2010).

### 2.2.2.4 Fishing

Occasionally a part of the drill-string can break off and be left in the borehole. This lost part is called a “fish” (Huenges 2010, Devereux 2012). The process of retrieving the fish is termed “fishing”.

The main causes for fish occurring in the hole includes;

- Failure somewhere in the drill string caused it to break and drop into the hole.
- A stuck pipe cannot be freed and has to be either cut or unscrewed downhole.
- An object falls into the borehole.

### 2.2.2.5 Side-tracking

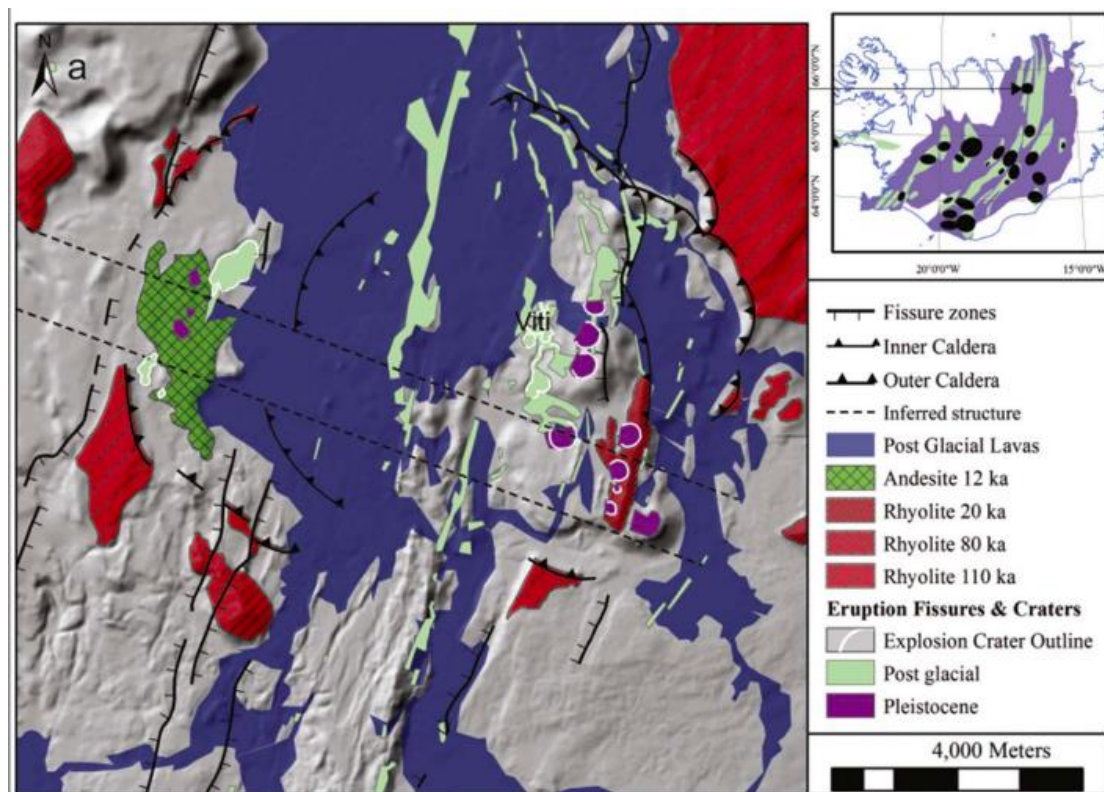
If fishing is unsuccessful, occasionally wells will be side-tracked. Side-tracking is the process of leaving an existing hole at a certain depth and branching off at an angle to form a new hole (Huenges 2010). This can be achieved by forming a cement plug of a certain length at the desired depth in the borehole. After the cement has hardened, a directional drilling assembly is used to side-track the drill string from the cement plug into the borehole wall. When the new hole is drilled, the new well path can be drilled.

## 2.3 Geological setting, Krafla Iceland

The Krafla central volcano is located on the Northern Iceland Rift Zone (Fig. 2.2). The Krafla geothermal field sits within an 8 km collapsed caldera (Sæmundsson, 1991, Zierenberg et al. 2013). The caldera is bisected by the NNE-SSE fissure swarm that marks the North Iceland Rift Zone. The caldera is largely filled by basaltic lavas and hyaloclastites. Rhyolite has erupted periodically, but intermediate composition lavas are not abundant (Sæmundsson, 1991; Jónasson, 1994).

Extensive drilling for geothermal development in the caldera began in 1974 (Zierenberg et al. 2013). The results of which revealed the extent of several active geothermal activities and subsurface geology (Sæmundsson, 1991; Jónasson, 1994). The drilling showed that near surface postglacial basaltic lavas are underlain by hyaloclastite erupted during the last glacial stage (Guomundsson, 1983). An older sequence of hyaloclastite is overlain and underlain by interglacial lavas down to about 1.2 km, where intrusive rocks begin to dominate. In the deeper regions of the geothermal reservoir doleritic intrusions are abundant, however felsic intrusions were also intersected in several drill holes (Zierenberg et al. 2013). Intrusions generally become more abundant and coarser grained at deeper depths.

The most recent eruptions in the Krafla central volcano were the Krafla Fires; a period of rifting that started with basaltic fissure eruptions in 1975 and continued irregularly until 1984 (Holmgeirsson et al. 2010; Zierenberg et al. 2013). This eruptive sequence consisted of nine volcanic eruptions and ~15 swelling and subsidence events. This volcanic activity affected the geothermal reservoir substantially and limited steam production in existing power plants (Zierenberg et al. 2013).



**Figure 2.2: Geological map of the Krafla central volcano, highlighting key eruptive material, volcanic and structural features.** Figure from Kennedy et al. (2018)

## 2.4 Boreholes hit magma

### 2.4.1 Previous magma intersection

Prior to the IDDP-1 drill hole, drilling into magma has occurred on two previous occasions in Hawaii (2005) and in Krafla (2008) (Teplow et al. 2009, Mortensen et al. 2010, Zierenberg et al. 2013).



In 2005, dacite melt was encountered during a routine commercial operations of injection well KJ-13 at Puna Geothermal Venture wellfield, Big Island of Hawaii (Teplow et al. 2009). This was located along a segment of the Kilauea Lower East Rift Zone. At a depth of 2488 m, a dacite melt was intercepted. The melt flowed up the borehole and was repeatedly redrilled over a depth interval of 8 m. Clear vitric glass cutting were recovered to the surface. These cuttings were dacitic and has a silica content of 67 wt%. The intercepted magma is interpreted as an end-stage differentiate of basaltic magma injection into the Kilauea East Rift zone.

Magma was probably first intersected in the Krafla Geothermal system in 2008. The KJ-39 well in the Suourhlioar well field, intercepted magma at 2,571 m depth. Glass retrieved from this event was partially hydrated with silica content ranging from 69.0 to 78.8 wt% (Mortensen et al. 2010, Zierenberg et al. 2013).

## 2.4.2 IDDP-1 well

The drilling contract for the IDDP-1 well was signed in August 2008 (Holmgeirsson et al. 2010). The IDDP-1 exploration well was designed to drill to completion at 4.5 km depth, where injected water was expected to reach supercritical conditions (Zierenberg et al. 2013). Based on seismic and magnetotelluric surveys, a magma body was inferred to underlie Krafla Caldera at depths of 3-7 km (Einarsson 1978) and the well was designed to intercept the supercritical hydrothermal fluids circulating above it.

The IDDP-1 well and drilling design plan was as follows:

- Surface casing to be drilled to 100 m below the surface. A wellbore rotary drill with a 26" tricone bit with a 36" under-reamer to be used with a drilling fluid of a bentonite mud mixture.
- Intermediate casing I to be emplaced following drilling to 300 m below the surface. Drilling using a wellbore rotary drill with a 23" tricone bit with a bentonite mud mixture as drilling fluid.

- Intermediate casing II to be emplaced upon reaching 800 m below the surface. Drilling with a wellbore rotary with a 23" tricone bit and a drilling fluid of a bentonite mud mixture.
- Anchor casing to be installed on reaching 2400 m below the surface using a wellbore drill with 16-1/2" tricone bits. Drilling fluid to be a bentonite mud mixture as long as it could be sufficiently cooled, if not the drilling fluid would be switched to water.
- Production casing to be drilled to 3500 m below the surface using a wellbore drill with 12-1/4" tricone bits. Drilling fluid to be a bentonite mud mixture as long as it could be sufficiently cooled, if not the drilling fluid would be switched to water.
- Slotted liner to be installed up to 4500 m below the surface, drilled by a wellbore with a 8-1/2" tricone bits. Water to be used as the drilling fluid.

#### 2.4.2.1 Surface casing

The pre-drilling of IDDP-1 began in June 2008. The well was drilled with a 26" tricone pilot and a 36" under reamer (Pálsson et al. 2014). Bentonite-based mud was used as the drilling fluid and the well was drilled to a depth of 87 m below the surface in nine days. At this depth the formation was stable for the casing shoe and it was decided to run the casing (Holmgeirsson et al. 2010, Pálsson et al. 2013).

#### 2.4.2.2 Intermediate casing I (18/Nov/2008 to 24/Nov/2008)

Drilling for the intermediate casing I started on 18 November 2008. Due to the hard formations encountered, drilling progress was slower than expected with a rate of penetration (ROP) of 2.5 m/h. No circulation losses were detected while drilling and the depth of this section was 275 m below the surface.

The well was cased to 260 m below the surface depth due to problems in the wellbore (Holmgeirsson et al. 2010, Pálsson et al. 2014). On 24 November cementing was successful.

#### 2.4.2.3 Intermediate casing II (27/Nov/2008 – 10/Dec/2008)

Drilling for the intermediate II casing continued through 27 November (Holmgeirsson et al. 2010, Pálsson et al. 2014). To keep the hole relatively vertical, drilling progress was slow and the weight on bit (WOB) was kept low. On 8 December, minor drilling fluid losses were detected and healed with LCM at a depth of 788 m below the surface. The well was cased to a depth of 784 m below the surface and the casing was successfully cemented on 10 December (Holmgeirsson et al. 2010, Pálsson et al. 2014).

#### 2.4.2.4 Anchor casing (24/March/2009 – 18/May/2009)

To drill further (anchor and production casing) it was proposed to use a new drilling rig. A Drillemec HH-300 rig was chosen and mobilised to the well site in the middle of March 2009 and drilling recommenced on 24 March (Holmgeirsson et al. 2010, Pálsson et al. 2014).

Anchor casing of the bottom hole assembly (BHA) consisted of a 16-1/2" roller cone tricone bit, a 9-1/2" mud motor, two 16-1/2" stabilizers, an Anderdrift tool, 9-1/2" and 8" drill collars, a shock sub, a jar, and heavy weight drill pipes (Holmgeirsson et al. 2010, Pálsson et al. 2014). Initially the drilling fluid used for the anchor casing was water to drill out the cement, however, when drilling into the formation at 803 m below the surface, the drilling fluid was changed back to mud. Drilling progress commenced well up to a depth of 2000 m below the surface although several unusual events occurred. These events included:

1. At 1194 m depth, a sudden pressure drop in the standpipe was observed and torque dropped. A float sub twisted apart in the BHA and the drill-string broke. Fishing took 48 hours to recover the broken piece. During this time, the mud motor was blocked and had to be replaced.
2. When drilling resumed the ROP was rather low (3-5 m/h). ROP decreased and testing showed either the bit or the mud motor was failing. A decision

was made to pull out of hole (POOH) from 1400 m below the surface. The mud motor was not working properly and was again replaced.

3. On reaching 1432 m below the surface, losses of 20 l/s were detected and a decision was made to POOH and cement off the loss zone to minimise mud losses, and to prevent interflow between loss zones during casing. After only drilling for a further 47 m, the bit was in a bad condition with almost all carbides broken off on the outer rows of the cones. On 7 April a junk basket was used to retrieve the missing carbides, however only formation cuttings and small metal fragments were retrieved. At the time the well appeared to be tight and the cement plugging was abandoned.
4. On 8 April the BHA was returned in hole (RIH) without a mud motor. No circulation losses were detected. An Anderdrift tool was used to measure the wells inclination which was 1.5°. On 13 April when drilling had reached 1907 m below the surface, the drill was POOH to replace the bit. The bit was badly worn and needed replacing. Drilling recommenced on April 15.
5. On 18 April at 2074 m below the surface a sudden drop on the standpipe was observed and the torque dropped. The drill collar had twisted apart and had dropped the lower string assembly. Fishing occurred for two hours to recover the BHA. A new BHA was run in the hole on 20 April and drilling resumed again from 2074 m below the surface.

On the morning of 21 April when drilling had reached 2101 m below the surface the torque values were fluctuating and the drill was pulled out for reaming (Holmgeirsson et al. 2010). When the last single was pulled out the weight of the bit dropped by 20 tons and the standpipe pressure decreased. The BHA had broken and fishing attempts were carried out over six days to recover the broken assembly. These attempts were unsuccessful. (Holmgeirsson et al. 2010, Pálsson et al. 2014). During this time the cause of the drill string loss was unknown, it was later interpreted as the drill string becoming stuck in response to interception with a magma body (Pálsson et al. 2014; Zierenberg et al. 2013; Saubin et al. 2017).

To continue drilling, a side-track to pass the fish was necessary. A 100 m thick cement plug was placed above the fish. On 12 May the drill bit was in formation at 1934 m below the surface, however the conditions in the well became difficult (Holmgeirsson et al. 2010, Pálsson et al. 2014). Sufficient hole cleaning was hindered by washout and circulation losses. To make hole cleaning more effective and to reduce the risk of a cave-in from the surface to 2000 m below the surface, a modification to the casing program was decided. Instead of having the anchor casing to 2400 m below the surface, it was changed to 2000 m below the surface. Drilling for the anchor casing was completed at 2005 m depth (Holmgeirsson et al. 2010, Pálsson et al. 2014).

The casing job started on the 18 May. The anchor casing consists of two sections of different thicknesses. The top 300 m were 13 5/8" lb/ft T9, the remaining casing was 13 3/8" 72 lb/ft K-55 (Holmgeirsson et al. 2010, Pálsson et al. 2014). The casing shoe was set at 1949 m below the surface and 24 m above the shoe was a stab-in float collar (Pálsson et al. 2014).

#### 2.4.2.5 Production casing (25-May-2009 – 25-June-2009)

Drilling for the production casing started on the 25 May. Once the cement had been drilled out, the challenge was to circulate the cuttings from the bottom with 12-1/4" drill bit in a 16-1/2" hole (Holmgeirsson et al. 2010, Pálsson et al. 2014). To overcome this, the plan was to drill with high viscous pills with a relatively low ROP and low pumping rate, after which the pumping rate would be increased and the string rotated and moved up and down hole (Holmgeirsson et al. 2010, Pálsson et al. 2014).

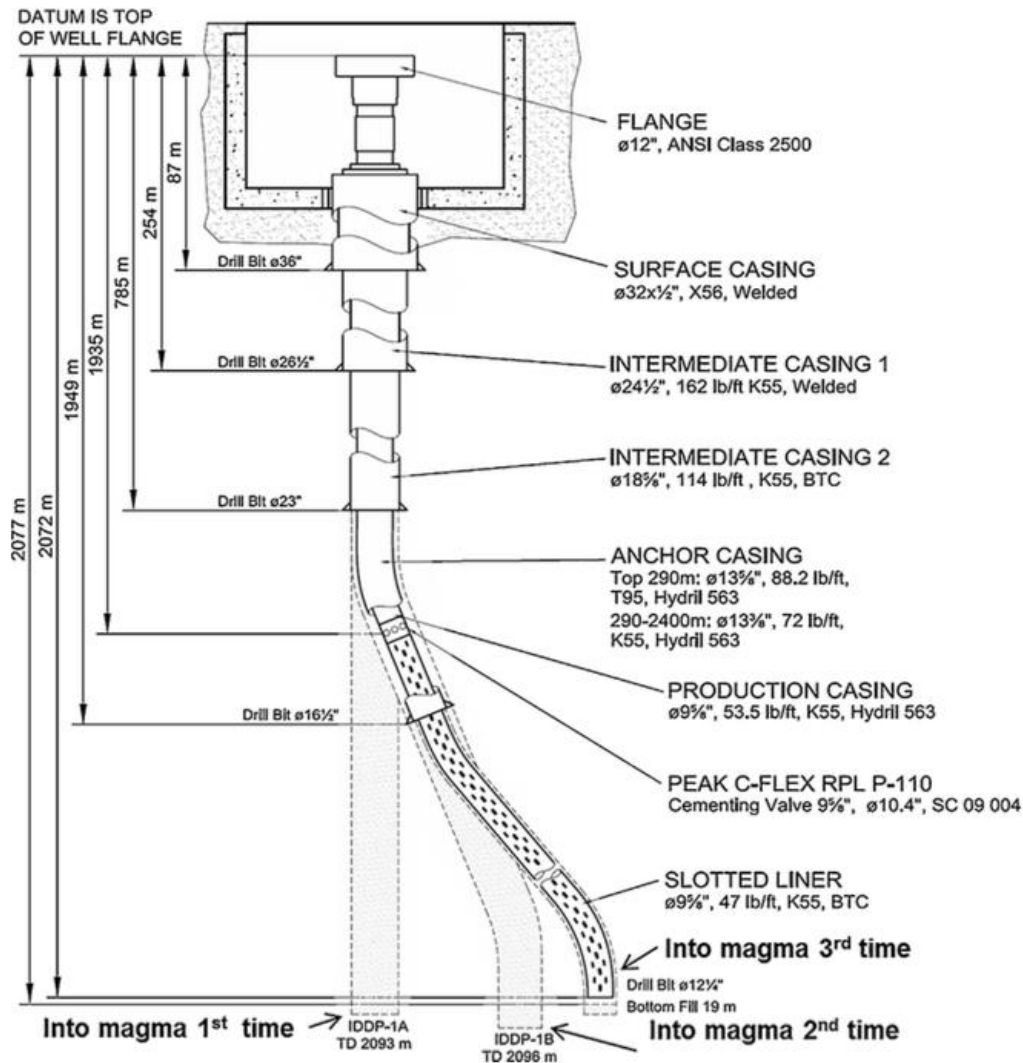
On 8 June at 2103 m below the surface, there was a sudden rise in torque and the string became stuck for two minutes (Pálsson et al. 2014). Using 160 tons the string was freed and one single was pulled out. Several high viscous pills were used to circulate the drill for one and a half hours. After circulation, a string was lowered to the bottom of the hole. When the bit reached the well bottom, the torque once again increased and the bit was pulled up 13 m. A few minutes later

the torque increased and the drill string got stuck (Holmgeirsson et al. 2010, Pálsson et al. 2013). Immediately after the standpipe pressure increased and there was a loss in circulation (Holmgeirsson et al. 2010, Pálsson et al. 2014). On 10 June the string was on the surface and fishing attempts with jarring and pulling was commenced. Fishing was carried out for two days without success. A possible reason identified for the string being stuck was that the cement had caved in. It was decided to attempt to dissolve the cement by pumping down two pills of hydrochloric acid. However, the drill string was not freed and the drill string was detached again and a second side-track was planned (Holmgeirsson et al. 2010, Pálsson et al. 2014). The cause behind the stuck bit at the time was unknown, this event was later interpreted as a second interception with the magma body (Pálsson et al. 2014; Zierenberg et al. 2013; Saubin et al. 2017).

Two attempts were needed to cement above the fish. On 15 June, a cement plug was placed from the top of the fish at 2072 m below the surface, up to 1927 m below the surface (Holmgeirsson et al. 2010, Pálsson et al. 2014). On the same day a side-track was attempted, however, the cement was too soft for side-tracking and the bit was pulled out to allow the cement to harden. On 19 June side-tracking commenced at a depth of 1985 m below the surface (Holmgeirsson et al. 2010, Pálsson et al. 2014). During side-tracking the bit was in the country rock but no cement was observed. Torque and pressure fluctuated and big chips of fine-grained basalt were recovered. These unexpected returns were explained as resulting from thermal cracking of the formation, which is believed to be above 340 °C and was cooled repeatedly as a result of the length of time that the hole was open below the casing shoe (Holmgeirsson et al. 2010, Pálsson et al. 2014). On reaching 1992 m below the surface it was decided to POOH and lay down the mud motor to reduce the chances of the drill string becoming stuck again. On 23 June, drilling resumed at a low ROP. At 2071 m below the surface total loss of circulation occurred. High viscous pills were pumped down two times to keep the well clean and maintain this state. Drilling was running fine until reaching 2100 m below the surface (Holmgeirsson et al. 2010, Pálsson et al. 2014).

On 25 June at 2104.4 m depth the rate of penetration doubled from 2 m/hr to 4 m/hr, the torque increased and the string was stuck and had to be pulled free (Holmgeirsson et al. 2010, Pálsson et al. 2014). A single was pulled out and the well was circulated for 90 minutes. The bit was run down to the bottom again and the torque increased again at the same depth. When running in again (2 m) the top drive and the single on the floor moved upwards and the weight decreased by 45 tons and immediately the string was stuck. Circulation was managed to be maintained and the returns were pulsating to the surface initially but became steady (Holmgeirsson et al. 2010; Zierenberg et al. 2013). The returns included quenched glasses, which initially were clear containing relatively spherical bubbles, however, they became red-brown in colour (further detail on the recovered glass cuttings in section 2.4). Circulation was maintained for 24 hours without moving the string, but after a while there were no returns. After the string was pulled out the POOH commenced and the bit was in excellent condition. It was clear that at 2104.4 m the drill had drilled into magma. Further drilling was pointless and the drill hole was capped (Holmgeirsson et al. 2010; Zierenberg et al. 2013).

Fig 2.3 summarises the IDDP-1 drill, the depths reached for each casing and the estimates of the magma intercepts. During the drilling of the IDDP-1 magma was intercepted three times between 2101 – 2104.4 m depth. Magma interception occurred when there was an increase in standpipe pressure followed by a decrease in weight on the bit. This was interpreted as representing the ascent of magma within the borehole (Saubin et al. 2017).



**Figure 2.3: Overview of the IDDP-1 drilling configuration illustrating key changes in casing used at depth over the campaign.** Figure from Pálsson et al. (2014).

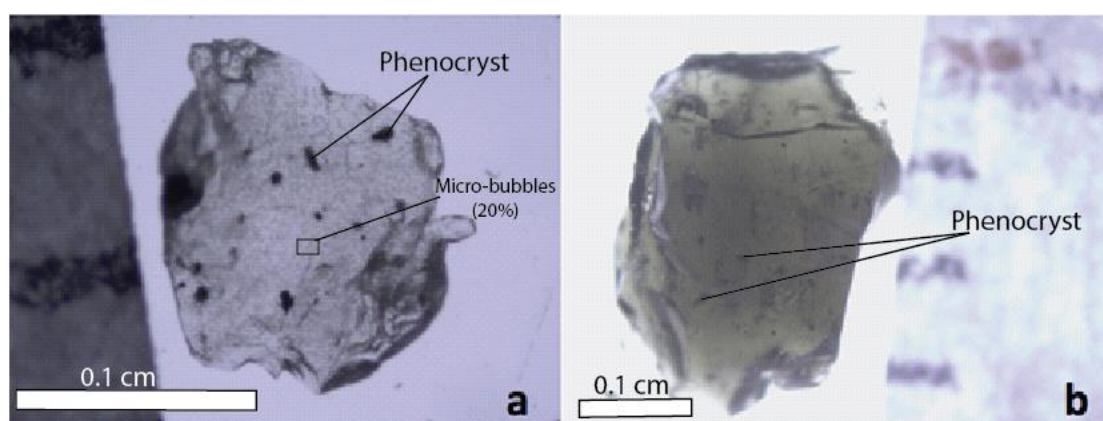
## 2.5 The recovered glass cuttings

### 2.5.1 Time constrained glass cuttings

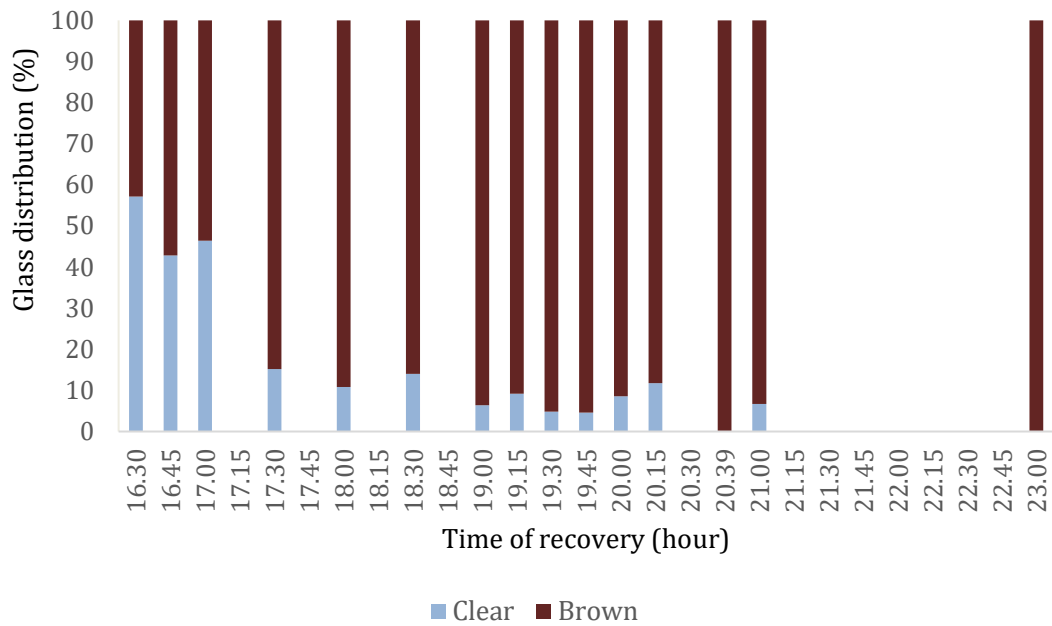
The set of samples received for this project were retrieved in a time-series from 15:17 on 24 June 2009 to 00:50 on 26 June 2009. The colour and vesiculation of these glass cuttings have been determined by observations made through a binocular microscope (Saubin et al. 2017).



Initially, clear sparsely phyric pumiceous glass fragments (Fig 2.4a) were most abundant in the returned cuttings (Holmgeirsson et al. 2010, Pálsson et al. 2014, Zierenberg et al. 2013). The proportion of the clear glass cuttings decreases significantly after 17:00 (Fig 2.5, Saubin et al. 2017) and the dominant cuttings recovered from the IDDP-1 become sparsely phyric, dark brown, poorly vesiculated glass (Fig 2.4b). Bubbles in the glass cuttings were initially spherical in shape, suggesting that the melt was not being significantly sheared at the time it was quenched. In contrast the bubbles in the later recovered cuttings were highly elongated, suggesting shearing due to movement of the magma. Both clear and brown glass have a refractive index of 1.49 (Zierenberg et al. 2013). Zierenberg et al. (2013) suggest that the textural change of the glass indicates that the pressure or quench time of the magma varied over time during penetration by the drill bit.

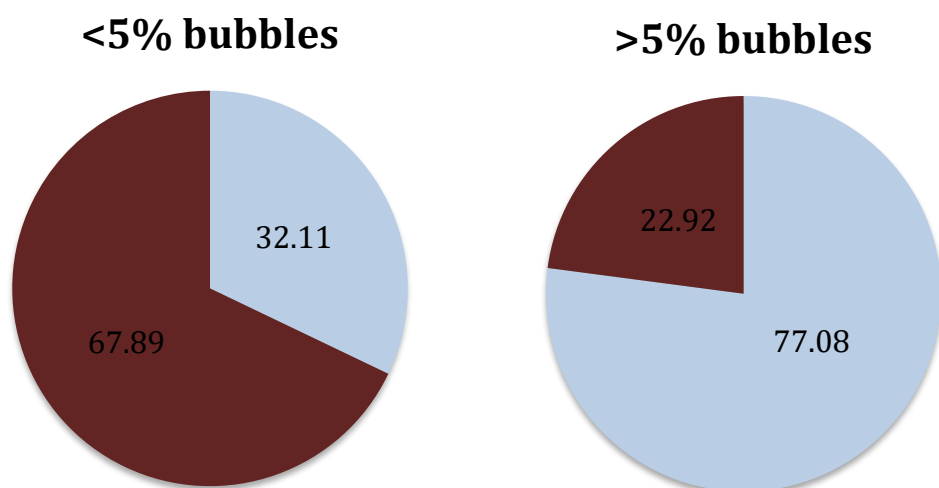


**Figure 2.4: Example of glass cuttings from IDDP-1 showing colour variation.** a) Clear glass with micro-vesicles and phenocrysts; b) Brown glass with no micro-vesicles.



**Figure 2.5: Variation of colour of the recovered glass cuttings with time.**  
Figure obtained from Saubin et al. 2017)

The glass cuttings have been separated into low vesicularity (<5% bubbles) and high vesicularity (>5%). The majority of the cuttings that have a high vesicularity are clear (77.08%), whereas the majority of glass cuttings that have a low vesicularity are brown (Figure 2.6).



**Figure 2.6: Variation of glass colour with vesicularity.** Most high vesicularity glasses are clear, and most low vesicularity glasses are brown. Figure from Saubin et al. (2017)

## 2.6 Previous work on the IDDP-1 glass cuttings

The unexpected interception of the magma provides many opportunities for research as it gives a direct insight into a resting magma body. This not only presents opportunities for understanding the evolution of rhyolitic magma chambers, but also raises possibilities for future geothermal development. All glass cuttings used in previous studies were received at 1700 (Saubin pers.comm., 2017).

### 2.6.1 Glass and mineral chemistry

Previous work on the glass cuttings by Zierenberg et al. (2013) has included characterising their composition and determining the origin of the magma. Based on petrography, Zierenberg et al. (2013) identified three chemically distinct high-silica rhyolite glasses quenched from melt by the drilling fluid.

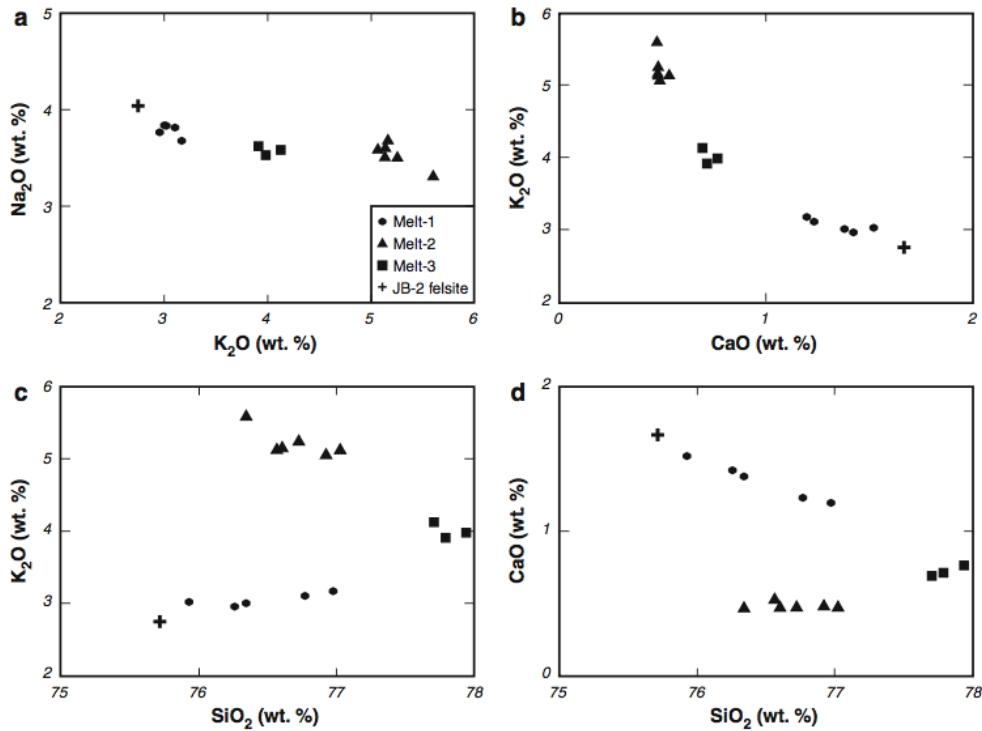
- Melt-1 is the dominant type of glass recovered from the IDDP-1 and is interpreted to be the quenched intersected magma. Zierenberg et al., (2013) defines Melt 1 to include both clear and brown glass cuttings described in this thesis. The clear glass is termed Melt-1a in this thesis, and the brown glass is termed Melt-1b.
- Melt-2 is a crystalline felsic intrusive rock with quenched rhyolite glass. It is believed to be representative of the host rock into which the rhyolite melt was emplaced. The felsite is composed of alkali feldspar, plagioclase, quartz, titanomagnetite and augite; the rhyolite glass makes up variable amounts of the felsite (up to ~20%).
- Melt-3 is a glass that is textually, and mineralogically similar to the felsite (except that it contains less alkali feldspar or contain small irregular grains of residual alkali feldspar). It contains abundant crystals and/or glomerocrysts of quartz, plagioclase, augite and magnetite. There are signs of resorption with some plagioclase crystals exhibiting irregular compositional zoning and quartz crystals highly embayed. The glass in these crystal rich cuttings is typically moderately vesicular. The textures of Melt-3 suggest that it has invaded and partially assimilated the felsite.

Zierenberg et al. (2013) used an electron probe microanalysis (EPMA) to determine the glass composition of the glass cuttings and X-ray fluorescence to determine the whole rock chemistry. Table 2.1 summarises their glass and whole rock compositions.

**Table 2.1: Summary of the major elements composition of glass (wt%) and the whole rock chemistry determined by X-ray fluorescence.** Major oxide and the whole rock chemistry normalised to 100 wt% on an anhydrous basis. Data obtained from Zierenberg et al. (2013).

	<b>Melt-1</b>	<b>STDEV</b>	<b>Melt-2</b>	<b>STDEV</b>	<b>Melt-3</b>	<b>STDEV</b>	<b>JB-2</b>
	<b>(n=5)</b>		<b>(n=6)</b>		<b>(n=3)</b>		<b>(felsite)</b>
<b>SiO<sub>2</sub></b>	76.47	0.4	77.04	0.43	77.81	0.09	75.66
<b>TiO<sub>2</sub></b>	0.33	0.04	0.21	0.05	0.2	0.02	0.42
<b>Al<sub>2</sub>O<sub>3</sub></b>	12.02	0.1	11.68	0.08	11.69	0.05	11.98
<b>FeO</b>	2.75	0.16	1.84	0.14	1.85	0.06	2.99
<b>MnO</b>	0.07	0.02	0.09	0.04	0.02	0.02	0.07
<b>MgO</b>	0.19	0.04	0.07	0.01	0.10	0.01	0.35
<b>CaO</b>	1.35	0.13	0.48	0.02	0.72	0.04	1.67
<b>Na<sub>2</sub>O</b>	3.79	0.07	3.55	0.12	3.58	0.05	4.04
<b>K<sub>2</sub>O</b>	3.05	0.08	5.05	0.21	4.01	0.11	2.75
<b>P<sub>2</sub>O<sub>5</sub></b>	n.a	-	n.a	-	n.a	-	0.07

The major oxide analysis has shown that all the melts are high-silica rhyolites. The difference between Melt-1 and Melt-2 is the higher K<sub>2</sub>O and lower CaO content in Melt-2. The composition of Melt-1 is similar to the whole rock composition of crystalline felsite (JB-2). The composition of Melt-3 glass lies between Melt-1 and Melt-2 on a plot of Na<sub>2</sub>O versus K<sub>2</sub>O (Fig 2.7a) and K<sub>2</sub>O versus CaO (Fig 2.7b), suggesting it is a mixture of Melt-1 and Melt-2. However, Melt-3 shows higher SiO<sub>2</sub> content than the other two melts (Fig 2.7c-d), which is consistent with the textural evidence of quartz resorption.



**Figure 2.7: Chemical variation diagrams for the different glasses and the crystalline felsite (JB-2).** Data are normalised on an anhydrous basis. a) K<sub>2</sub>O versus Na<sub>2</sub>O, b) CaO versus K<sub>2</sub>O, c) SiO<sub>2</sub> versus K<sub>2</sub>O and d) SiO<sub>2</sub> versus CaO. Figure from Zierenberg et al. (2013).

## 2.6.2 H<sub>2</sub>O and CO<sub>2</sub>

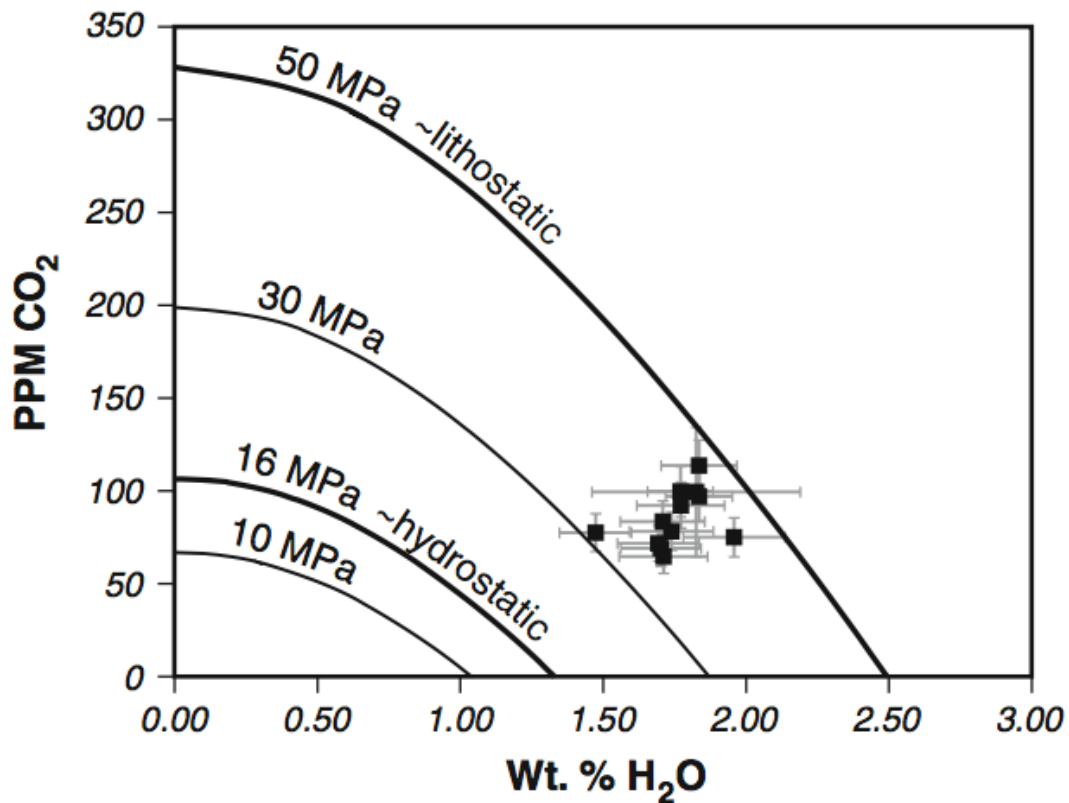
Table 2.2 summarises the volatile measurements on the quenched Melt-1 glasses done by Zierenberg et al. (2013). Vesiculated glass have volatile contents identical to the more common non-vesiculated glass. The high ratio of OH to H<sub>2</sub>O<sub>m</sub> in the glass suggests that the melt was quenched to glass at a high temperature by the drilling fluid (Zierenberg et al. 2013).

**Table 2.2: H<sub>2</sub>O and CO<sub>2</sub> content of Melt-1 from Zierenberg et al. (2013). Volatile content is determined by FTIR.**

<b>Sample</b>	<b>H<sub>2</sub>O<sub>m</sub> (wt%)</b>		<b>OH (wt%)</b>		<b>H<sub>2</sub>O<sub>t</sub> (wt%)</b>		<b>CO<sub>2</sub> (ppm)</b>	
<b>Vesiculated (n=2)</b>	0.58	0.05	1.12	0.06	1.7	0.01	78	8
<b>Non Vesiculated (n=11)</b>	0.6	0.11	1.18	0.09	1.78	0.15	87	16
<b>Average</b>	0.6	0.02	1.17	0.04	1.77	0.17	85	15

Zierenberg et al. (2013) calculated the H<sub>2</sub>O-CO<sub>2</sub> saturation pressures of the glasses using VolatileCalc (Newman and Lowenstern 2002) assuming a temperature of 900 °C (section 2.4.3). The H<sub>2</sub>O-CO<sub>2</sub> saturation pressures are mostly between 35 and 45 MPa (Zierenberg et al. 2013). These pressures are below the lithostatic pressure but greater than the hydrothermal system (Fig. 2.8). Pressures above the hydrostatic pressure in an active geothermal field suggest that there must be a low permeability barrier that separates the melt from the overlying convection hydrothermal system (Zierenberg et al. 2013). Loss of circulation prior to the intercepting the magma suggests that a high permeability zone existed above the magma.

Trewick et al. (2016) investigated the vesiculation history of the magma and attempted to characterise the average H<sub>2</sub>O concentration and the bubble population density in quenched glass with different textures. The results show that H<sub>2</sub>O concentrations have some spatial variability, and they observed enrichment of H<sub>2</sub>O towards the bubbles, which they suggested was the result of resorption. However, mean H<sub>2</sub>O values for the clasts they analysed ranged from 1.41 – 1.68 wt %, with no statistically significant difference between glass cuttings that exhibit different textures.



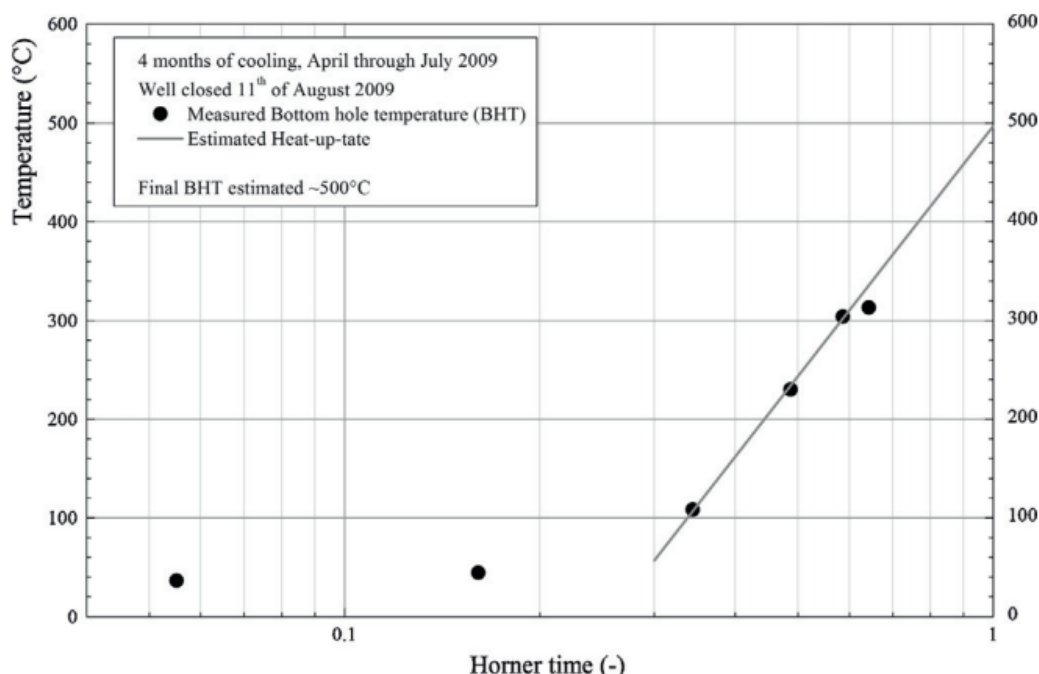
**Figure 2.8: H<sub>2</sub>O versus CO<sub>2</sub> in Melt-1 glass.** Saturation pressure contours were calculated using VolatileCalc (Newman and Lowenstern 2002). Figure obtained by Zierenberg et al. (2013).

### 2.6.3 Temperature estimates

Previous work on the IDDP-1 glass cuttings includes estimating the temperature of the intercepted melt. Based on two-pyroxene geothermometry and modelling of the crystallisation sequences, Zierenberg et al. (2013) estimate the in situ temperature of the magma to be between 850 – 920 °C. The high ratio of dissolved OH to H<sub>2</sub>O<sub>t</sub> (0.66) indicates that the magma was quenched at temperatures between 760 and 940 °C (Elders et al. 2014).

Using a Horner plot the temperature of the host rock (Felsite) the magma was sitting was indicated. Horner plots are an experimental analysis which can be used in drilling operations to derive reservoir pressure and/or temperature conditions. This is done through the derivation of a heating rate at hole bottom following temporary well closure (Dowdle and Cobb, 1975). Based on temperature logs from 13 October 2009, 3 December 2009 and 28 January 2010,

the host rock temperature was indicated to be  $\sim 500^{\circ}\text{C}$  (Fig 2.9), (Morstensen et al. 2014).



**Figure 2.9: Horner plot for six recovery logs from the ICCP-1 borehole.** Assuming no active convection processes near the bottom, temperature recovery indicates  $\sim 500^{\circ}\text{C}$  near the bottom of the well. From Mortensen et al. (2014).

## 2.7 Project objective

This project will differ from previous work on the Krafla glass cuttings as it will investigate the magmatic volatile ( $\text{H}_2\text{O}$  and  $\text{CO}_2$ ) content, the glass composition and the glass texture in glass cuttings that were collected over the time period that circulation was maintained after the third interception of the magma body and returns of quenched glass reached the surface. The time of return of each glass cutting has been recorded allowing these samples to be used to understand how the magma responded to being intercepted by the drilling. Samples that Zierenberg et al. (2013) interpreted as the magma body (i.e. Melt-1) were selected. A total of 18 glass cuttings with different colour and vesicularity were selected (Table 2.3). For this project, the glass cuttings have been grouped into two categories based on textural characteristics; clear glass cuttings (termed Melt-1a) and brown glass cuttings (termed Melt-1b).



**Table 2.3: Summary of the sample analysed by FTIR. Sample name based on time of the glass cutting recovery.**

<b>Sample</b>	<b>Colour</b>	<b>vesicularity</b>
1545	Melt-1a	<5%
1645B	Melt-1b	>5%
1645C	Melt-1b	<5%
1645D	Melt-1a	>5%
1730B	Melt-1b	<5%
1730C	Melt-1b	>5%
1800A	Melt-1b	>5%
1800B	Melt-1b	<5%
1800C	Melt-1b	>5%
1800E	Melt-1b	<5%
1830C	Melt-1a	<5%
1915	Melt-1b	<5%
1945A	Melt-1b	<5%
1945B	Melt-1b	<5%
2015	Melt-1b	<5%
2039C	Melt-1b	<5%
2039D	Melt-1b	<5%
0050	Melt-1b	>5%

## 2.7.1 Additional glass cuttings

In addition six glass cuttings collected together at 1700 on 25 June have been included. These glass cuttings consist of brown (Melt-1b) clasts ranging in vesicularity (Table 2.4). H<sub>2</sub>O contents for these samples had already been collected using Fourier transform infrared spectroscopy using a conventional bench-top global source and focal-plane-array camera detector (FPA-FTIR) and these data (Joyce, pers. Comm. 2017) have been processed to allow comparison to the time-series of samples.

**Table 2.4: Summary of the textural description of the additional Krafla glass cuttings.**

<b>Sample</b>	<b>Colour</b>	<b>Vesiculation</b>
Ice 216-3	Melt-1B	<5%
Ice 220-8	Melt-1B	>5%
Ice 214-7	Melt-1B	>5%
Ice 217-2	Melt-1B	<5%
Ice 219-5	Melt-1B	<5%
Ice 221-1	Melt-1B	>5%

## 2.8 Summary

Preparation of the IDDP-1 well drilling started in the year 2000 and continued until drilling activity started in mid-2008. The IDDP-1 well was designed to drill down to ~4-5 km in order to reach fluids. However drilling problems occurred for nearly three months from March 2009. Drilling problems included getting stuck, having twist offs, which results in side-tracking at two different times (21 April and 8 June). The causes of these issues become clearer at 2104 m depth. On 24 June the rate of penetration increased, the torque increased and the string got stuck and the drill moved upwards. It was clear the drill bit drilled into a magma body of some size. Drilling stopped when the magma was intercepted, however circulation continued and glass cuttings were recovered and time of recovery recorded. Initially clear glass cuttings were the most abundant and then decreased and dark brown glass cuttings became dominant. Previous studies have indicated that the magma body consists of high silica (76.5% SiO<sub>2</sub>) rhyolite melts with an in-situ temperature estimated between 850 and 920 °C emplaced in a country rock at ~500 °C. Previous volatile studies show that the quenched melt contains 1.77 wt% H<sub>2</sub>O and 85 ppm CO<sub>2</sub>, indicating that the in situ pressure is above hydrostatic (~16 MPa) and below lithostatic (~55 MPa). In summary, these previous studies have not looked at how the characteristics of the volatiles dissolved in the melt changed with time. This project will differ from other work by examining the time-series of cuttings. The clear glass (Melt-1a) and the brown

glass (Melt-1b) will also be examined to determine whether they came from different sources.

## Chapter 3: H<sub>2</sub>O, CO<sub>2</sub> and silicate melts

### 3.1 Introduction

Volatiles are a very important factor in magmatic processes (Sparks, 1978). Dissolved volatiles change the physical and chemical properties of melts and are a driving force of volcanic eruptions. Silicate melts like the Krafla glass cuttings contain dissolved volatile species that are weakly soluble, and exsolve to form bubbles. The most common magmatic volatiles are H<sub>2</sub>O and CO<sub>2</sub>. This chapter presents an overview of the structure and composition of silicate melts, the mechanisms and controls of magmatic volatile solubility and diffusion in the melt, and how these factors influence bubble growth and resorption. Understanding the role magmatic volatiles have on melts and their properties can help determine the causes of the volatile distributions observed in the Krafla glass cuttings (Chapter 5) and give an insight to the magmatic evolution which occurred owing to the interception of the magma by IDDP-1 drilling (Chapter 6).

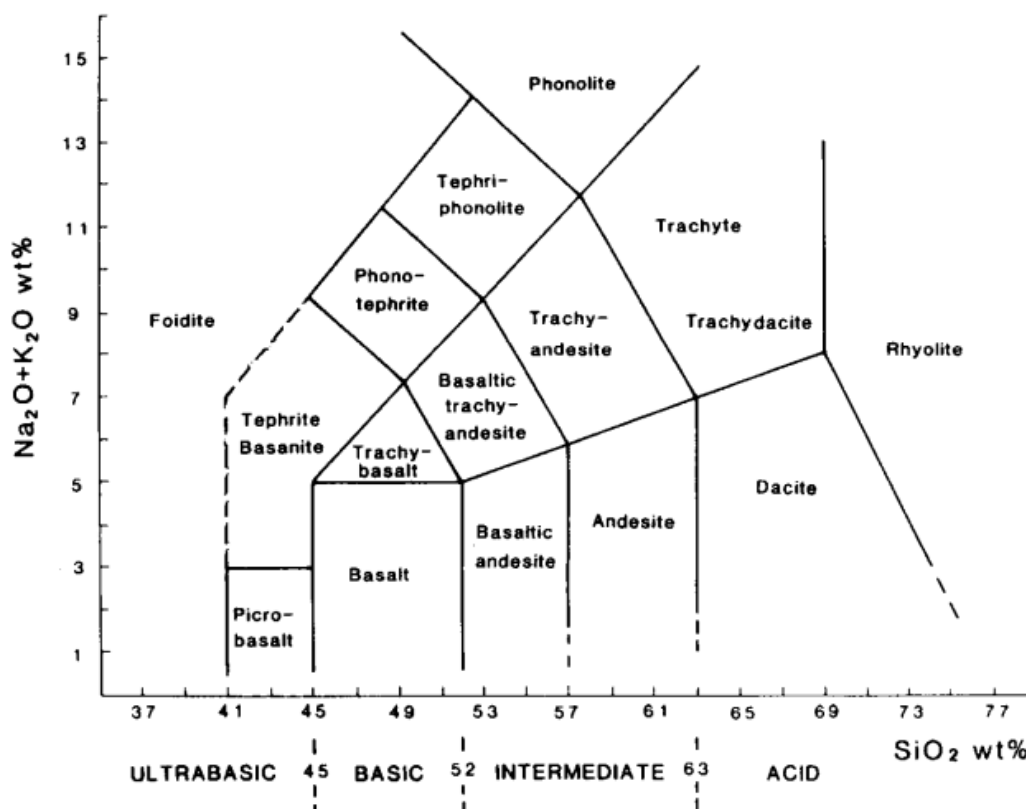
### 3.2 Silicate melts

Silicate melts consist primarily of silicon, oxygen, aluminium, alkalis, magnesium, titanium and iron. The basic unit is a silicon atom in tetrahedral co-ordination with four oxygen atoms. These silicate tetrahedra are usually found in conjunction with other silicate tetrahedra or charge balancing cations. A pure silicate melt consists only of Si and O and forms an array of silicate tetrahedra linked together at all corners by strong Si-O covalent bonds (Bottinga & Weill 1972). In the presence of other elements in the magma, the silicate tetrahedra network becomes altered. Network formers are those elements which can substitute for Si in tetrahedral sites without affecting the melt structure. Network formers in silicate melts include Ti<sup>4+</sup>, Al<sup>3+</sup> and Fe<sup>3+</sup>. The lower charges of Al and Fe means that they can only act as network formers in the presence of other cations, which are required to maintain charge balance. If the supply of charge-balancing cations is exhausted, Al and Fe will act as network modifiers.

Other network modifiers include Na, K, Mg and Ca, which bond with O to form metal oxides (MO). Network modifiers cause Si-O-Si linkages to break, with some Si-O bonds replaced with M-O bonds. This weakens the adjacent Si-O bonds because the cations polarise the shared O (Bottinga and Weil, 1972).

### 3.2.1 Melt classification

A silicate melt entirely composed of network formers in tetrahedral sites is fully polymerised. Melt structure becomes increasingly depolymerised with addition of network modifiers as they break up the tetrahedral network (Bottinga and Weil, 1972). The degree of polymerisation controls the melt viscosity. Silicate melts are classified by their composition, typically by the proportions of SiO<sub>2</sub> and alkalis (Na and K). Generally, melt polymerisation increases with increasing SiO<sub>2</sub> content and decreases with increased Alkali content. Total Alkali Silica (TAS) diagram (Fig 3.1) are used to represent these properties.



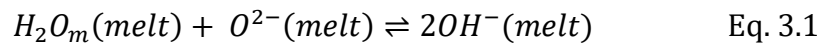
**Figure 3.1: Total Alkali Silica (TAS) plot.** Figure obtained from Le Bas and Streckeisen (1991).

## 3.3 Magmatic volatile

Naturally occurring silicate melts will contain a variety of dissolved volatiles. These species are weakly soluble in the melt. The most common volatiles are H<sub>2</sub>O and CO<sub>2</sub>. Other volatiles include S, Cl and F-bearing species.

### 3.3.1 H<sub>2</sub>O speciation in silicate melt

In silicate melts, total H<sub>2</sub>O (H<sub>2</sub>O<sub>t</sub>) is dissolved in two species, molecular H<sub>2</sub>O (H<sub>2</sub>O<sub>m</sub>) and dissociated hydroxyl groups (OH). H<sub>2</sub>O<sub>m</sub> is neutral and does not affect the melt structure. OH groups are anions (OH<sup>-</sup>) and act as network modifiers. Stolper (1982) showed these two species interconvert via the equilibrium reaction:



in which H<sub>2</sub>O<sub>m</sub> reacts with a bridging oxygen (O<sup>0</sup>) to form OH groups that are bounded to the silicate network and act as network modifiers (Stolper, 1982).

The H<sub>2</sub>O speciation reaction (Eq. 3.1) is a reversible reaction. The equilibrium position of the reaction (*K*), represents what proportions of the two species are expected to be present in a range of given conditions. This equilibrium constant (*K*) for H<sub>2</sub>O can be defined as

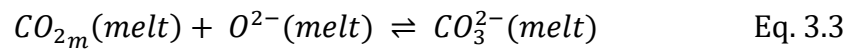
$$K = \frac{X_{OH}^2}{X_{H_2O_m} X_O} \quad \text{Eq. 3.2}$$

where *X* represents the mole fractions on a single oxygen basis. The value of *K* determines the relative proportions of H<sub>2</sub>O<sub>m</sub> and OH in a melt for a given H<sub>2</sub>O<sub>t</sub> content when the system is in equilibrium, and the proportions vary with H<sub>2</sub>O<sub>t</sub> content. At low H<sub>2</sub>O<sub>t</sub>, OH is the dominant species, but H<sub>2</sub>O<sub>m</sub> becomes dominant at

higher  $H_2O_t$ . Higher  $K$  values indicate higher proportions of OH, which requires heat input to break bonds and thus is an endothermic process. In addition, as  $K$  varies with pressure, temperature and melt composition, equilibrium speciation also varies (Stolper 1982; Stolper 1989; Silver et al. 19990; Hui et al. 2008).

### 3.3.2 CO<sub>2</sub> speciation in silicate melt

CO<sub>2</sub> is the second most abundant volatile in natural silicate melts. CO<sub>2</sub> has two distinct species in silicate melts, molecular CO<sub>2</sub> (CO<sub>2m</sub>) and carbonate groups (CO<sub>3</sub><sup>2-</sup>). CO<sub>2m</sub> is neutral and does not affect melt structure, but CO<sub>3</sub><sup>2-</sup> reacts with network modifying cations therefore is not associated with depolymerisation of the melt (Fine and Stolper, 1985; Lowenstern 2001). These two species interconvert via the equilibrium reaction:



in which CO<sub>2m</sub> reacts with bridging oxygen to form CO<sub>3</sub><sup>2-</sup>. The equilibrium constant for CO<sub>2</sub> is as follows:

$$K = \frac{X_{CO_3^{2-}}}{X_{CO_{2m}}X_{O^{2-}}} \quad \text{Eq. 3.4}$$

where  $X$  represents the mole fraction. In contrast to dissolved H<sub>2</sub>O in silicate melts, in which the species concentration ratio depends strongly on  $H_2O_t$ , as total CO<sub>2</sub> varies,  $CO_{2m}/CO_{2total}$  and  $CO_3^{2-}/CO_{2total}$  are roughly the same (Zhang et al. 2007), although they do vary with melt composition. In rhyolitic melt, CO<sub>2</sub> is present as CO<sub>2m</sub> and in basaltic melt, CO<sub>2</sub> is present as CO<sub>3</sub><sup>2-</sup> (Zhang et al. 2007).

## 3.4 Temperature estimation

### 3.4.1 Equilibrium speciation temperature

Temperature is the strongest control on  $K$  for any composition. The relationship with  $K$  and temperature ( $T$ ) is expressed by:

$$\ln K = A + \frac{B}{T} \quad \text{Eq. 3.5}$$

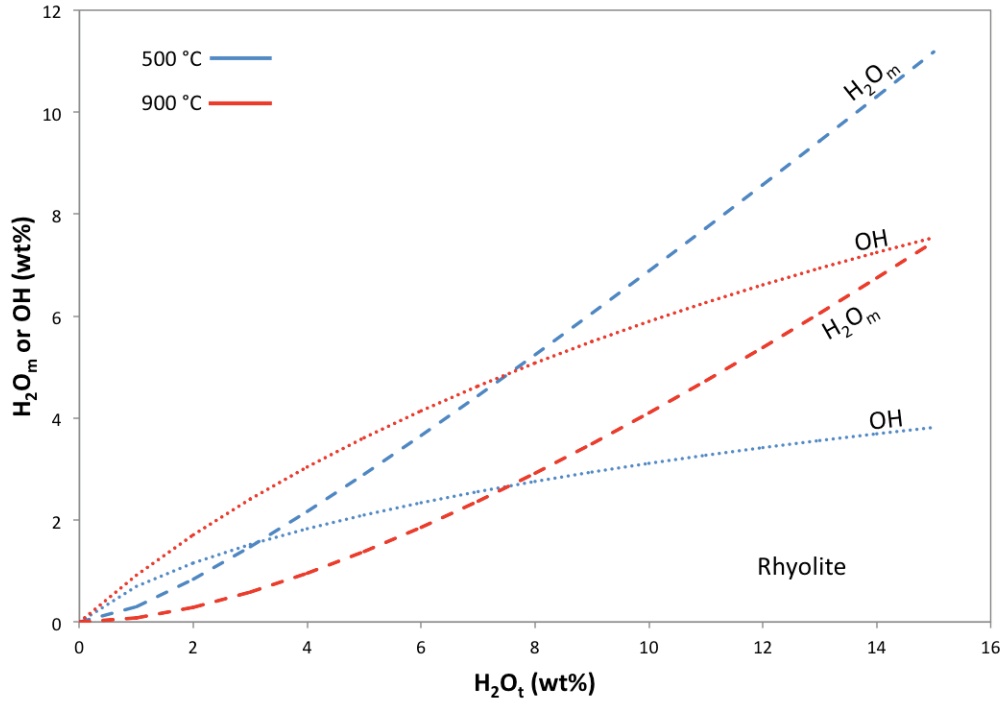
where  $A$  and  $B$  are constants related to the standard state entropy and enthalpy of Eq. 3.1 (Zhang and Ni 2010).

Studies of the effect of  $K$  with temperature are impacted by the quench effect (Section 3.4.2), which limits the temperatures that can be investigated as high temperature melts have to be quenched to glass for analysis at room temperature. Measurements of high temperature melts at room temperature are difficult because melt density and molar absorptivity coefficients vary with temperature (Zhang and Ni 2010). A study by Nowak and Behrens (2001) that investigated  $K$  in haplogranite is calibrated for 500 – 800 °C, 100 – 300 MPa and 1.27-5.15 wt% H<sub>2</sub>O<sub>t</sub>. Fig 3.2 shows the variation in speciation using the  $K$  derived by Nowak and Behrens (2001).

$K$  also varies with composition, however, the variation for melt composition is less than for temperature. Zhang and Ni (2010) showed rhyolite melts exhibited a sharper decrease in  $K$  with temperature than dacite and andesite melts. The constants  $A$  and  $B$  in Eq. 3.5 vary with composition, however, the relationship between  $K$  and melt composition is not well characterised. Data for other compositions, such as basalt, trachyte and phonolite, are limited.

Hui et al. (2008) investigated the effect of pressure on  $K$ . They found there to be no dependence. The effect of pressure at <1 GPa is negligible therefore, the pressure effect can be considered negligible.





**Figure 3.2: Variation in equilibrium speciation with temperature and  $H_2O_t$  concentration.** Equilibrium speciation contours taken from Nowak and Behrens (2001). The data for the anhydrous glass composition used to calculate  $K$  are from the EPMA analysis on the Krafla glass cuttings (Chapter 5).

### 3.4.2 Quenching effect

The equilibrium speciation data obtained by experimental studies of melts at temperature above  $\sim 1000$  K (eg., Silver and Stolper 1989; Stolper 1989; Silver et al. 1990) are now known to have been affected by the quench effect (Zhang et al. 2010). During the initial stages of quenching the  $H_2O_m$ -OH interconversion reaction (Eq. 3.1) keeps pace with cooling and maintains equilibrium speciation until the reaction slows at lower temperatures. This results in a final volatile speciation that reflects an equilibrium at a cooler intermediate temperature rather than the original experimental temperature. Speciation data affected by the quench effect can be described by the apparent equilibrium constant,  $Q$ :

$$Q = \frac{[OH]^2}{[H_2O_m][O]} \quad \text{Eq. 3.6}$$

The expression of  $Q$  is identical to the expression of  $K$  (Eq. 3.4), except  $K$  is only appropriate when the system is in equilibrium (Zhang and Ni, 2010). For a volcanic glass affected by the quench effect, the measured speciation will have an apparent equilibrium constant  $Q$  that is lower than the true equilibrium constant,  $K$ , of the original hotter temperature. Therefore, volcanic glasses affected by the quench effect have higher  $\text{H}_2\text{O}_m/\text{OH}$  ratios at higher temperature. The magnitude of the difference between  $K$  and  $Q$  is determined by the quench rate, with bigger differences caused by slower quench (Fig 3.3).

Like  $K$ ,  $Q$  can be related to temperature as follows,

$$\ln Q = A + \frac{B}{T_{ae}} \quad \text{Eq. 3.7}$$

where  $T_{ae}$  is the apparent equilibrium temperature.  $T_{ae}$  is calculated from  $\ln Q$  assuming  $Q$  is the same as  $K$  even though equilibrium is not reached at any temperature. Fig 3.3 illustrates the variation of  $Q$  and  $T_{ae}$  as a function of temperature during cooling at various quench rates (Zhang and Ni, 2010).

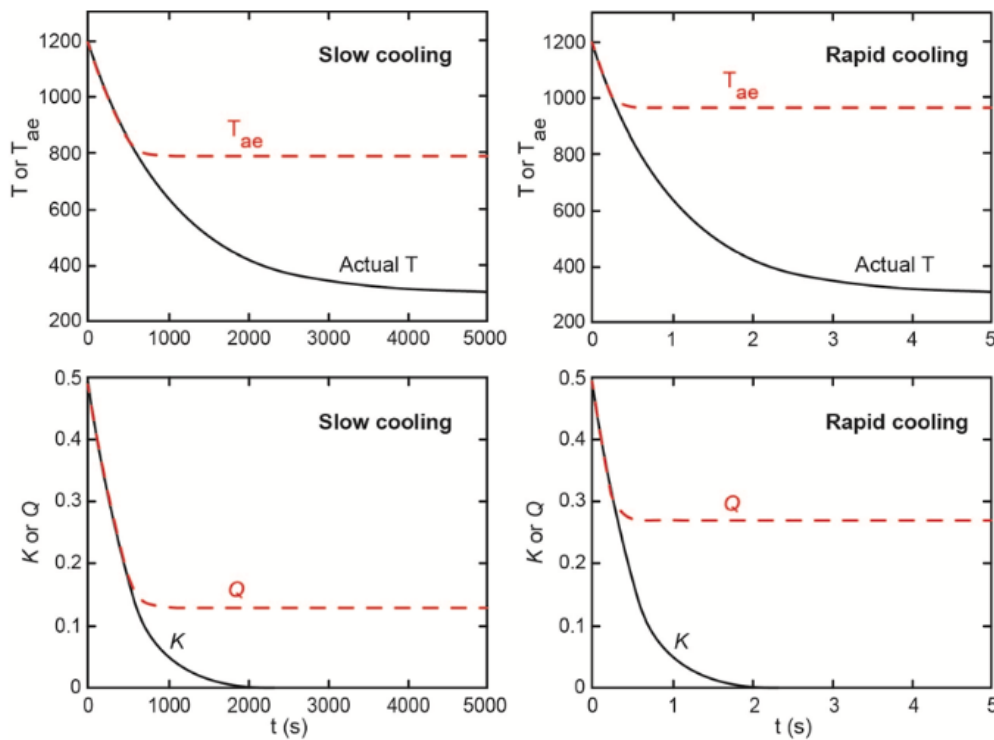


Figure 3.3: **Schematic evolution of  $Q$  and  $T_{ae}$  for rapid cooling and slow cooling of a hypothetical reaction.** Figure from Zhang and Ni (2010).

## 3.5 Melt viscosity

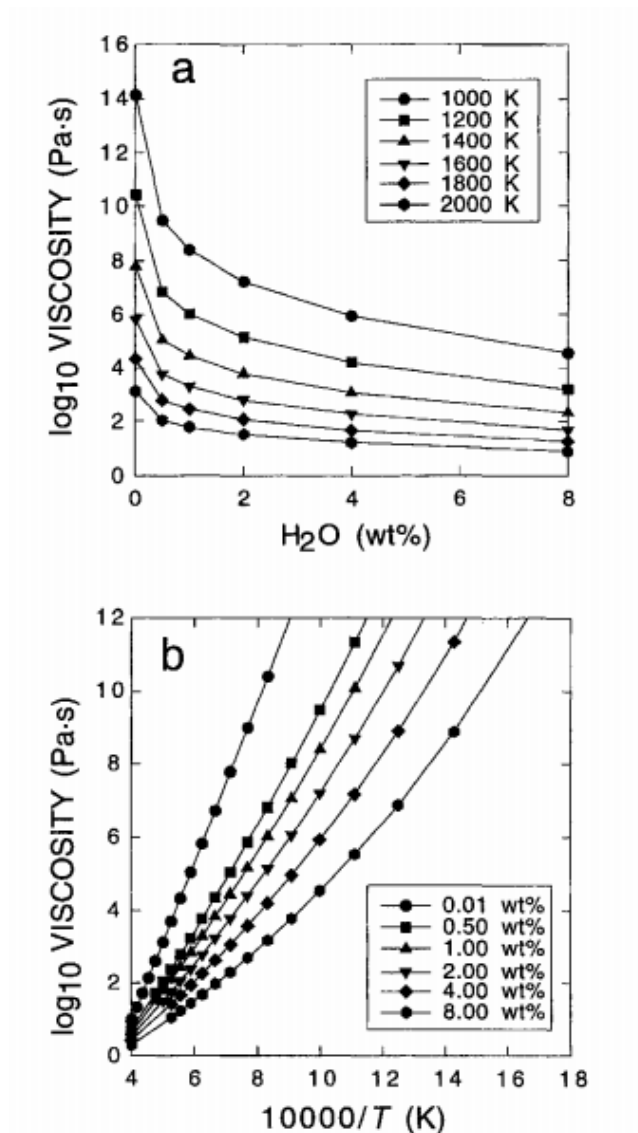
### 3.5.1 Effect of H<sub>2</sub>O and CO<sub>2</sub>

Viscosity is a measure of a melt's resistant to flow (Pa s). Dissolved OH groups act as network modifiers that reduce melt polymerisation, and therefore melt viscosity decreases with increasing H<sub>2</sub>O<sub>t</sub>. Studies by Hess and Dingwell (1996) have shown that viscosity of rhyolite melts vary by more than seven orders of magnitude for H<sub>2</sub>O<sub>t</sub> concentrations of 0.01 to 8 wt% (Fig 3.4a). The decrease of viscosity with increasing H<sub>2</sub>O<sub>t</sub> is the most significant with the addition of the first 1 wt% of H<sub>2</sub>O<sub>t</sub> (Hess and Dingwell, 1996). The effect diminishes for every additional wt%.

Data on the effect of CO<sub>2</sub> on silicate melt viscosity is limited with studies indicating that CO<sub>2</sub> does not appear to affect the viscosity (Lange, 1994; Morizet et al. 2007). It is believed that it has no effect because its dissolution does not alter the fundamental silicate melt structure. However, in recent study by Morizet et al. (2017) have measured glass transition temperature (section 3.6.1) in low silica melt decreases with increasing CO<sub>2</sub> content, therefore viscosity decreases. This contradicts the melt polymerisation because CO<sub>2</sub> are neutral in melt. However, since the Krafla melt is high in silica, the effect CO<sub>2</sub> on melt viscosity would not be considered.

### 3.5.2 Effect of temperature

Temperature is the strongest control on melt viscosity, where melt viscosity increases with decreasing temperature. Melt viscosity displays non-Arrhenian behaviour (non-linear) with temperature, usually expressed as  $1/T$  (Hess and Dingwell, 1996). Fig 3.4b shows the variation of viscosity with temperature, with all melt compositions showing a decrease of viscosity with increasing temperature. The exact nature of this temperature-dependence varies with the anhydrous composition and volatile content of the melt.



**Figure 3.4: Variation of rhyolite viscosity with temperature and  $H_2O$  concentration.** a) Variation of viscosity with  $H_2O$  concentration at different temperatures. b) Variation of viscosity with temperature at constant  $H_2O$  concentration. Figure from Hess and Dingwell (1996).

## 3.6 Melt to glass

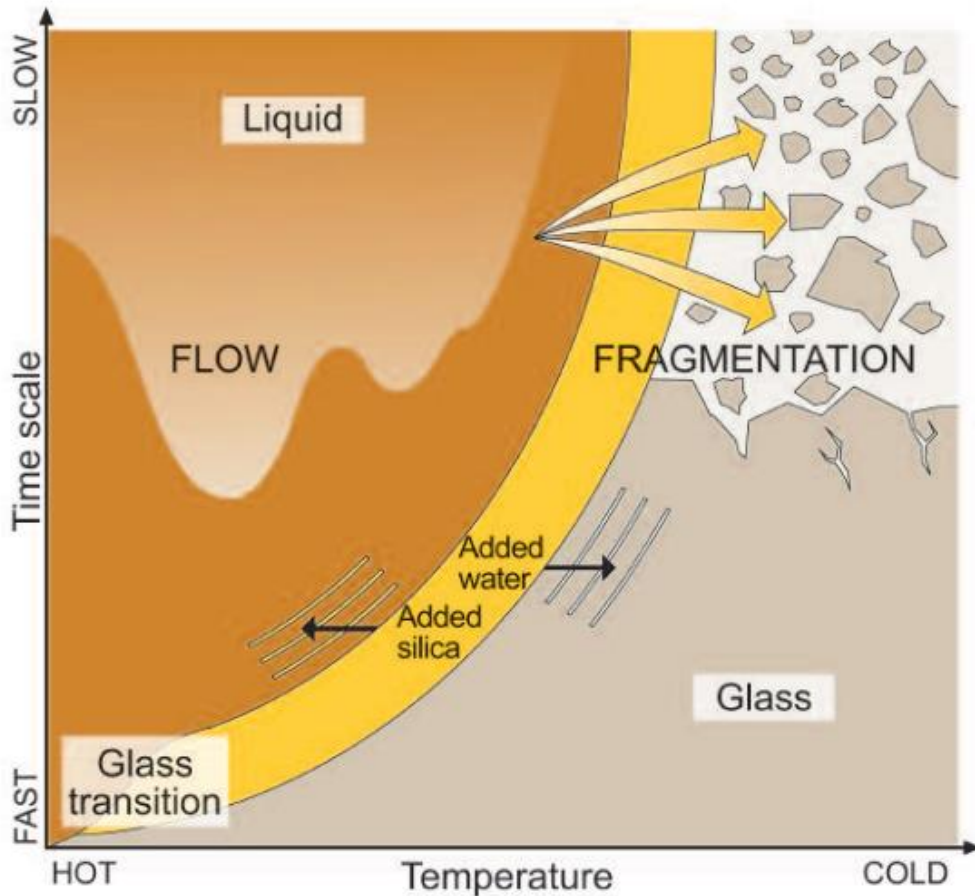
The transition between liquid-like and glass-like behaviour over a temperature interval is known as the glass transition (Fig. 3.5). The transition from liquid to glass occurs when the liquid is unable to achieve equilibrium due to external perturbation (Dingwell & Webb 1989). In volcanic contexts this is most likely due to rapid cooling, or from deformation during flow. The ability of the melt to

maintain an equilibrium state is controlled by the structural relaxation timescale of the melt. Structural relaxation timescale is controlled by the rate of self-diffusion of Si and O atoms in the liquid matrix (Dingwell 2006). The self-diffusion involves breaking Si-O bonds (Dingwell and Webb, 1989). If the timescale of perturbation is more rapid than this structural relaxation timescale, the silicate structure is unable to deform rapidly enough to maintain an equilibrium liquid state and the melt begins to act like a solid and therefore becomes a glass.

The structural relaxation timescale of a melt is given by the Maxwell relationship:

$$\tau_s = \frac{\eta_s}{G^\infty} \quad \text{Eq. 3.8}$$

where  $\tau_s$  is the structural relaxation timescale,  $\eta_s$  is the shear viscosity and  $G^\infty$  is the shear modulus and can be considered to be constant ( $\sim 10^{10}$  Pa; Dingwell and Webb, 1989). The structural relaxation timescale of a melt varies with temperature and volatile content, like melt viscosity.



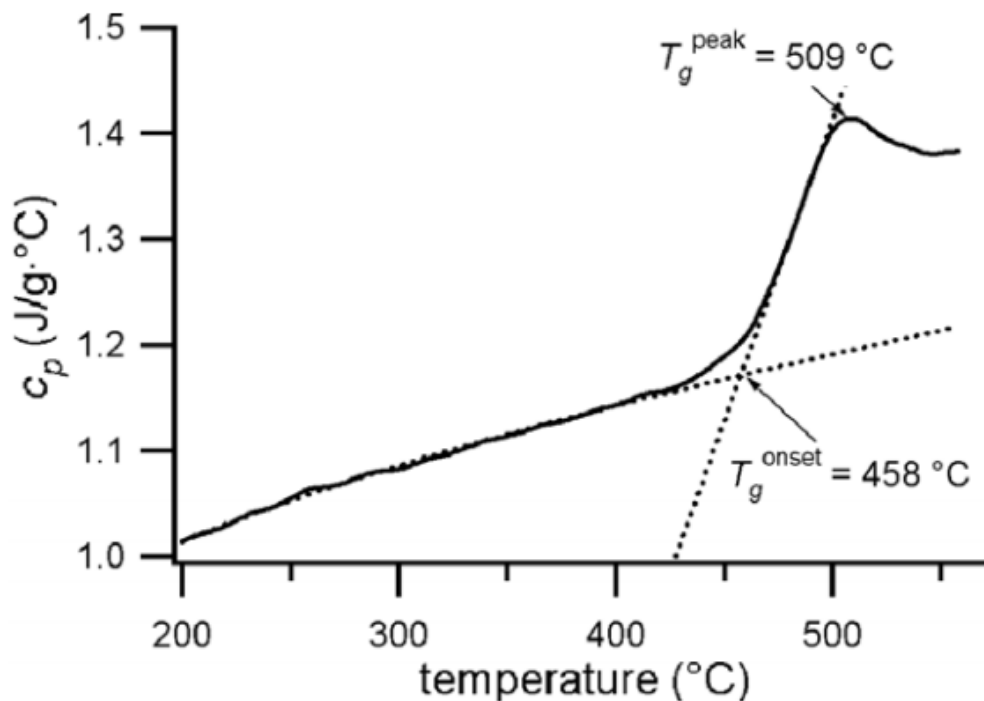
**Figure 3.5: Schematic drawing of glass transition for a silicate melt.** The glass transition separates the liquid and solid response of the molten state. Figure from Dingwell and Webb (1989).

### 3.6.1 Glass transition temperature

The glass transition occurs over a temperature interval, the exact definition of  $T_g$  differs depending on how it is determined. The two most common techniques used to determine  $T_g$  are dilatometry and calorimetry. In dilatometry studies,  $T_g$  is defined as the temperature at which the melt viscosity is  $10^{12}$  Pa s (Giordano et al. 2008). In calorimetric studies, a glass sample is repeatedly heated and cooled across the glass transition at different heating/cooling rates. This method is able to measure the variation in the specific heat capacity ( $c_p$ ) of the glass with temperature. Fig 3.6 shows a result from the study by Giordano et al. (2005) for determining the  $T_g$  using a calorimetry method. When the transition from glass to liquid is reached there is a rapid increase in  $c_p$ . The glass transition is defined

as the temperature range between which  $c_p$  deviates from the glassy state and the temperature at which  $c_p$  is in a constant state, and from this two definitions of  $T_g$  are obtained (Giordano et al., 2005). First, is  $T_{g \text{ peak}}$ , which is the temperature at which  $c_p$  is at the highest value. The second is  $T_{g \text{ onset}}$ , which is intersecting the extrapolations of the gradual  $c_p$  in the glassy state and the tangent of the increase in  $c_p$  during the glass transition interval.

Calorimetry studies have shown that the value of  $T_g$  is influenced by the heating/cooling rate. Slower cooling rates create a longer effective structural relaxation timescale and therefore lower  $T_g$  (Dingwell and Webb 1989; Giordano et al. 2005)



**Figure 3.6: Determination of  $T_g$  using a calorimetry method.** Figure from Giordano et al. 2005.

### 3.6.2 Controls on glass transition temperature

$T_g$  is affected by quench rate, with  $T_g$  being higher for faster quench rates and lower for slow quench rates. Like viscosity,  $T_g$  varies with composition including volatile content. Volatiles dissolved in the melt can alter the structural relaxation timescale of the melt, resulting in an increase of volatile decreasing  $T_g$  (Giordano et al. 2005).

## 3.7 Solubility

Solubility is the property of a substance that can be dissolved in a solvent at equilibrium. For this project the solvent is the silicate melt and the solute is H<sub>2</sub>O and CO<sub>2</sub> species. If the silicate melt is over-saturated with the volatile then two phases, (melt and fluid) exist in equilibrium. The term fluid is used since many temperature and pressure conditions relevant to volcanology will be above the critical point (Barker and Alletti, 2012). Henry's Law quantifies the solubility of a gas in a solvent and shows that at a fixed temperature the solubility of a gas is proportional to its partial pressure:

$$C = k P_{gas} \quad \text{Eq. 3.9}$$

where  $C$  is the concentration in the solvent,  $P_{gas}$  is the partial pressure of the gas, and  $k$  is Henry's Law constant, which depends on the solvent, temperature and gas.

### 3.7.1 H<sub>2</sub>O solubility in silicate melt

Due to the non-unity stoichiometric coefficient for OH in Eq. 3.1 and H<sub>2</sub>O speciation, the dependency of H<sub>2</sub>O solubility on pressure is non-Henrian (Zhang et al. 2007). The solubility of total H<sub>2</sub>O can be written as;



$$X_{H_2O_t} = \frac{X_{OH}}{2} + X_{H_2O_m} \quad \text{Eq. 3.10}$$

where  $X$  denote mole fractions of the dissolved species in the melt. From the gas-liquid equilibrium of  $H_2O$  (gas)  $\rightleftharpoons H_2O$  (melt) and assuming ideal gas behaviour for the gas phase and ideal mixing in the melt,  $X_{H_2O_m}$  is proportional to  $P_{H_2O}$ .  $X_{OH}$  is proportional to square root of  $X_{H_2O_m}$  (Eq. 3.2) and therefore is proportional to  $\sqrt{P_{H_2O}}$ . The solubility of  $H_2O_t$  in silicate melt depends on pressure as;

$$X_{H_2O_t} = k_1\sqrt{P_{H_2O}} + k_2P_{H_2O} \quad \text{Eq. 3.11}$$

where the first term on the right-hand side of the equation results from OH groups and the second term results from  $H_2O_m$  (Zhang et al. 2007). According to Eq. 3.11 at low pressures the dissolved  $H_2O_t$  is low and OH will be the dominant species. As pressure increases,  $H_2O_t$  also increases, whereby the proportion of  $H_2O_m$  will increase. As  $H_2O_m$  species becomes more dominant,  $H_2O_t$  solubility will begin to increase more rapidly than if it continues to follow a square root dependency on pressure (Zhang et al. 2007).

### 3.7.2 CO<sub>2</sub> solubility

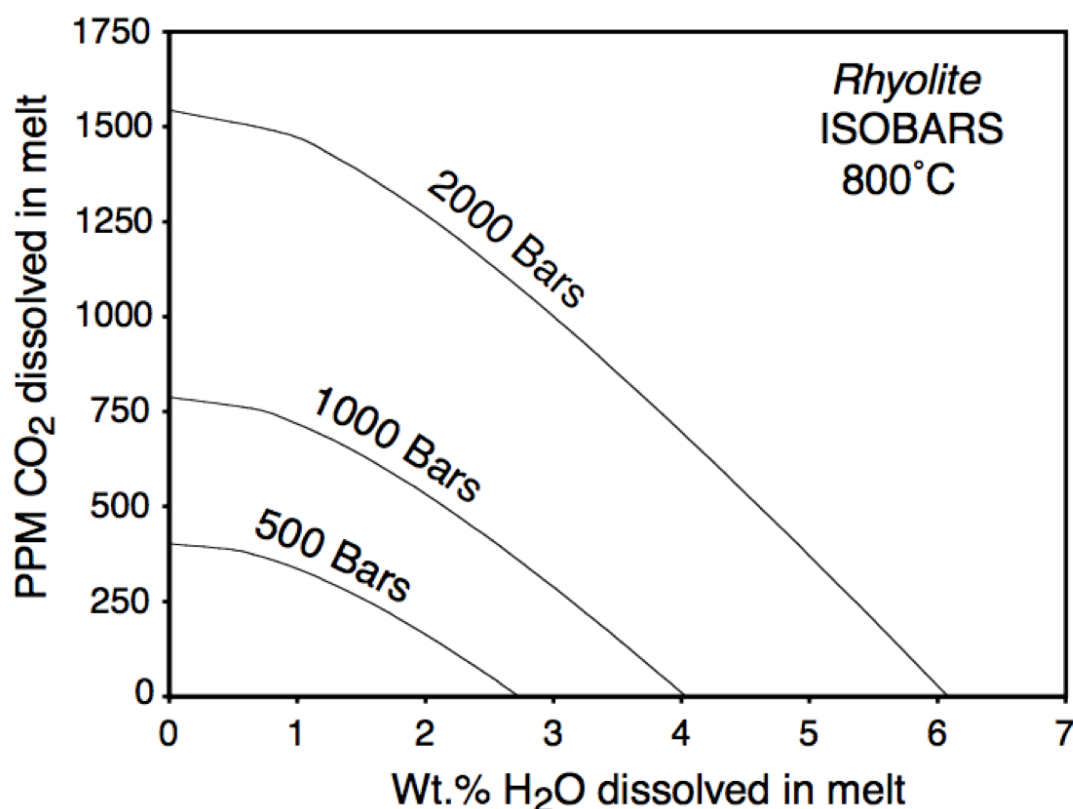
For CO<sub>2</sub> although there are also two species like  $H_2O$ , it is more simple as the concentration of the two species are simply proportional to each other (Eq. 3.3). Therefore the solubility of CO<sub>2</sub> follows Henry's law (Eq. 3.9), (Zhang and Ni, 2010).

### 3.7.3 H<sub>2</sub>O-CO<sub>2</sub> solubility

Silicate melt is unlikely to be in equilibrium with a fluid consisting of a single volatile species. The presence of additional volatile species would lower the

solubility of H<sub>2</sub>O compared to a melt that is in equilibrium with a pure H<sub>2</sub>O fluid (Barker and Alletti, 2012). Experimental studies have shown that H<sub>2</sub>O solubility can be assumed to be independent of CO<sub>2</sub> concentration, for pressures below 400 MPa (Blank et al. 1993). However, addition of CO<sub>2</sub> gas lowers the concentration of H<sub>2</sub>O in a melt by diluting the H<sub>2</sub>O vapour phase (Lowenstern 2001).

For rhyolitic melts the model VolatileCalc by Newman and Lowenstern (2002), can model the solubility of mixed H<sub>2</sub>O-CO<sub>2</sub> fluids in different composition, pressure and temperature conditions (Fig 3.7). The VolatileCalc model is valid for  $\leq 11$  wt% H<sub>2</sub>O<sub>t</sub> and 10,000 ppm CO<sub>2</sub>, pressures of  $\leq 500$  MPa and temperature in a range of 600 – 1500 °C.



**Figure 3.7: Example of pressure estimates from VolatileCalc. Isobars represents range of values for dissolved H<sub>2</sub>O and CO<sub>2</sub> in rhyolitic melt in equilibrium with H<sub>2</sub>O-CO<sub>2</sub> gas at 800 °C and selected pressures. Figure from Newman and Lowenstern (2002).**

## 3.8 Diffusion

Diffusion occurs as a result of the thermally activated random motion of particles such as atoms, ions or molecules, which leads to a net flux of particles when the concentration of a component is not uniform (Zhang and Ni, 2010). This net flux is described by Fick's first law:

$$J = -D \frac{\partial C}{\partial x} \quad \text{Eq. 3.12}$$

where  $J$  is the diffusive mass flux (a vector),  $D$  is the diffusion coefficient ( $\text{m}^2/\text{s}$ ),  $C$  is concentration in units of mass per unit volume and  $\frac{\partial C}{\partial x}$  is the concentration gradient (a vector). The negative sign indicates that flux moves from high to low concentration. Therefore, large concentration gradients are associated with large diffusive fluxes.

To understand how a concentration profile will evolve, the relationship between concentration, space and time is necessary. The general equation for one-dimensional diffusion is given by Fick's second law as;

$$\frac{\partial C}{\partial t} = \frac{\partial}{\partial x} \left( D \frac{\partial C}{\partial x} \right) \quad \text{Eq. 3.13}$$

where  $t$  is time,  $x$  is distance,  $D$  is constant.

### 3.8.1 Diffusion of $\text{H}_2\text{O}$

As already discussed,  $\text{H}_2\text{O}$  is present in the melt as both  $\text{H}_2\text{O}_m$  and  $\text{OH}$  whose proportions vary according to the equilibrium interconversion reaction (Eq. 3.1). Fick's second law can be modified to give the diffusion of  $\text{H}_2\text{O}_t$  as the sum of the diffusion of each species (Zhang and Ni, 2010):

$$\frac{\partial X}{\partial t} = \frac{\partial}{\partial x} \left( D_{\text{H}_2\text{O}_m} \frac{\partial X_{\text{H}_2\text{O}_m}}{\partial x} + \frac{1}{2} D_{\text{OH}} \frac{\partial X_{\text{OH}}}{\partial x} \right) \quad \text{Eq. 3.14}$$

where  $X_{H_2O_m}$  is the mole fraction of  $H_2O_m$ ,  $X_{OH}$  is the mole fraction of OH and  $D_{H_2O_m}$  and  $D_{OH}$  are the diffusivities of  $H_2O_m$  and OH. The factor of  $\frac{1}{2}$  accounts for the 2 mol of OH that are produced with every 1 mol of  $H_2O_m$  (Eq. 3.1).

Shaw (1974) carried out the first experimental study on  $H_2O$  diffusion. Shaw (1974) investigated the mass gain of obsidian melt during hydration and concluded that  $H_2O_t$  diffusivity ( $D_{H_2O_t}$ ) is dependent on the concentration of  $H_2O_t$ . The next major advance in  $H_2O$  diffusion study was the development of infrared spectroscopy, which allowed the microanalytical determination of the concentration of  $H_2O_t$  and the two species,  $H_2O_m$  and OH (Stolper 1982, Newman et al. 1986). Zhang and Stolper (1991) measured  $H_2O_m$  and OH concentration profiles in experimentally dehydrated obsidians. The results of the experimental profiles show that  $D_{OH}$  is much slower than  $D_{H_2O_m}$  and  $D_{H_2O_m}$  is roughly the same as for  $D_{H_2O_t}$ . This suggests that for  $H_2O_t$ ,  $H_2O_m$  is the mobile species and OH is immobile (Zhang and Stolper 1991; Zhang et al. 2007), essentially the diffusion of  $H_2O_m$  is controlling the diffusion of  $H_2O_t$ . Therefore, OH concentration profiles are formed indirectly by diffusion of  $H_2O_m$  and then the conversion of  $H_2O_m$  to OH via Eq 3.1. The results of the difference in diffusivity may be explained by  $H_2O_m$  being a small, neutral molecule that can move relatively easily whereas the OH groups are bounded to a cation and therefore cannot move easily (Zhang and Stolper 1991).

### 3.8.2 Diffusion of $CO_2$

The diffusivity of  $CO_2$  does not depend on its own concentration and is roughly independent of the anhydrous melt composition. Diffusivity of  $CO_2$  in melts are expressed by the following (Zhang et al. 2007):

$$\ln D_{CO_2t}^{all\ melts} = -13.99 - \frac{17367+1944.8P}{T} + \frac{(855.2+271.2P)}{T} C_w \quad \text{Eq. 3.15}$$

where  $D$  is in  $m^2/s$ ,  $T$  is in K,  $P$  is in GPa, and  $C_w$  is the wt% of  $H_2O_t$ .

### 3.8.3 Comparison of CO<sub>2</sub> and H<sub>2</sub>O diffusivities

$D_{H_2O_m}$  is greater than  $D_{CO_{2t}}$ , and  $D_{H_2O_t}$  is usually greater than  $D_{CO_{2t}}$ . This can be explained by the size of the molecules, where the effective radius of H<sub>2</sub>O<sub>m</sub> is smaller than the effective radius of CO<sub>2m</sub>.

The diffusion of H<sub>2</sub>O and CO<sub>2</sub> shows similarities and distinctions. In both H<sub>2</sub>O and CO<sub>2</sub>, speciation plays an important role where the neutral species (H<sub>2</sub>O<sub>m</sub> and CO<sub>2m</sub>) is the dominant diffusing species and diffusivity of these species increases with H<sub>2</sub>O<sub>t</sub>. The major differences are that H<sub>2</sub>O speciation results in  $D_{H_2O_t}$  that depends strongly on H<sub>2</sub>O<sub>t</sub>. CO<sub>2</sub> speciation results in CO<sub>2</sub> diffusivity that is independent of CO<sub>2t</sub> (Zhang and Ni 2010).

## 3.9 Bubbles in melt

### 3.9.1 Bubble formation

Formation of bubbles occurs when volatiles dissolved in a melt become over-saturated and exsolve to form a gas phase. Bubble formation may occur when there is a change in pressure and/or temperature in the melt causing the equilibrium solubility value to fall below the concentration of volatiles dissolved in the melt, or when there is addition of volatiles to the melt, which increases the concentration of dissolved volatiles above the equilibrium volatile solubility of the system (Mourtada-Bonnetoi and Laporte, 1999; Mangan and Sisson, 2000). In volcanism, a decrease in the confining pressure of the system caused by magma ascent or magma depressurisation results in a reduction in the amount of volatiles that can be dissolved in the magma and they exsolve to form bubbles. Bubble formation can also occur with growth of anhydrous (H<sub>2</sub>O deprived) crystal phases. This increases the concentration of H<sub>2</sub>O dissolved in the remaining melt and can drive them above the equilibrium solubility value.

Bubble formation can occur by homogeneous and/or heterogeneous nucleation. Homogeneous bubble nucleation is the nucleation of a bubble in a crystal and bubble free melt. Homogeneous nucleation may therefore require oversaturation pressures of up to 200 MPa (Mourtada-Bonnetoi and Laporte, 1999). Heterogeneous nucleation occurs in the presence of a nucleation sites such as a crystal, microlite, inclusion or xenolith. Heterogeneous bubble nucleation is more efficient and does not require as high supersaturation pressures (only ~5 MPa) (Gardner et al. 1999).

### 3.9.2 Bubble growth

Bubble growth is controlled by both depressurisation and exsolution (Proussevitch and Sahagian, 1996). Depressurisation contributes to bubble growth by oversaturation of the melt in volatiles and a decrease of the density of the volatiles. Volatiles diffuse into bubbles if the melt is oversaturated due to depressurisation, increase of temperature or changes in the composition. During bubble growth, volatile concentrations at the bubble wall will be the equilibrium volatile solubility value for the current local pressure and temperature conditions. Away from the bubble wall the melt will contain more volatiles than the new equilibrium solubility value. As a result, volatiles will diffuse down a concentration gradient towards the bubble wall and into the bubble (McIntosh et al. 2014). Thus, if a bubble is growing, the volatile concentration decreases towards the bubble wall (Fig 3.8b, McIntosh et al. 2014).

### 3.9.3 Bubble resorption

During bubble resorption the volatile concentration at the bubble wall will increase above the concentration of volatile dissolved in the melt due to a decrease in the equilibrium volatile solubility caused by an increase in pressure and/or decrease in temperature (Fig 3.8C). This transfer of volatiles back into the melt will cause the bubble to shrink (Watkins et al 2012; McIntosh et al 2014) and volatiles will diffuse down a concentration gradient away the bubble wall and into the melt (McIntosh et al. 2014).

Studies have investigated bubble resorption in magmatic melts. Yoshimura and Nakamura (2008) have shown that bubble resorption can occur in an open degassing scenario by holding vesicular melt in an apparatus that maintains a pressure difference between the melt and the exterior. In higher-pressure sample chambers, H<sub>2</sub>O gas is capable of escaping, and so H<sub>2</sub>O was lost by diffusion from sample margins resulting in the resorption of bubbles along the margins. Bubble resorption was also investigated in a H<sub>2</sub>O-CO<sub>2</sub> system (Yoshimura and Nakamura, 2010). A H<sub>2</sub>O-rich melt containing bubbles of H<sub>2</sub>O was surrounded by a pure CO<sub>2</sub> fluid, and bubbles were seen to shrink initially as fast-diffusing H<sub>2</sub>O moved from the melt into the surrounding fluid and H<sub>2</sub>O moved from the bubbles into the melt to maintain equilibrium. CO<sub>2</sub> moved from the fluid into the melt and ultimately into the bubbles, however due to the slower timescale of CO<sub>2</sub> diffusion temporary bubble resorption occurred.

Bubble resorption in magmatic contexts have been attributed to an increase in pressure causing resorption. H<sub>2</sub>O concentration profiles around bubbles in obsidian pyroclasts from Mono Craters, California (Watkins et al. 2012), shown that bubbles were resorbing prior to the glass transition. Watkins et al. (2012) interpreted this as evidence of pressure increase cycles in the conduit prior to eruption.

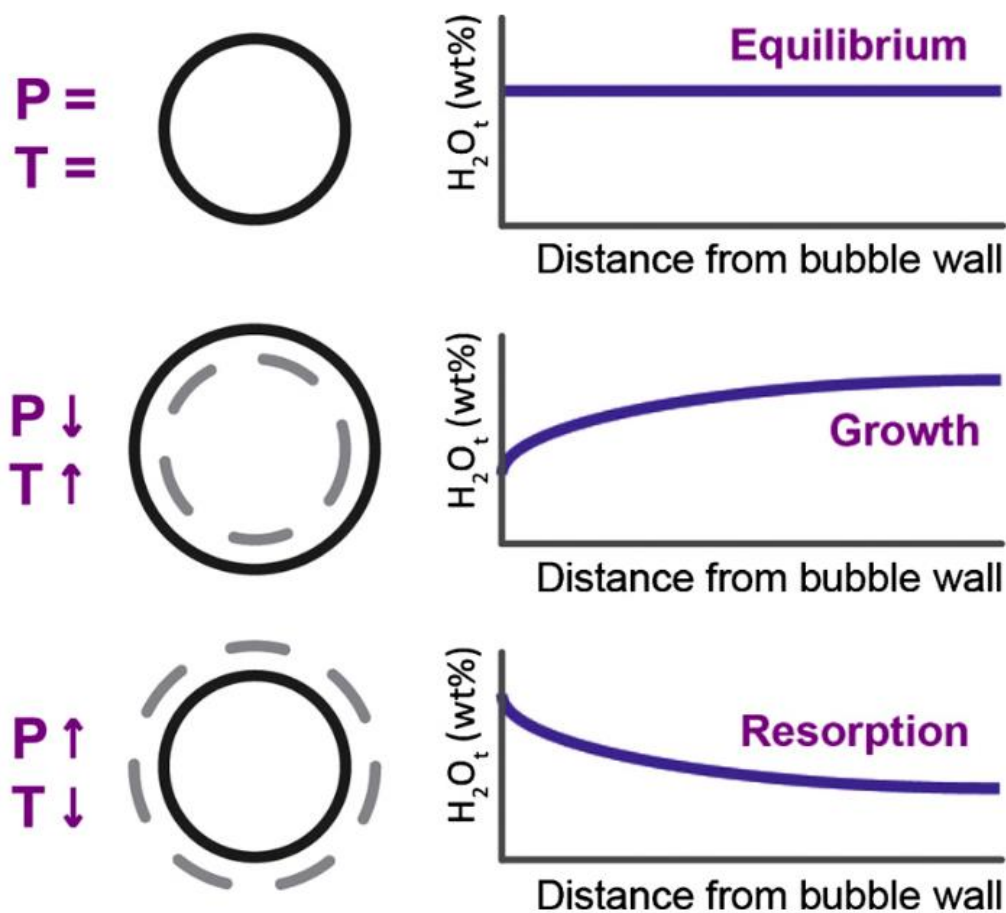
Second study of bubbles resorption been attributed to an increase in pressure by Carey et al. (2013) have also observed in direct evidence of bubble resorption in

basaltic pyroclasts from explosive lake eruption in Kilauea, Hawaii.  $H_2O$  concentrations in the samples were too low to analyse but the bubble resorption was inferred with a halos consisting of small bubbles surrounding a large bubble. This texture is thought to represent bubbles which had begun to resorb as they sank during conduit convection. This causes an increase in pressure. During pressure decrease caused by either as a result of further convection of eruption, the  $H_2O$ -rich melt surrounding the partially resorbed bubble became supersaturated and nucleated a secondary bubble population.

Bubble resorption has also been attributed to cooling of the magmatic melt. McIntosh et al. (2014) have shown if resorption is caused by temperature decreases during quench, previous growth history maybe preserved. Speciation data indicate that the  $H_2O_m$  distribution records resorption, while the OH distribution records earlier decompressive growth.

If a bubble is not growing or shrinking it is in equilibrium with the melt, and the volatile concentration in the surrounding melt is constant throughout from the bubble wall (Fig 3.8a).





**Figure 3.8: Schematic figure linking the pressure and temperature conditions, state of bubble growth and associated concentration profiles in the melt.** Solid black lines indicate current bubble size; dashed grey lines represent the previous bubble size. Figure from McIntosh et al. 2014.

### 3.10 Summary

Many experiments have been carried out investigating many aspects of magmatic volatiles and its influence of the melt and bubble property. This project will look at the evolution of the magmatic  $H_2O$  and  $CO_2$  from the Krafla glass cuttings in response to the unexpected interception by the IDDP-1 drill. Chapter 4 discusses analytical techniques used to achieve an understanding of the insights of the  $H_2O$  and  $CO_2$  content of the Krafla melt and also its melt properties.

# Chapter 4: Analytical methods and data production

## 4.1 Introduction

This chapter provides an overview of how the samples were analysed, and the methodologies used to produce the data presented in Chapters 4 and 5. For this project Fourier transform infrared spectroscopy (FTIR) was used to measure volatile elements, H<sub>2</sub>O and CO<sub>2</sub>, dissolved in the glass and electron probe microanalysis (EPMA) was used to measure major element compositions of the glass.

## 4.2 FTIR measurements of magmatic H<sub>2</sub>O and CO<sub>2</sub>

H<sub>2</sub>O and CO<sub>2</sub> concentrations in the Krafla glass cuttings were analysed by FTIR. As well as total H<sub>2</sub>O (H<sub>2</sub>O<sub>t</sub>) and CO<sub>2</sub> (CO<sub>2t</sub>) concentrations, FTIR can determine the speciation of H<sub>2</sub>O (i.e., the amount of H<sub>2</sub>O present as molecular H<sub>2</sub>O, H<sub>2</sub>O<sub>m</sub>, versus hydroxyl, OH, groups) and CO<sub>2</sub> (i.e., the amount of CO<sub>2</sub> present as molecular CO<sub>2</sub>, CO<sub>2m</sub>, versus carbonate, CO<sub>3</sub><sup>2-</sup>, groups) (see Chapter 3 section 3.3). H<sub>2</sub>O speciation data can be used to investigate the magmatic conditions immediately prior to quenching, including species gradients resulting from pressure changes (e.g., Watkins et al., 2012; Carey et al., 2013), temperature changes (e.g., McIntosh et al., 2014), and post-quenching hydration (e.g., Dixon et al. 1995; Anovitz et al. 2008; McIntosh et al. 2017). FTIR results will therefore help to establish the evolution of magmatic volatiles caused by the magma being intersected by drilling during IDDP-1.

## 4.2.1 Principle

FTIR is a non-destructive technique and can resolve H<sub>2</sub>O and CO<sub>2</sub> concentrations ranging from a few ppm to several wt % (Newman et al. 1986). In FTIR, infrared (IR) radiation is passed through a glass sample of known thickness. As the radiation passes through the sample it interacts with the molecular species present in the glass, triggering transitions in the molecular vibrations of covalent bonds (Stolper 1982). These transitions absorb the wavelength of IR radiation that corresponds to the frequency of the molecular vibration (Stolper 1982; Newman et al. 1986). Different bonds have different molecular vibration frequencies and hence absorb different wavelengths of IR radiation. This allows identification of the specific molecular species present in the samples. These absorption values can be converted into concentrations using the Beer-Lambert law;

$$c = \frac{M.A}{\rho.d.\varepsilon} \times 100 \quad \text{Eq. 4.1}$$

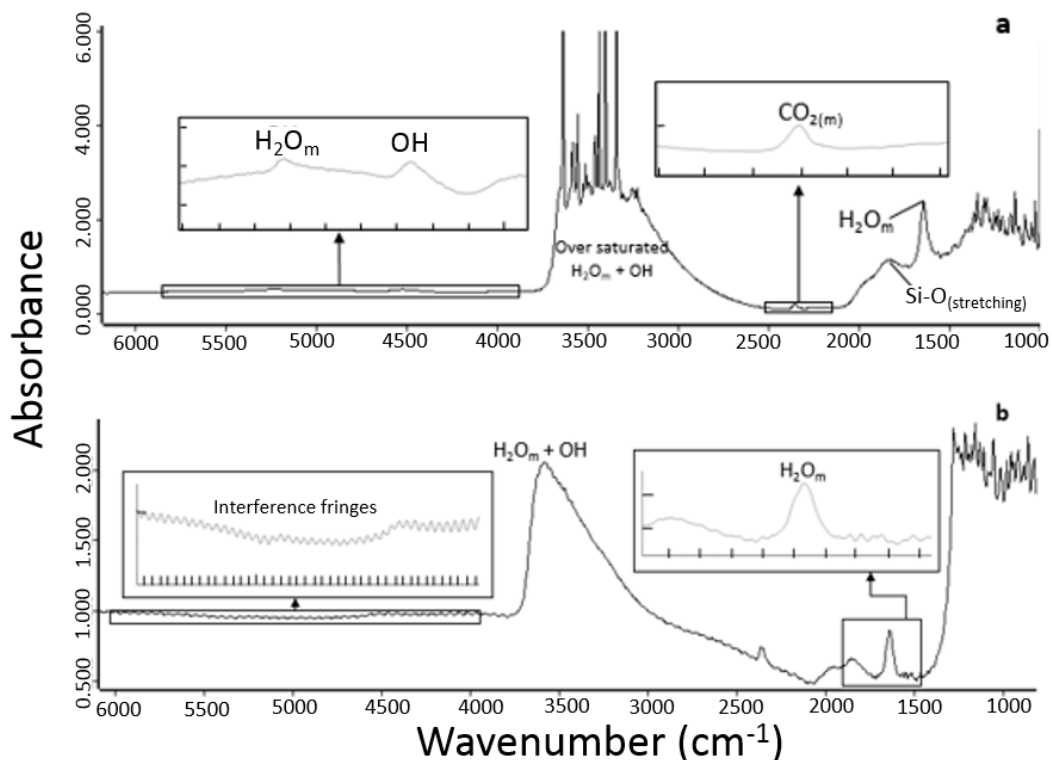
where  $c$  is the concentration of species of interest (wt.%),  $M$  is the molar mass of H<sub>2</sub>O or CO<sub>2</sub> (g.mol<sup>-1</sup>),  $A$  is the absorbance of the species of interest, the height of the peak above background (unitless),  $\rho$  is the density of the sample (kg.m<sup>-3</sup>),  $d$  is sample thickness (cm) and  $\varepsilon$  is the molar absorptivity for the absorbance band of interest (l.mol<sup>-1</sup>.cm<sup>-1</sup>) (Stolper 1982, Newman et al. 1986).

FTIR is especially useful for identifying H-O species as the bonds are highly polarised and therefore are efficient absorbers of IR radiation (Della Ventura et al. 2010). Several different absorption bands exist for the H<sub>2</sub>O species (Fig 4.1):

- In the near infrared (near-IR) region (wavenumbers 14000-4000 cm<sup>-1</sup>), absorption bands for H<sub>2</sub>O exist at ~7100, ~5200, ~4500 and ~4000 cm<sup>-1</sup>, of which the 5200 and the 4500 cm<sup>-1</sup> bands are the strongest and were used in this study. The 5200 cm<sup>-1</sup> band is attributed to the bending and stretching of H<sub>2</sub>O<sub>m</sub> (Scholze 1960; Bartholomew et al. 1980). The 4500 cm<sup>-1</sup> band is attributed to combination modes of Si-OH and Al-OH groups (Stolper 1982). The weaker absorbance band at 7100 cm<sup>-1</sup> is a first

overtone of the O-H stretching, and the origin of the  $4000\text{ cm}^{-1}$  band is not fully understood (Bartholomew et al. 1980; Newman et al. 1986).

- In the mid infrared (mid-IR) region ( $4000\text{--}400\text{ cm}^{-1}$ ) the bands related to  $\text{H}_2\text{O}$  exist at  $\sim 3500$  and  $\sim 1630\text{ cm}^{-1}$ . The  $3500\text{ cm}^{-1}$  band is broad and asymmetric, and attributed to the fundamental OH stretching of both  $\text{H}_2\text{O}_m$  and OH. The shape of the band reflects the distribution of H-bond strength within the glass. The  $1630\text{ cm}^{-1}$  is symmetric and is attributed to the fundamental bending mode of  $\text{H}_2\text{O}_m$  (Nakamoto 1978). The absorption bands for  $\text{CO}_2$  exist at  $\sim 2350$  and at a doublet peak at  $\sim 1515$  and  $\sim 1435\text{ cm}^{-1}$ . The  $2350\text{ cm}^{-1}$  band is related to the antisymmetric stretching mode of  $\text{CO}_{2m}$  (Newman et al. 1986, Stolper et al. 1989). The doublet is attributed to a doubly degenerate asymmetric stretching of  $\text{CO}_3^{2-}$ .



**Figure 4.1: Representative spectra showing the positions of the volatile absorption bands in the near-IR (A) and mid-IR (B).** The bands in the near-IR are weaker and hence require thicker samples in order to be amplified.

#### 4.2.1.2 Synchrotron radiation FTIR

Synchrotron radiation Fourier-transform infrared spectroscopy (SR-FTIR) produces a much brighter light than conventional global sources (von Aulock et al. 2014), resulting in a higher signal to noise ratio even when using small aperture sizes. A synchrotron source not only emits powerful light but also delivers the light at a smaller range of incidence angles, meaning less light is blocked by the aperture. These features allow significantly higher spatial resolution and spectra with higher signal-to-noise ratios, enabling H<sub>2</sub>O and CO<sub>2</sub> gradients to be measured in greater detail, their concentrations to be measured with greater accuracy, and their detection limits to be reduced.

#### 4.2.1.3 Focal Plane Array Detectors

Focal plane array detectors (FPA) consist of multiple channel detectors (Miller and Smith, 2005; von Aulock et al. 2014). All individual detector elements on the FPA simultaneously measure the absorbance of the spectral range of the system (Miller and Smith 2005). As a result, maps of absorbance over a large area can be created in a relatively short time. However, while FPA analyses are ideal for making large absorbance maps, the signal to noise ratios are lower than those from a synchrotron sourced FTIR due to the weaker light source, making smaller peaks harder to resolve. The spatial resolution of FPA is not only limited by the pixel size but also by signals off the focal plane as FPA are not confocal. Currently, FPA is not routinely used with a synchrotron source because of the difficulty in setting the FPA detector up on the synchrotron beamline and artefacts that appear in the image that are most apparent when tiling images and may be produced by sample specific scattering (or diffraction) and/or coherence speckle noise (von Aulock et al. 2014).

## 4.2.2 Sample preparation

For this project, we want to understand how the magmatic volatiles evolved and what that can tell us about how the magma responded to being intercepted by the IDDP-1 drilling project. To achieve this, 18 glass cuttings recovered from the IDDP-1 borehole at different times were selected for analysis and detailed mapping using synchrotron-sourced FTIR and the FPA at the infrared microspectroscopy beamline of the Australian Synchrotron (AS). Mapping consisted of measuring and displaying H<sub>2</sub>O and CO<sub>2</sub> distributions around a feature, such as bubbles in the glass.

The samples were chosen to best represent the evolution of the magma body. At each time interval a sample that exhibited the most common texture was selected (See chapter 2 section 2.5.1 on the textural details of the glass cuttings). At time intervals with a mixture of textures, multiple samples (up to 4 glass cuttings) with different textures were selected.

### 4.2.2.1 Target sample thickness

Measurements of both H<sub>2</sub>O and CO<sub>2</sub> were desired and therefore the samples needed to be prepared to a thickness that allowed the relevant vibrational bands to be detected. The aim was for each glass cutting to be analysed twice in the FTIR as thick and thin samples. Thick samples would allow H<sub>2</sub>O<sub>m</sub> and OH to be detected in the near-IR and CO<sub>2</sub> to be detected at the 2350 cm<sup>-1</sup> band. The thin samples would allow measurements of the H<sub>2</sub>O peaks in the mid-IR region, because H<sub>2</sub>O peaks in the mid-IR are stronger than the peaks in the near-IR. Larger peaks are more desirable as they are easier to measure. However, if the samples are too thick all the radiation at that wavenumber is absorbed and no radiation reaches the detector, so the peak will no longer be proportional to concentration and thickness (von Aulock et al. 2014). On the absorbance FTIR spectra this appears as a noisy peak that is cut off (Over saturated H<sub>2</sub>O<sub>m</sub> + OH;

Fig 4.1a). Thicker samples will also increase the likelihood of the beam hitting bubbles or crystal, which would interfere with the results.

To achieve the desired peak heights in the near-IR, concentrations of H<sub>2</sub>O and CO<sub>2</sub> measured in the IDDP-1 glass cutting by Zierenberg et al. (2013) were used to estimate target sample thicknesses by rearranging (Eq. 4.1):

$$d = \frac{M.A}{\rho.c.\epsilon} \times 100 \quad \text{Eq. 4.2}$$

Zierenberg et al. (2013) used molar absorption coefficient values of 1.86 l.mol<sup>-1</sup>.cm<sup>-1</sup> for H<sub>2</sub>O<sub>m</sub> and 1.5 l.mol<sup>-1</sup>.cm<sup>-1</sup> for OH (from Ihinger et al. 1994), and 1214 l.mol<sup>-1</sup>.cm<sup>-1</sup> for molecular CO<sub>2</sub> (from Behrens et al. 2004). A density value of 2350 kg.m<sup>-3</sup> for rhyolite (Stevenson et al. 1994) was used. H<sub>2</sub>O<sub>m</sub> and OH concentrations provided by Zierenberg et al. (2013) are 0.60 ± 0.11 wt%, and 1.08 ± 0.09 wt%, respectively; CO<sub>2</sub> concentration is 85.08 ± 15.07 ppm. The desired minimum absorbances of the 5200 cm<sup>-1</sup> and the 4500 cm<sup>-1</sup> bands are 0.05, and 0.1 for the CO<sub>2</sub> band at 2350 cm<sup>-1</sup>.

Using the above absorbance values, the minimum required wafer thickness for optimum results for the near-IR peaks and 2350 cm<sup>-1</sup> (CO<sub>2</sub>) peak was estimated to be between 200 to 250 µm. Table 4.1 summarises the values of each parameter from the Beer-Lambert law (Eq. 4.1) used in the thickness calculations.

**Table 4.1: Parameter values from Zierenberg et al. (2013), which was used to calculate the required sample thickness for the near-IR and the CO<sub>2</sub> band.**

Parameter	H <sub>2</sub> O <sub>mol</sub>	OH	CO <sub>2</sub>	Units	Reference
<i>C<sub>average</sub></i>	0.6	1.17	85.08	wt%, ppm	Zierenberg et al. 2013
<i>C<sub>minimum</sub></i>	0.49	1.08	70.01	wt%, ppm	Zierenberg et al. 2013
<i>C<sub>maximum</sub></i>	0.7	1.26	100.15	wt%, ppm	Zierenberg et al. 2013
$\rho$	2350	2350	2350	kg.m <sup>-3</sup>	Stevenson et al. 1994
<i>A</i>	>0.05	>0.05	>0.1	unitless	-
<i>M</i>	18.02	18.02	44.01	g/mol	-
$\epsilon$	1.86	1.5	1214	l.mol <sup>-1</sup> .cm <sup>-1</sup>	lhinger et al. 1994 (H <sub>2</sub> O); Behrens et al., 2004 (CO <sub>2</sub> )

For measurements in the mid-IR, the samples were thinned to prevent too much absorbance and cut-off of the peaks (Fig. 4.1b). A maximum target thickness for each sample was calculated after the samples had been analysed in the near-IR region. When the values for concentration and molar absorptivity coefficient are known, the thickness can be directly calculated for each sample. Therefore, unlike the near-IR thickness calculations, the sample-specific H<sub>2</sub>O<sub>m</sub> and OH concentrations measured by the preceding near-IR analyses can be used as input in the calculations of the thickness required for analysis of H<sub>2</sub>O peaks in the mid-IR region (more detail of concentration calculation in section 4.2.4).

Newman et al. (1986) show that the molar absorptivity coefficient for the 3500 cm<sup>-1</sup> band depends on H<sub>2</sub>O speciation:

$$\epsilon_{3500} = X_{OH}\epsilon_{3500_{OH}} + X_{H_2O_m}\epsilon_{3500_{H_2O_m}} \quad \text{Eq. 4.3}$$

where  $X_{OH}$  and  $X_{H_2O_m}$  are the mass fractions of water dissolved as OH and H<sub>2</sub>O<sub>m</sub> respectively. Newman et al. (1986) calculated that if all H<sub>2</sub>O is present as H<sub>2</sub>O<sub>m</sub> the molar absorptivity for the 3500 cm<sup>-1</sup> band is 56 l.mol<sup>-1</sup>.cm<sup>-1</sup> ( $\epsilon_{3500_{H_2O_m}}$ ), and if all H<sub>2</sub>O is present as OH it is 100 l.mol<sup>-1</sup>.cm<sup>-1</sup> ( $\epsilon_{3500_{OH}}$ ). The molar



absorptivity used for the 1630  $\text{cm}^{-1}$  band was 55  $\text{l.mol}^{-1}.\text{cm}^{-1}$  (Newman et al., 1986).

The density value used for the target thickness calculation was the same, 2350  $\text{kg.m}^{-3}$  (Stevenson et al. 1994).

Based on the near-IR  $\text{H}_2\text{O}$  species measurements, target thickness for the samples ranged from 40 to 60  $\mu\text{m}$  to produce the desired peaks. Table 4.2 summarises the parameters needed to calculate the maximum target thickness for the mid-IR analysis.

**Table 4.2: Parameter values required to calculate the thickness for the thin samples.**

Parameter	$\text{H}_2\text{O}_m + \text{OH}$ (3500 $\text{cm}^{-1}$ )	$\text{H}_2\text{O}_m$ (1630 $\text{cm}^{-1}$ )	Units	Reference
<b>C</b>	from near-IR for each sample	from near-IR for each sample	wt%	-
<b><math>\rho</math></b>	2350	2350	$\text{kg.m}^{-3}$	Stevenson et al. 1994
<b>A</b>	>0.1	>0.1	unitless	-
<b>M</b>	18.02	18.02	g/mol	-
<b><math>\epsilon</math></b>	speciation dependent (77.6 – 147.8) (Eq. 4.3)	25	$\text{l.mol}^{-1}.\text{cm}^{-1}$	Newman et al. 1986

#### 4.2.2.2 Doubly polished sample wafer preparation

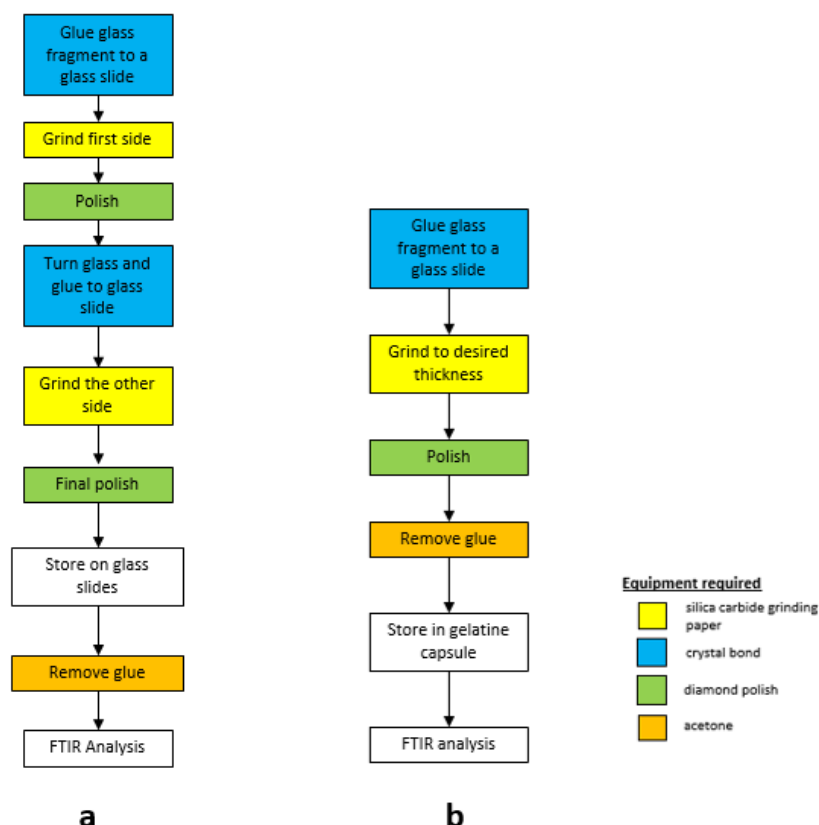
A major challenge in using FTIR is preparing samples (von Aulock et al. 2014), since thin, doubly polished unsupported wafers are needed. This can be difficult for very small or fragile samples, and hence limits the samples that can be analysed. Variations of sample thickness can cause major errors in the volatile calculation from FTIR measurements. The measured values are directly proportional to both

the thickness and volatile content of the glass. Therefore the accuracy of the thickness measurements is often the biggest limitation to the accuracy of quantitative measurements, especially in the case of very thin samples.

Preparation of wafers of thick samples was undertaken in the Department of Geological Sciences, University of Canterbury. Fig 4.2a illustrates the series of processes that were followed in preparing the samples. Preparation started by mounting a chip of the glass cuttings onto a glass slide using the adhesive Crystalbond™ 509. Crystalbond™ 509 has a relatively low melting temperature, ~121 °C, and is soluble in acetone, allowing it to be easily removed from the sample when desired. This enables easier handling of samples during preparation. Mounted samples were polished using silica carbide grinding paper with a grade of P1200 and then P2000 to achieve a flat surface, before final polishing using diamond paste of 3 µm followed by 1 µm on a polishing disc. After polishing, the sample was turned over and remounted in Crystalbond™ 509 so that the other side could be ground and polished in the same way to the target thickness. The two surfaces were prepared parallel to avoid thickness variations across the samples. Sample thickness was checked throughout the grinding and polishing of the second surface while the sample was mounted on the glass slide using a digital indicator micrometer. This ensured the target thickness was reached and ensured variations across the samples were minimised. Finally, the Crystalbond™ 509 was removed by placing the samples in an acetone bath and allowing the adhesive to dissolve. Once the Crystalbond™ 509 had dissolved, the sample was moved from the bath and left to dry, leaving a freestanding sample wafer for analysis. To reduce risk of damage during transport, all the samples prepared for near-IR and CO<sub>2</sub> analysis were transported mounted on the glass slide and removed just prior to analysis. Once analysed the freestanding wafer was stored in a gelatine capsule. Samples were handled using a paintbrush rather than tweezers to avoid breaking the samples.

Thin sample wafers for analysis of H<sub>2</sub>O peaks in the mid-IR region were prepared from the thick wafers by grinding down and re-polishing one of the surfaces to the target thickness. This was undertaken at the Australian Synchrotron after the

thick wafers had been analysed. The thickness of the sample was checked throughout the grinding and polishing using a digital indicator micrometer. This ensured that the target thickness was reached and ensured thickness variation across the sample was minimised. Fig 4.2b illustrates the series of processes to prepare the thin samples from the thick samples



**Figure 4.2: Schematic diagram showing the sample preparation process for a) the thicker samples and b) the thinner samples.**

### 4.2.3 FTIR data collection

For this project, the infrared microspectroscopy beamline at the Australian Synchrotron (AS) was used for four days. Absorbance maps for both thick ( $\sim 200 \mu\text{m}$ ) and thin ( $\sim 30\text{-}40 \mu\text{m}$ ) samples were created using a Bruker Hyperion 2000 infrared microscope attached to a Bruker Vertex 80v spectrometer bench on the AS infrared microspectroscopy beamline. Maps were collected over 128 scans (background every 5 measurements) at a spectral resolution of  $8 \text{ cm}^{-1}$  using a

synchrotron source, an extended potassium bromide (KBr) beamsplitter, and a liquid-nitrogen cooled mercury cadmium telluride (MCT) detector. A pinhole of 0.2  $\mu\text{m}$  were used to create a spatial resolution of 5.56  $\mu\text{m}$  but by overlapping the spectra a spatial resolution for each maps is  $\sim 3 \mu\text{m}$ . Samples were placed inside a purge box under 19-20% humidity. In total, 44 maps were created for 19 thick samples, and 8 maps were created for 8 thin samples. The size of the maps for the thicker samples ranges from 40 x 40  $\mu\text{m}$ , to 40 x 80  $\mu\text{m}$ . The maps for the thinner samples ranged from 35 x 35  $\mu\text{m}$ , to 50 x 50  $\mu\text{m}$ .

During the time at the AS, 13 additional maps were created for 8 samples from spectral images using the offline (i.e., non-synchrotron source) Bruker Hyperion 3000 infrared microscope with focal plane array detector (FPA) attached to a Bruker Vertex 70 spectrometer bench. The images were collected over 256 scans with a spectral resolution of 8  $\text{cm}^{-1}$ , using a conventional globar source, a KBr beamsplitter and the FPA detector (which had 64 by 64 channels covering 180 x 180  $\mu\text{m}$ , giving a spatial resolution of 2.8  $\mu\text{m}$ ). The size of the maps compiled from the 180 x 180  $\mu\text{m}$  FPA images range from 180 x 700  $\mu\text{m}$ , to 720 x 720  $\mu\text{m}$ .

#### 4.2.4 FTIR data processing

The peaks analysed for the near-IR region were the 5200  $\text{cm}^{-1}$  band for  $\text{H}_2\text{O}_{\text{mol}}$  and 4500  $\text{cm}^{-1}$  for OH. The peaks analysed for the mid-IR were the 3500  $\text{cm}^{-1}$  band for  $\text{H}_2\text{O}_{\text{t}}$ , 1630  $\text{cm}^{-1}$  peak for  $\text{H}_2\text{O}_{\text{mol}}$  and the 2350  $\text{cm}^{-1}$  peak for  $\text{CO}_2$ . In rhyolites total  $\text{CO}_2$  is present as  $\text{CO}_{2\text{m}}$  and  $\text{CO}_3^{2-}$  are not preserved (Chapter 3, section 3.2.2.2). The following describes how each parameter required in the Beer-Lambert law (Eq. 4.1) was established to calculate the concentrations of the Krafla glass cuttings.

#### 4.2.4.1 Absorbance

Bruker's OPUS (v7.2, 2012) spectroscopy software for was used to determine the height above the baseline of each relevant peak. The integration command on OPUS was used to calculate peak heights at a given peak with a linear baseline fit (method R in OPUS).

##### 4.2.4.1.1 CO<sub>2</sub> absorbance

As expected, a single CO<sub>2m</sub> peak at  $\sim 2350\text{ cm}^{-1}$  was observed in the thicker samples. However, in the thin samples there is a double peak at  $\sim 2350\text{ cm}^{-1}$  in spectra collected by SR-FTIR (Fig 4.3). These doublets are always positive, with the higher wavelength peak (peak 1) absorbance value higher than the lower wavelength peak (peak 2). The difference between the two peaks heights tends to change with sample thickness, where the greatest absorbance differences are in the thinnest samples. In all thick samples, the location of the CO<sub>2</sub> single peak is in the trough of the two peaks in the thin samples (Fig 4.3). These doublets were only observed from thin samples measured by the SR-FTIR and are not observed in the FPA-FTIR analysis. Watkins et al. (2012) identified such doublet peaks at this wavenumber to be caused by gaseous CO<sub>2</sub> (CO<sub>2gas</sub>). As bubbles were not in the beam path, it is presumed that CO<sub>2gas</sub> would be in the atmosphere in which the measurement was conducted. However, measurements were conducted in a purge box under 19-20% humidity to ensure measurements were taken in a clean, dry atmosphere.

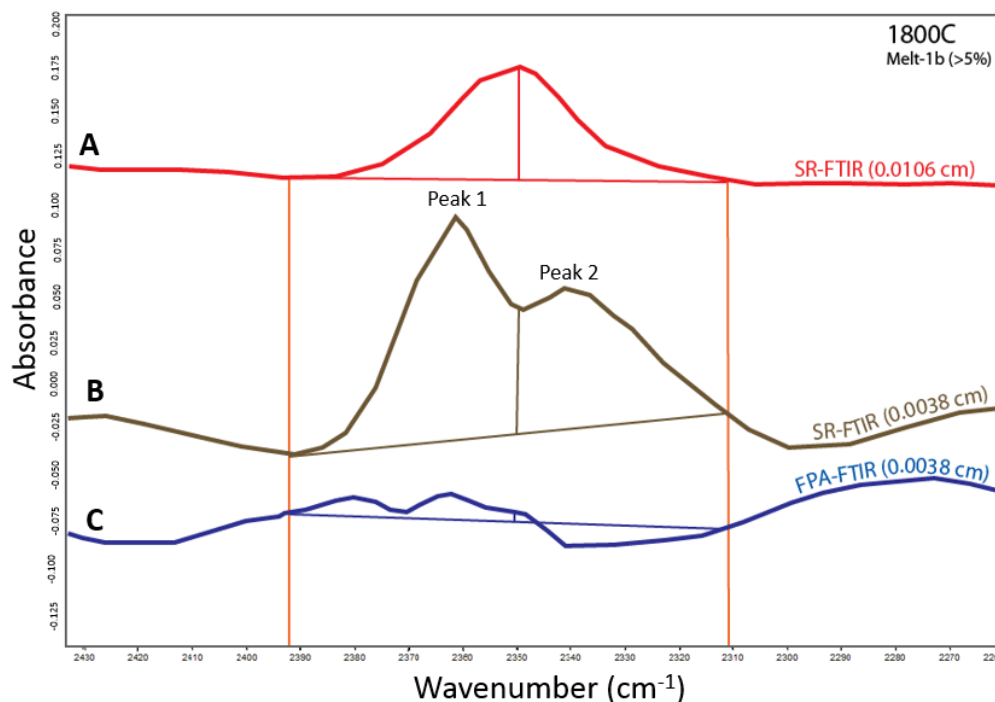
The following suggests that these doublet peaks are not caused from CO<sub>2gas</sub> in the atmosphere surrounding the measurements;

1. The doublet peaks only appear in the spectra for the thin samples; in the thick samples the peak is a single peak indicative of dissolved CO<sub>2m</sub> (Watkins et al. 2012). Measurements of the thick and thin samples were conducted under the same conditions. If the CO<sub>2gas</sub> were in the atmosphere surrounding the measurements it would be expected to be in both sets of measurements.

2. One possibility to explain the absence of the doublet in the thick samples is that the  $\text{CO}_{2\text{m}}$  vibrations overwhelm those from the  $\text{CO}_{2\text{gas}}$ . However, the doublet peaks in the thin samples all have higher absorbance than the single peaks in the thicker samples (Fig 4.3).
3. The doublet peaks are always positive indicating that  $\text{CO}_{2\text{gas}}$  has increased between background and sample measurements. Background measurements were taken every five sample measurements. If there was  $\text{CO}_{2\text{gas}}$  leaking into the purge box during the measurement series the levels, and thus the doublet peak, would be expected to increase gradually throughout the measurement sequence. However, during the five sample measurements the absorbance of the doublet does not change. This suggests that the  $\text{CO}_{2\text{gas}}$  always increases after every background and then remains stable for five sample measurements before increasing again, but only after the next background. This seems hard to explain, unless the actual movement to conduct a background is causing some sort of leak. However, once again, why does this only affect the analyses of thin samples and not the thick samples?

Based on these observations, it seems more likely that these doublet peaks relate to something specific about the thin sample rather than  $\text{CO}_{2\text{gas}}$  in the atmosphere around the measurements because bubbles in the samples were avoided during data collection.

The causes of the double peak at  $\sim 2350\text{ cm}^{-1}$  in the thin samples are currently unknown. For the purpose of this project, the concentrations of  $\text{CO}_2$  were taken from the peaks in the thick samples. This is because the peak is always positive and a single peak, indicative of  $\text{CO}_{2\text{m}}$  dissolved in rhyolite (Watkins et al., 2012).



**Figure 4.3: Variation of the CO<sub>2m</sub> peak (~2350 cm<sup>-1</sup>) between, A) the thick (106 μm) sample measured by the SR-FTIR (red), B) thin (39μm) sample measured by the SR-FTIR (brown) and C) a thin sample measured by the FPA-FTIR (blue). Spectra displaced along y-axis for clarity. All three peaks are from the same sample (1800C).**

#### 4.2.4.2 Sample thickness

Sample thickness is a major concern as it can introduce significant error to the volatile concentration calculations (von Aulock et al. 2014). Therefore, in addition to careful sample preparation, the sample thicknesses were determined at each spot using the following three methods:

A digital indicator, with a needle-tip contact point, mounted on a stand with a marble base was used to directly measure thickness. Placing the sample on the marble base and measuring the distance with and without the sample under the needle-tip and obtaining the difference was the simplest and quickest method to measure sample thickness. The digital indicator has an accuracy of 1 μm, however, it was difficult to be certain that the thickness was measured at precisely the same spot as the FTIR measurement was taken. Firstly, because the needle tip is larger than the analytical spot size, and secondly, as the needle was positioned by eye after the sample had been measured and transferred from the

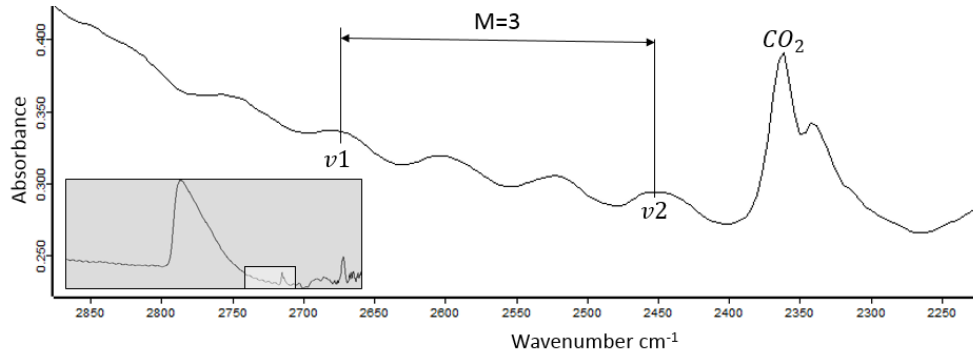
infrared microscope, it was difficult to be certain that the thickness was measured at exactly the same spot as the analysis. There was also the danger of the needle tip breaking the sample.

Interference fringes were used to measure sample thickness from the thinner samples (Nishikida et al. 1996; Tamic et al. 2001; Wysoczanski and Tani, 2006; Nichols and Wysoczanski, 2007). Interference fringe patterns were seen as a regularly undulating background in the FTIR spectra of the thinner samples even in transmitted light (Fig 4.4). As sample thickness decreases, the amplitude and wavelength of the interference fringes increases. Nishikida et al. (1996) have shown that, in thin samples, the wavelength of the interference fringe pattern is directly proportional to the thickness and the refractive index of the sample:

$$t = \frac{m}{2n(\nu_1 - \nu_2)} \quad \text{Eq. 4.4}$$

where  $t$  is the thickness of the analysed area (cm),  $m$  is the number of waves in the selected interval of wavenumbers (between  $\nu_1$  and  $\nu_2$ ), and  $n$  is the refractive index of the glass. For the Krafla glass a refractive index of 1.5 for rhyolite was used (Liu et al. 2005). Interference fringes in the region either side of the 2350  $\text{cm}^{-1}$  band between 2000 and 2700  $\text{cm}^{-1}$  were chosen, as this is a “clear” domain with no overlap with volatile peaks. Advantages of this method include a precision up to  $\pm 3 \mu\text{m}$  (Sun et al. 2007) and as the fringes appear on the FTIR spectra the thickness is calculated for the exact spot analysed. However, the limitation in estimating thickness by interference fringes is that it is only useful for samples that have been polished thin enough for the spectra to exhibit interference fringes (Nichols & Wysoczanski, 2007) and it requires knowledge of the refractive index of the sample.





**Figure 4.4: Representative spectra showing how the interference fringes are used to determine the thickness of the sample.**

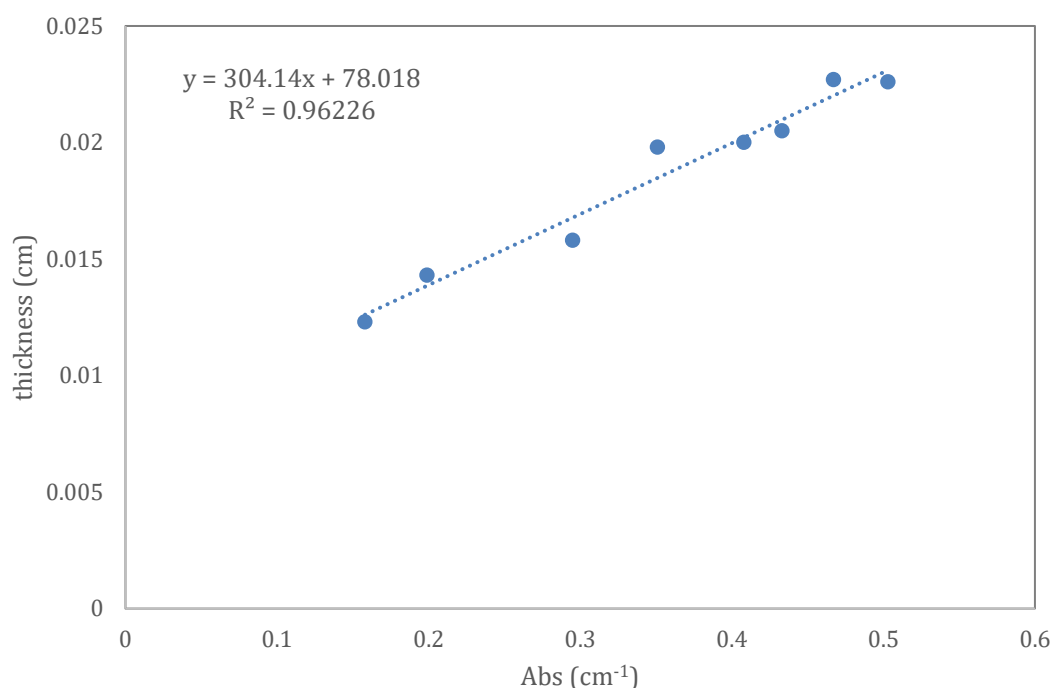
Another way to calculate thickness is by using the Si-O stretching on the 1830  $\text{cm}^{-1}$  band. Miwa and Toramaru (2013) have shown a clear relationship between the peak intensity of the 1830  $\text{cm}^{-1}$  band and the thickness of the samples (Fig 4.4). They also suggest that the volatile content in the glass does not appear to affect this relationship. Therefore, Miwa and Toramaru (2013) were able to calculate sample thickness using the following equation:

$$d = 0.034 \pm 0.0006 \text{Abs}_{1830} \quad \text{Eq. 4.5}$$

where  $d$  is the thickness of the sample (cm) and  $\text{Abs}_{1830}$  is the absorption at the 1830  $\text{cm}^{-1}$  band. However, this equation is for the specific composition in the Miwa and Toramaru (2013) study, and as the dependency of this relationship on composition has not yet been determined. It is thus not known if this constant applies to the composition of the IDDP-1 glass cuttings. In this project, the thicknesses of seven samples with a known analysis location were measured using the digital indicator and compared to the absorbance of the 1830  $\text{cm}^{-1}$  band measured at the same location. This allowed a constant of  $0.0304 \pm 0.0008$  to be calculated specifically for the Krafla samples (Fig 4.5) and hence the sample thickness at other locations was then estimated using the 1830  $\text{cm}^{-1}$  absorption band measurements.

This method can only be done on the thick samples because the 1830  $\text{cm}^{-1}$  does not appear in the thin samples. For this project, the thickness from the interference fringes and digital indicator measured thickness were not compared

due to the potential of breaking the fragile sample when using the digital indicator.



**Figure 4.5: The relationship between glass thickness and the absorbance at 1830 cm<sup>-1</sup> for the Krafla glass cuttings.**

A major limitation of thickness estimations by the 1830 cm<sup>-1</sup> peak is that the directly measured thickness used the digital indicator. As discussed above, the thickness will not be measured at precisely the same spot as the FTIR measurement was taken.

#### 4.2.4.3 Density

Due to the very small sample size, it was too difficult to directly measure the glass density. Sample-specific glass density was calculated using major element data obtained by electron microprobe (section 4.3) and H<sub>2</sub>O<sub>t</sub> from FTIR using the model for silicate melts at high temperatures from Lange & Carmichael (1987) and Lange (1997), extrapolated to room temperature. Partial molar volumes were taken from Lange & Carmichael (1987) and Lange (1997), except for H<sub>2</sub>O,

which was from Ochs and Lange (1997). For samples in which major elements were not analysed, a density of  $2410 \text{ kg.m}^{-3}$  was used. This was the average density for all the samples that were analysed by the electron microprobe analysis.

#### 4.2.4.4 Molar absorptivity coefficients

For the near-IR  $\text{H}_2\text{O}$  bands, the molar absorptivity coefficient used for the  $\text{H}_2\text{O}_{\text{mol}}$  band at  $5200 \text{ cm}^{-1}$  was  $1.61 \text{ l.mol}^{-1}.\text{cm}^{-1}$  and  $1.73 \text{ l.mol}^{-1}.\text{cm}^{-1}$  for the OH band at  $4500 \text{ cm}^{-1}$  (Newman et al. 1986).

The molar absorptivity coefficient used for the  $\text{CO}_2$  band at  $2350 \text{ cm}^{-1}$  in the thick samples was  $1214 \text{ l.mol}^{-1}.\text{cm}^{-1}$  (Behrens et al. 2004).

For the mid IR  $\text{H}_2\text{O}$  bands, the molar absorptivity coefficient used for the  $\text{H}_2\text{O}_{\text{mol}}$  band at  $1630 \text{ cm}^{-1}$  was  $55 \text{ l.mol}^{-1}.\text{cm}^{-1}$  (Newman et al. 1986). For the  $3500 \text{ cm}^{-1}$  band the molar absorptivity coefficient depends on speciation according to Newman et al. (1986). However, the equation provided to calculate the molar absorptivity coefficient requires advance knowledge of the  $\text{H}_2\text{O}$  speciation (Eq. 4.3). Recently, McIntosh et al. (2017) present a new methodology that accounts for the species-dependence of the molar absorptivity coefficient for the  $3500 \text{ cm}^{-1}$  band without requiring advance knowledge of the species proportions. This enables accurate water species concentrations to be determined from the  $3500$  and  $1630 \text{ cm}^{-1}$  absorbance bands.

##### *4.2.4.4.1 Species-dependent molar absorptivity coefficient for total $\text{H}_2\text{O}$ at $3500 \text{ cm}^{-1}$*

The equation that McIntosh et al. (2017) derive uses end-member molar absorptivity coefficient  $3500$  values (i.e. the theoretical  $\epsilon_{3500}$  value if all water within the glass were present exclusively as  $\text{H}_2\text{O}_{\text{mol}}$  or OH) to find accurate  $\text{H}_2\text{O}_{\text{total}}$  and OH concentrations from the  $3500 \text{ cm}^{-1}$  peak. It requires  $\text{H}_2\text{O}_{\text{mol}}$

concentration to be known (from 5200 or 1630  $\text{cm}^{-1}$  peak), but not the proportions of the different  $\text{H}_2\text{O}$  species.

The concentration of OH using the end-member  $\epsilon_{3500}$  values can be calculated by rearranging Equation 4.3, using the definitions  $X_{OH} = C_{OH}/C_{H_2O_t}$ , and  $X_{H_2O_m} = C_{H_2O_m}/C_{H_2O_t}$ :

$$\epsilon_{3500} = \frac{C_{OH}\epsilon_{3500_{OH}} + C_{H_2O_m}\epsilon_{3500_{H_2O_m}}}{C_{H_2O_t}} \quad \text{Eq. 4.6}$$

Rearranging for  $C_{H_2O_t}$ , and substituting into the Beer-Lambert law (Eq. 4.1), an expression for the concentration of OH is given;

$$C_{OH} = \frac{1}{\epsilon_{3500_{OH}}} \left( \frac{100M(A/l)_{3500}}{\rho} - \epsilon_{3500_{H_2O_m}} C_{H_2O_m} \right) \quad \text{Eq. 4.7}$$

This equation makes it possible to directly calculate the concentration of OH in a sample of known thickness, using the 3500  $\text{cm}^{-1}$  absorption peak and the  $\text{H}_2\text{O}_m$  concentration (which is found in the 1630  $\text{cm}^{-1}$  peak through Beer-Lambert law (Eq. 4.1), and the end-member  $\epsilon_{3500}$  values.  $\text{H}_2\text{O}_t$  is calculated simply by;

$$C_{H_2O_t} = C_{OH} + C_{H_2O_m} \quad \text{Eq. 4.8}$$

## 4.3 EPMA measurements on major elements

The same wafers of Krafla glass cuttings analysed by FTIR were analysed by electron probe microanalysis (EPMA) to determine their major element compositions. The EPMA analysis determine if there are any heterogeneities within and between samples. Major element compositions were also used to calculate the density of the glass cuttings for a more accurate magmatic volatile calculation and in other composition-dependent models used in the study.

### 4.3.1 Principle

EPMA is a technique for chemical analysis over a small, selected area of solid samples, in which X-rays are excited by a focused electron beam (Potts et al. 1995; Reed 2005). The resulting X-ray spectrum contains lines characteristic of the elements present. Determination of the geochemistry is achieved by measuring line intensities for each element in the sample and comparing their intensities with those emitted from standard samples (Potts et al. 1995; Reed 2005). Standard samples are pure elements or samples of known composition. EPMA can detect elements from atomic number 4 (Be) to 92 (U) (Potts et al. 1995).

### 4.3.2 Sample preparation

#### 4.3.2.1 Mounting

Glass wafers which had previously been analysed by the FTIR were mounted into 2.5 cm diameter epoxy mounts. Firstly, up to five wafers were placed on double-sided tape adhered to a clean glass plate. Epoxy resin (EPO-TEK® 301) was then poured over the sample wafers into the 2.5 cm cylindrical moulds and left to set overnight. Once the resin had set the mounts were extracted from the mould and the tape was removed, leaving a polished flat surface of the wafers exposed. The glass wafers were then re-polished using diamond paste of 1 µm, leaving the glass surface scratch free.

#### 4.3.2.2 Coating

Most geological samples are non-conductive and require a conductive coating to prevent charging under electron bombardment. Common coating elements are carbon and gold. Carbon coating is the preferred coat for X-ray analysis because it has a minimal effect on the X-ray spectrum (Reed, 2005). As the main aim for

this project was X-ray analysis, the Krafla samples were therefore carbon coated. The sample mounts were placed in a vacuum chamber with a carbon evaporation source consisting of pointed carbon rods in contact. A current of about 100 A was passed through the carbon rods for a few seconds to cause carbon evaporation from the contacts. The evaporated carbon atoms travel in straight lines and coat the mounted samples. The Krafla glass cuttings were carbon coated at the School of Geography, Environment and Earth Sciences, Victoria University of Wellington.

### 4.3.3 EPMA data collection

Major element compositions of the Krafla glass cuttings were determined using the JEOL JXA-8230 Superprobe at the School of Geography, Environment and Earth Sciences, Victoria University of Wellington. Major elements were determined using an accelerating voltage of 15 kV, current of 8.0 nA, peak/background count times of 30s/15s, and a beam defocused to 10  $\mu\text{m}$ . In order to minimize sodium volatilization during analysis, it was measured for shorter times (10s/5s), measured first and at a fixed peak position in order to eliminate the need for a peak search procedure. Major element analyses were calibrated using natural and synthetic compounds (Jaroewich et al. 1980) as follows: basaltic glass standard VGA-99 for Ca, Mg, Fe; rhyolitic glass standard VG568 for Si, Al, K; and synthetic oxides for Ti, Mn, Cr.

A total of 16 samples were analysed by EPMA. In each sample five points on a clean part of the sample were analysed. To check for any analytical drift during analytical runs, standard sample VGA-99 and VG568 were used as standards (Appendix 1). Each standard was analysed twice after every 25 measurements on the Krafla samples.

### 4.3.4 EPMA data use

Analyses that generated totals less than 97 wt.% or greater than 103 wt.% were discarded. Measurements from the EPMA were used to compare the glass chemistry of the clear glass and the brown glass and determine if there are any heterogeneities within the samples. Major element compositions were also used to calculate glass density (see Section 4.2.4.3) and in various compositionally dependent models used in the study (Chapter 6).

## 4.4 Summary

For this project FTIR and EPMA analytical techniques were used on the IDDP-1 glass cutting to help determine the magmatic degassing history of the Krafla magma body in response to a sudden loss of pressure as a result of penetration by drilling. FTIR was used to measure the concentration of  $\text{H}_2\text{O}_m$ , OH,  $\text{H}_2\text{O}_t$  and  $\text{CO}_2$  in the glass cuttings. The high resolution maps created using the SR-FTIR, allowed  $\text{H}_2\text{O}$  and  $\text{CO}_2$  species distributions to be examined around bubbles in the glass cuttings. The high resolution maps then can help determine the bubble behaviour. Data from the SR-FTIR and the FPA-FTIR are used to calculate the average magmatic volatile concentration of each sample. EPMA is used to determine the major elements compositions of the glass cuttings. In this study these were used to compare the compositions of different glass textures, calculate density to provide a more accurate calculation of the magmatic  $\text{H}_2\text{O}$  and  $\text{CO}_2$  contents, and as inputs into various compositionally dependent models used in the study (Chapter 6).

# Chapter 5: Glass composition and volatile distribution

## 5.1 Introduction

This chapter presents data from the major element and magmatic volatile (H<sub>2</sub>O and CO<sub>2</sub>) analyses of the Krafla glass cuttings. Results from these analyses will form the basis for the interpretations presented in Chapter 6.

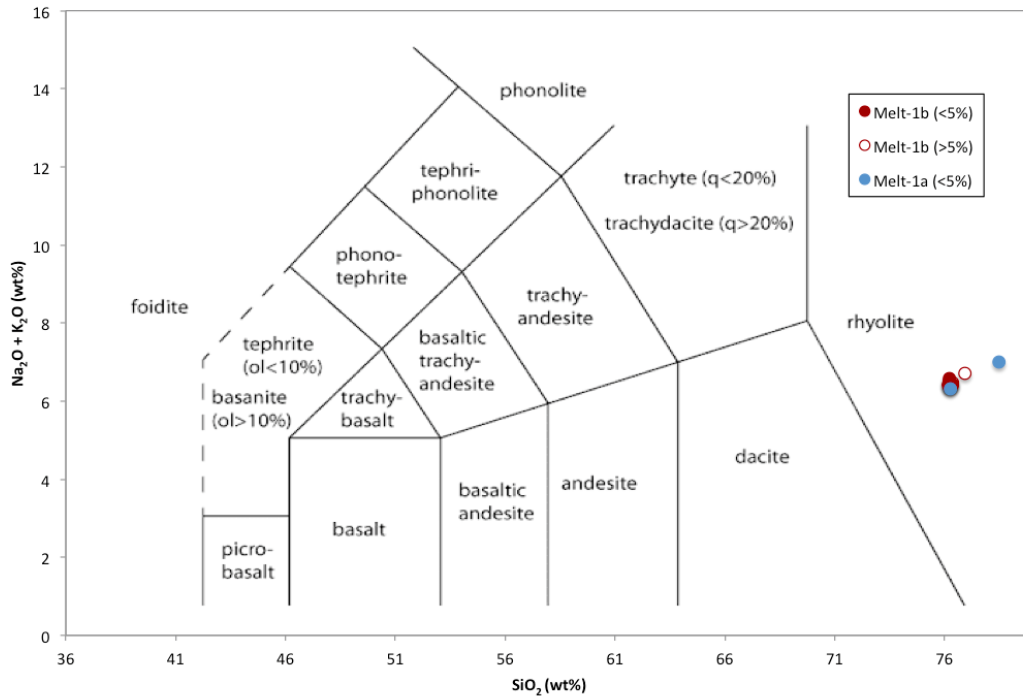
## 5.2 Major element glass chemistry

Table 5.1 provides the average major element compositions of Melt-1a and Melt-1b measured by EPMA. Melt-1a and Melt-1b have broadly similar compositions, but Melt-1a contains higher SiO<sub>2</sub> and K<sub>2</sub>O concentrations, while Melt-1b has higher TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, MgO and CaO. Both Melt-1a and Melt-1b plot within the rhyolite field on the total alkali silica (TAS) classification diagram (Fig 5.1).

**Table 5.1: Geochemistry of the glass cuttings obtained from EPMA.** Major oxides normalised to 100 wt% on an anhydrous basis. n is the number of samples analysed

	<b>Melt-1a (n=2)</b>	<b>STDEV</b>	<b>Melt-1b (n=12)</b>	<b>STDEV</b>
<b>SiO<sub>2</sub></b>	77.37	0.34	76.31	0.38
<b>TiO<sub>2</sub></b>	0.29	0.01	0.33	0.02
<b>Al<sub>2</sub>O<sub>3</sub></b>	11.78	0.08	12.12	0.11
<b>FeO</b>	2.50	0.10	2.97	0.11
<b>MnO</b>	0.06	0.01	0.07	0.02
<b>MgO</b>	0.16	0.02	0.22	0.01
<b>CaO</b>	1.19	0.02	1.50	0.05
<b>Na<sub>2</sub>O</b>	3.28	0.08	3.48	0.09
<b>K<sub>2</sub>O</b>	3.37	0.09	2.98	0.05





**Figure 5.1: TAS diagram with EPMA data of the Melt-1a and Melt-1b.**

## 5.3 Temporal variations of H<sub>2</sub>O speciation concentrations in the Krafla melt

### 5.3.1 H<sub>2</sub>O<sub>t</sub> concentration

Fig 5.2A shows the temporal variation of total H<sub>2</sub>O<sub>t</sub> of the Krafla glass cuttings. Overall the H<sub>2</sub>O<sub>t</sub> concentrations of Melt-1a are lower than those of Melt-1b. The average H<sub>2</sub>O<sub>t</sub> concentration of Melt-1a (n=3) is  $1.46 \pm 0.07$  wt% and  $1.8 \pm 0.07$  wt% for Melt-1b (n=13). No significant changes are observed in the concentration of H<sub>2</sub>O<sub>t</sub> over time for both Melt-1a and Melt-1b.

### 5.3.2 OH concentration

Fig 5.2B shows the concentration of OH over the time of glass recovery. The average OH concentration for Melt-1a (n=3) is  $1.00 \pm 0.06$  wt% and  $1.12 \pm 0.2$  wt% for Melt-1b (n=13). Over the time sampled the OH concentration of Melt -

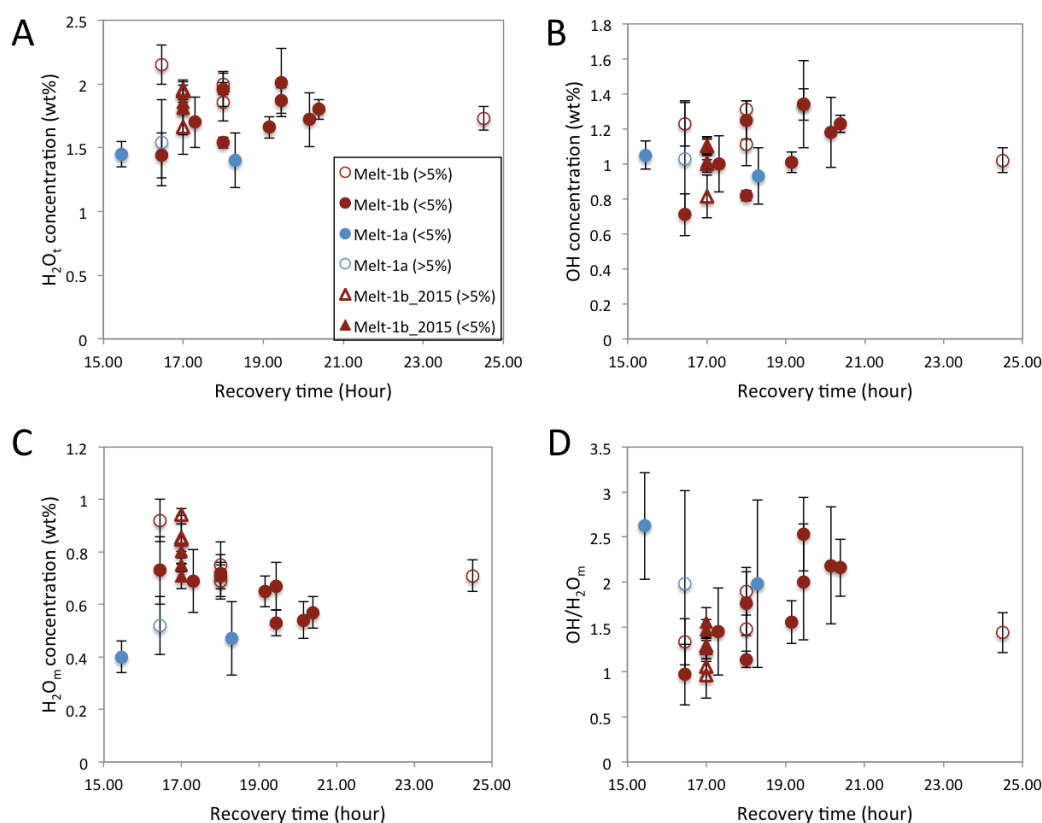
1a decreased from  $1.05 \pm 0.08$  to  $0.93 \pm 0.16$  wt%. At 16:45 Melt-1b had a large range of OH concentrations ( $0.71 \pm 0.12$  to  $1.23 \pm 0.13$  wt%). Apart from the range seen in 16:45, the OH concentrations in Melt-1b from 16:45 until 20:39 increased from  $0.71 \pm 0.12$  to  $1.23 \pm 0.11$  wt%. At 00:50 the OH concentration of Melt-1b was  $1.02 \pm 0.13$  wt%.

### 5.3.3 H<sub>2</sub>O<sub>m</sub> concentration

Fig 5.2C shows the concentration of H<sub>2</sub>O<sub>m</sub> over the time of glass recovery. The concentration of H<sub>2</sub>O<sub>m</sub> in Melt-1a is lower than Melt-1b. The average concentration of Melt-1a (n=3) is  $0.46 \pm 0.06$  wt%. The average H<sub>2</sub>O<sub>m</sub> concentration for Melt-1b (n=13) is  $0.68 \pm 0.1$  wt%. At 16:45 and 17:00 there was a large range of H<sub>2</sub>O<sub>m</sub> concentration in Melt-1b. From 17:30 to 20:39 the concentration of H<sub>2</sub>O<sub>m</sub> decreased from  $0.69 \pm 0.12$  to  $0.57 \pm 0.05$  wt%. At 00:50 the concentration of H<sub>2</sub>O<sub>m</sub> was  $0.71 \pm 0.06$  wt%.

### 5.3.4 OH/H<sub>2</sub>O<sub>m</sub>

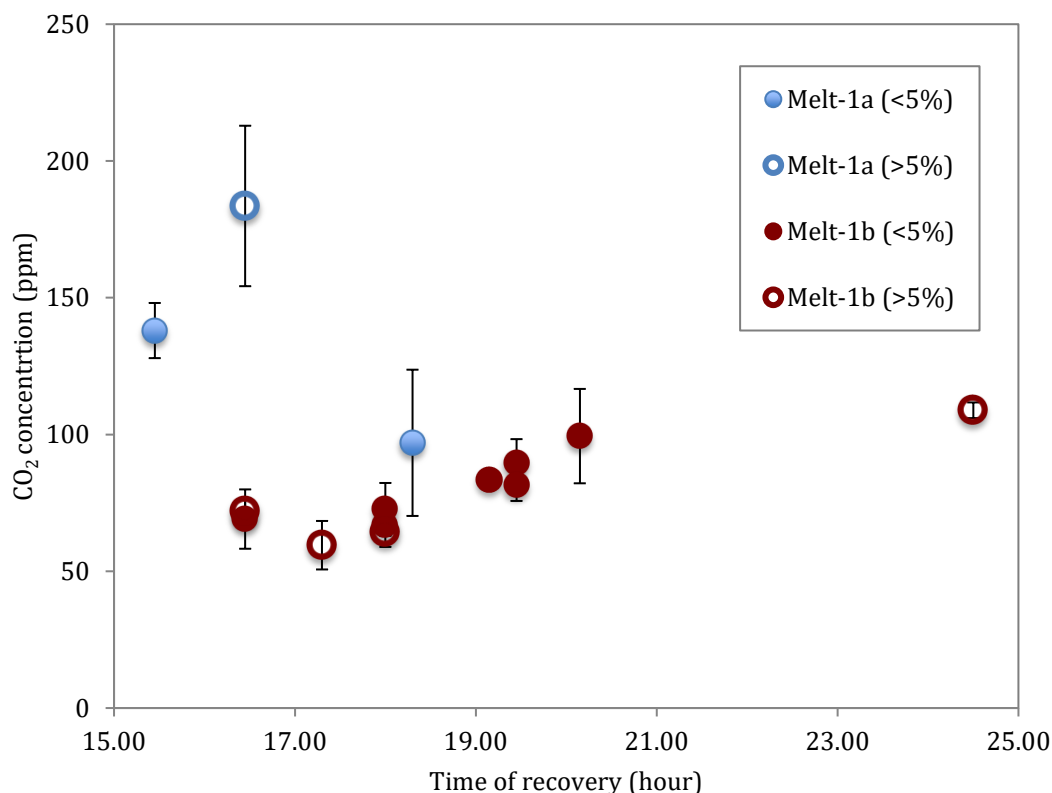
All sample have higher OH concentration than H<sub>2</sub>O<sub>m</sub>. Fig 5.2D shows the OH/H<sub>2</sub>O<sub>m</sub> ratio over the time of glass recovery. OH/H<sub>2</sub>O<sub>m</sub> increased from 0.97 to 2.53 over time for Melt-1b (1645 to 2039), although the large propagated errors mean most values are within error of each other. The glass cutting from 0050 does not follow this trend and has an OH/H<sub>2</sub>O<sub>m</sub> value of  $1.44 \pm 0.22$ . No apparent trend is observed with Melt-1a, though this may reflect the limited data points. The average OH/H<sub>2</sub>O<sub>m</sub> of Melt-1a is  $2.19 \pm 0.37$  and  $1.68 \pm 0.45$  for Melt 1-b.



**Figure 5.2: Volatile concentrations of the Krafla melt over recovery time.** Volatile concentrations were measured for  $H_2O_t$  (A), OH (B),  $H_2O_m$  (C), and OH/ $H_2O_m$  (D). Results from this study (circles) include samples from Melt-1a (blue) and Melt-1b (brown). Other sample results (triangles) were sourced from the addition glass cuttings (Chapter 2, section 2.7.1).

## 5.4 Temporal variations of $CO_2$ concentration in the Krafla melt

Fig 5.3 shows the temporal variation of the  $CO_2$  concentration of the Krafla melt. The concentration of  $CO_2$  in Melt-1a is generally higher than the concentration of  $CO_2$  in Melt-1b. The average  $CO_2$  concentration in Melt-1a ( $n=3$ ) is  $139.5 \pm 43$  ppm, and  $78.86 \pm 15.4$  ppm for Melt-1b ( $n=10$ ). At 16:45 the concentration of  $CO_2$  in Melt-1b was  $72.34 \pm 0.4$  then the  $CO_2$  decreased to  $59.53 \pm 8.9$  at 17:30. After 17:30, the  $CO_2$  concentration in Melt-1b increased from 63 to 115 ppm (Fig. 5.1D). No trend can be observed for Melt-1a, which may again reflect the limited data for this group. No systematic trend is observed in the vesiculation of the glass cuttings with time.



**Figure 5.3: CO<sub>2</sub> concentrations in the Krafla melt over the time of recovery.**

## 5.5 H<sub>2</sub>O and CO<sub>2</sub> distribution around bubbles

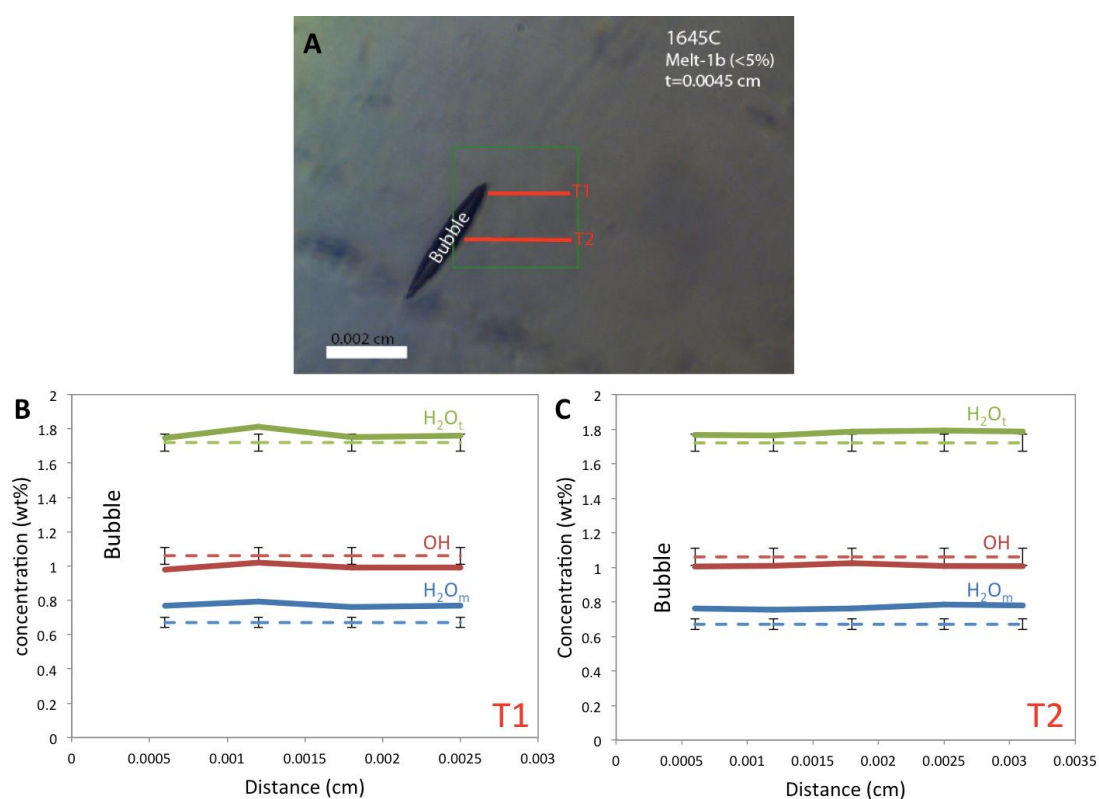
A total of six samples have been mapped for H<sub>2</sub>O speciation transects and three samples for CO<sub>2</sub> transects (Appendix 2). These transects have been categorised into three groups;

1. No changes in H<sub>2</sub>O and CO<sub>2</sub> concentration towards bubbles (Fig 5.4). This group consists of bubbles in samples 1645C, 1645D and 1945A.
2. H<sub>2</sub>O<sub>t</sub> and OH concentration decrease towards bubbles and the H<sub>2</sub>O<sub>m</sub> increases (Fig 5.5). CO<sub>2</sub> does not show systematic trends towards bubbles and may increase, decrease or remain constant. These transects are shown in samples 1545 and 1800C.
3. All H<sub>2</sub>O species and CO<sub>2</sub> increase towards the bubbles (Fig 5.6). This is shown in samples 1800A, 1915 and 1945A.

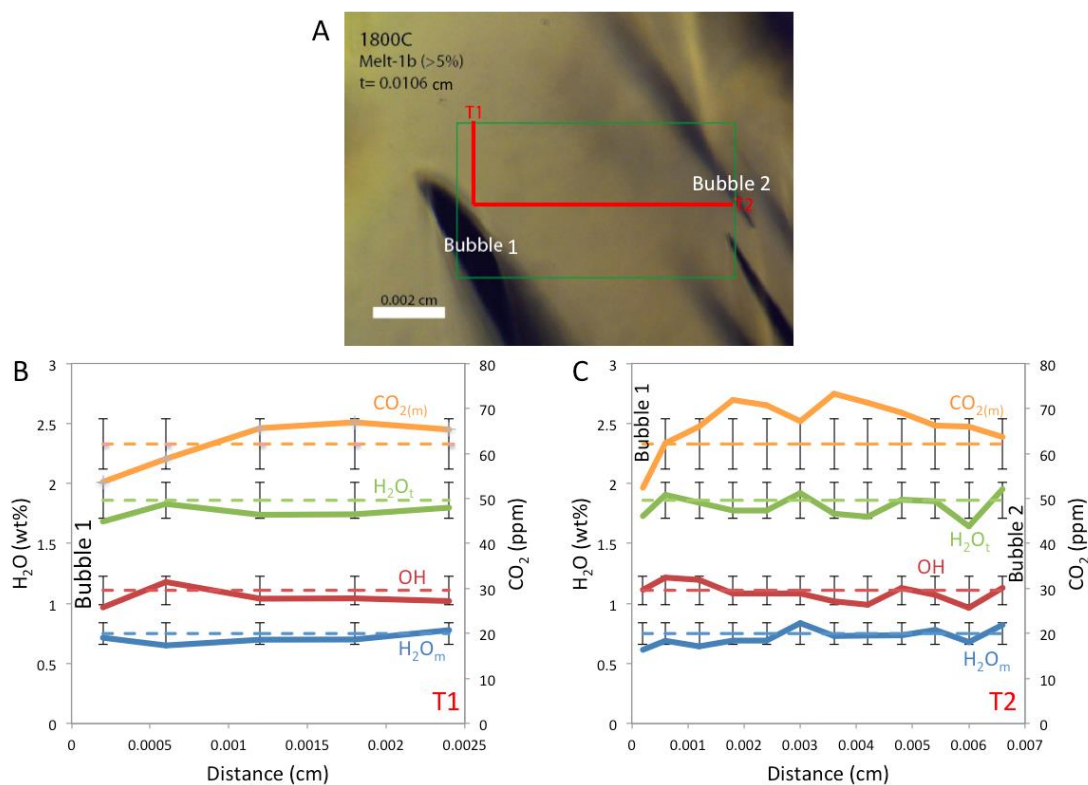
No correlation between time of recovery, glass colour, vesicularity or thickness of the samples has been observed in any of the three groups.

Sample 1800C displays behaviours from group 2 and group 3 (Fig 5.5) for different bubbles in the sample. Indicating that these different bubble behaviours can occur in the same sample.

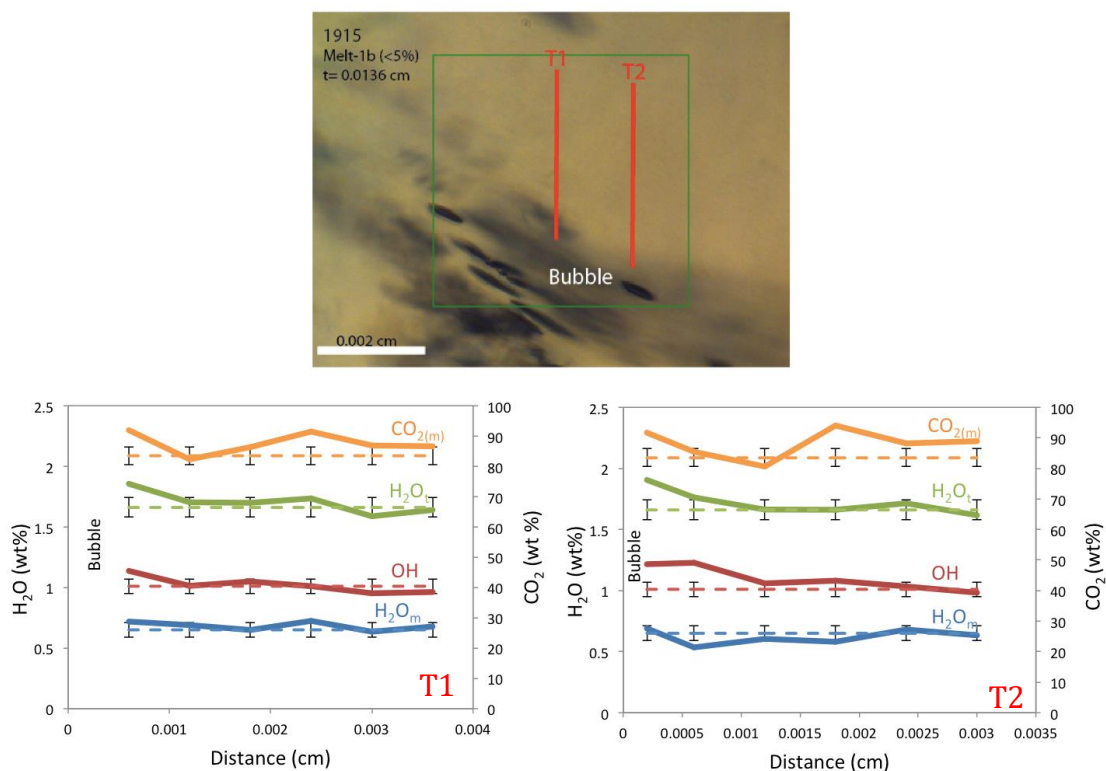
Sample 1945A (Fig 5.7) have displayed distribution pattern of group 1 and group 3 in the same bubble.



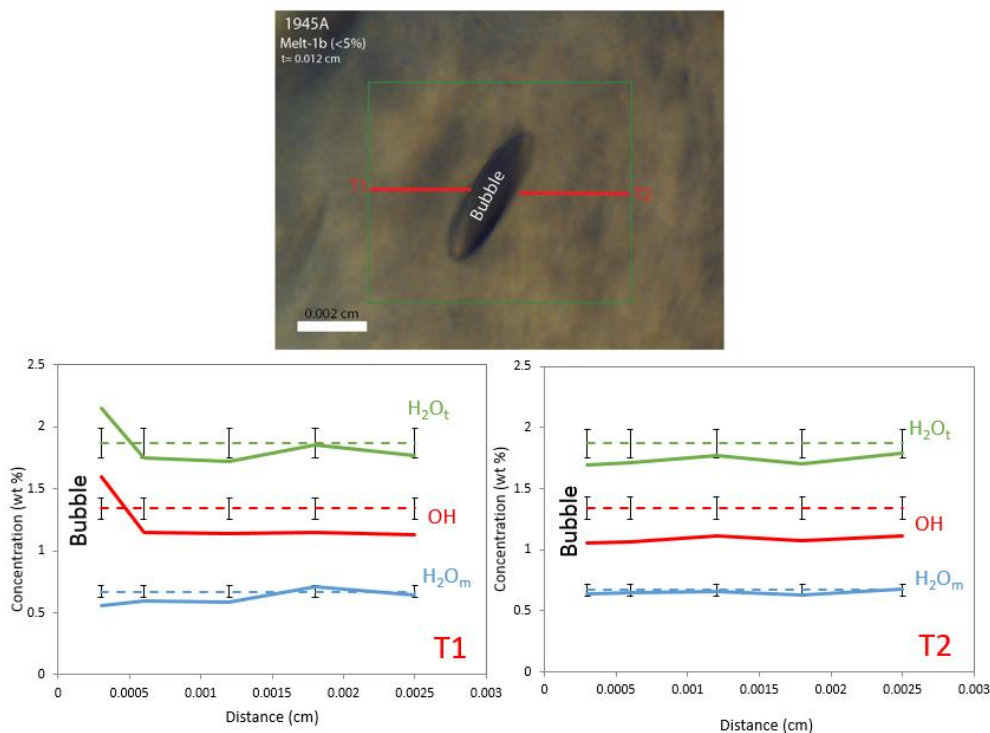
**Figure 5.4: Example of H<sub>2</sub>O transects showing no changes towards a bubble. Sample was too thin to measure CO<sub>2</sub>.**



**Figure 5.5: Example of samples where  $\text{H}_2\text{O}_t$ , OH and  $\text{CO}_2$  decrease and  $\text{H}_2\text{O}_m$  increases towards bubble 1.  $\text{H}_2\text{O}_t$ , OH and  $\text{H}_2\text{O}_m$  increase towards Bubble 2.**



**Figure 5.6: Example of bubbles where  $\text{H}_2\text{O}_t$ ,  $\text{H}_2\text{O}_m$ , OH and  $\text{CO}_2$  increase towards bubbles.**



**Figure 5.7: Example of multiple volatile distribution around a single bubble.** Volatile distribution in T1 shows H<sub>2</sub>O<sub>t</sub> and OH increases and H<sub>2</sub>O<sub>m</sub> slightly decreases towards the bubble. Volatile distributions in T2 are constant towards the bubble. Sample is too thin to measure CO<sub>2</sub>.

## 5.6 Summary

Major elements have shown that Melt-1a and Melt-1b are both rhyolite melts however, the two glass types have slightly different glass chemistry. Melt-1a is slightly higher in SiO<sub>2</sub> and K<sub>2</sub>O, while Melt-1b has higher concentrations of TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, MgO and CaO. The concentration of H<sub>2</sub>O<sub>t</sub> in the Krafla melt showed no change over time, however, OH/H<sub>2</sub>O<sub>m</sub> and the concentration of CO<sub>2</sub> increased over time. No correlations between volatile content and vesiculation of the glass cuttings have been observed. Based on seven samples, three volatile distribution around bubbles were identified. 1) The first group showed no changes in volatile distribution towards bubbles, 2) the second groups shows a decrease in H<sub>2</sub>O<sub>t</sub> and OH, and a increase of H<sub>2</sub>O<sub>m</sub> towards bubbles and 3) magmatic volatiles increases towards the bubbles. These transects showed no systematic trend with time or texture. Based on theses analyses, the next chapter will present interpretations of the response of the magma to being intercepted by the drilling.

# Chapter 6: Implications from the magmatic volatile evolution

## 6.1 Introduction

This project set out to examine the magmatic volatile ( $\text{H}_2\text{O}$  and  $\text{CO}_2$ ) evolution of the intersected magma from the IDDP-1. The results of the FTIR analyses have answered the two objectives below.

1. Determine any changes in magmatic volatile concentration through time.
2. Map the distribution of  $\text{H}_2\text{O}$  species and  $\text{CO}_2$  around bubbles in the samples from the time-series set of samples collected during drilling into the Krafla magma body.

These results, together with other published datasets, will allow the remaining objectives to be addressed:

3. Determine the magmatic degassing history of the Krafla magma body in response to a sudden loss of pressure as a result of intersection by the drilling.
4. Use the degassing history to assess quenches timing and changes in magmatic viscosity.
5. Investigate why the Krafla magma did not erupt while it was being drilled.

This chapter uses the results from the EPMA and FTIR analyses from the previous chapter in order to establish the magmatic volatile evolution of the Krafla melt after the unexpected depressurisation from the IDDP-1 drill. The findings from this project will also be compared to other glass chemistry and magmatic volatile studies from Zierenberg et al. (2013) and data from Saubin et al. (2017). Finally, possible future directions for studies on the magmatic volatiles in the intercepted Krafla melt will be discussed.

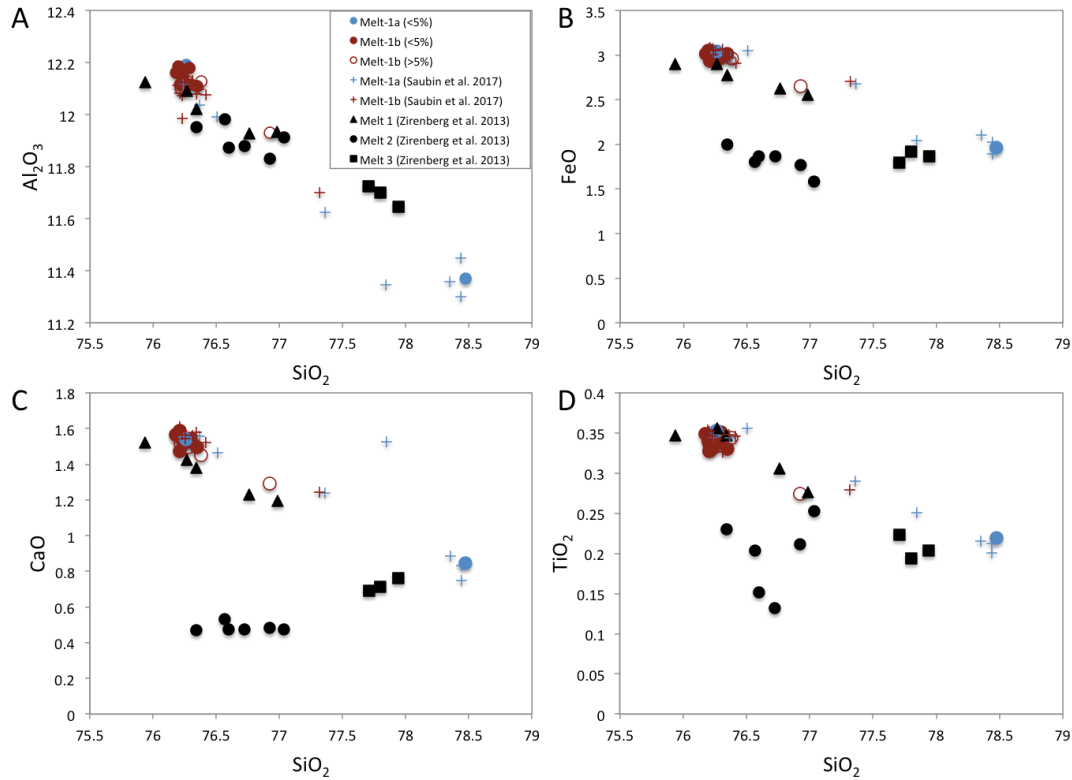


## 6.2. Pre-drilling magmatic conditions

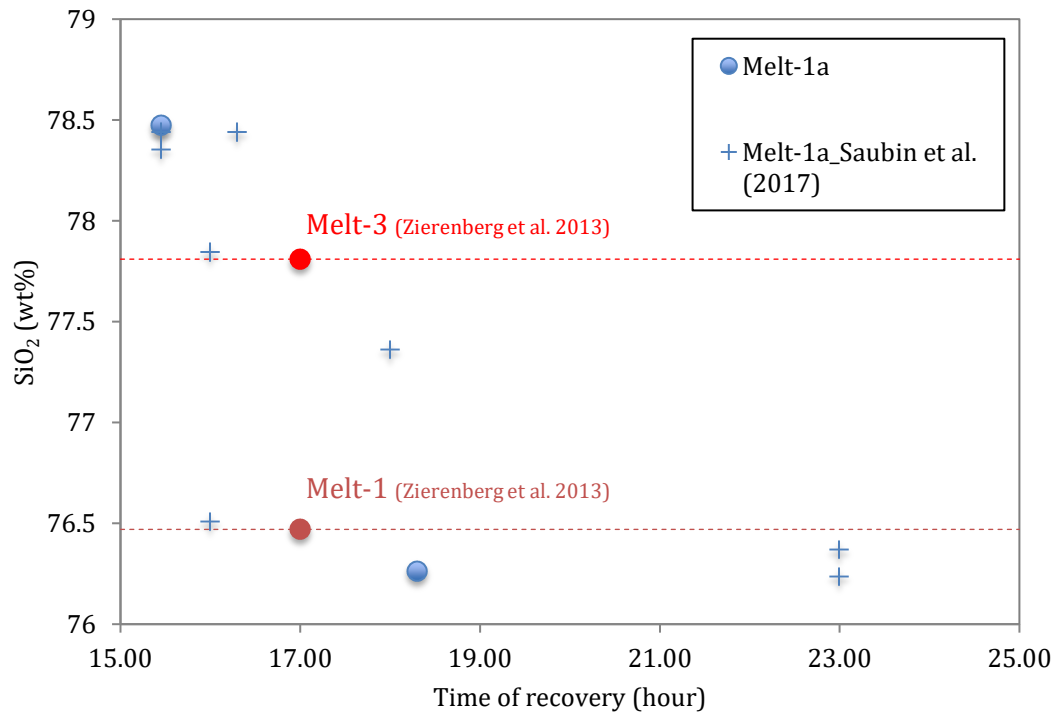
### 6.2.1 Melt chemistry

Glass chemistry of the glass cuttings from this project and Saubin et al. (2017) have shown that the texturally defined Melt-1a (clear glass) has a range of chemistries clustering into two distinctive end members. The two end-member glass chemistries are similar to the reported glass chemistries of Melt 1 and Melt 3 from Zierenberg et al. (2013), (Fig 6.1). Earlier recovered Melt-1a tends to have a similar composition to Zierenberg et al.'s (2013) Melt-3, however Melt-1a samples have slightly lower  $\text{Al}_2\text{O}_3$  than Zierenberg et al. (2013) Melt-3. This could reflect the difference of the standardisation procedure of the two different EPMA machines during data collection rather than reflecting a different melt. The later recovered Melt-1a has a similar composition to Zierenberg et al.'s (2013) Melt-1 (Figure 6.2). Two of the Melt 1a samples plot in a linear trend between the Melt-1 and Melt-3 chemistries. There are no systematic differences in vesicularity and microlite contents between the glass cuttings (Saubin et al. 2017). Therefore, the linear variation of the chemistry between the end-member compositions suggests the Melt 1a cuttings represent mixing between Melt-1 and Melt-3 (Fig 6.3). This interpretation is consistent with hypothesis invoked by Zierenberg et al. (2013).

The  $\text{SiO}_2$  contents of Melt-1b reported here and in Saubin et al. (2017), both measured in the same electron microprobe laboratory, are slightly lower than Zierenberg et al. (2013), where the  $\text{SiO}_2$  contents of earlier recovered Melt-1a are slightly higher than Melt-3. These differences are very small but are not likely to be a result from the interlaboratory bias because it would be expected that all the samples from this project and Saubin et al. (2017) would be systematically higher or lower than Zierenberg et al. (2013) contents. The samples from this project and Saubin et al. (2017) could reflect glass cuttings closer to two end-members.

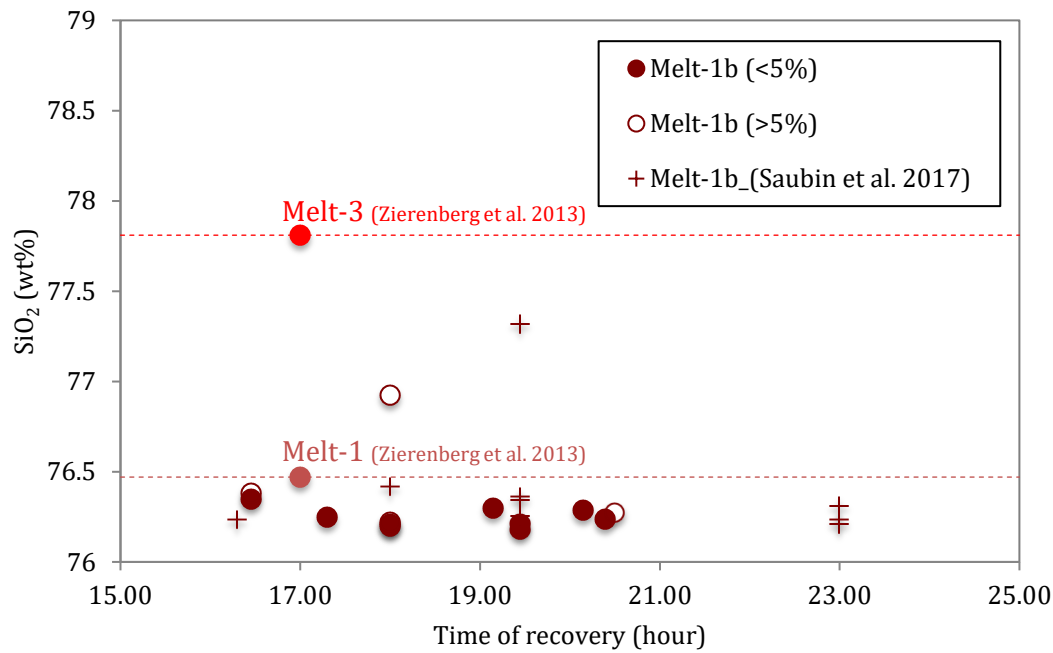


**Figure 6.1: Chemical variation diagrams for the different glasses from data obtained at the EPMA, additional data from Saubin et al. (2017) and Zierenberg et al. (2013). Data are normalised to 100% on an anhydrous basis. A)  $\text{SiO}_2$  versus  $\text{Al}_2\text{O}_3$ , B)  $\text{SiO}_2$  versus  $\text{FeO}$ , C)  $\text{SiO}_2$  versus  $\text{CaO}$  and D)  $\text{SiO}_2$  versus  $\text{TiO}_2$ .**



**Figure 6.2:  $\text{SiO}_2$  variation of Melt-1a (clear glass) over time of recovery.**





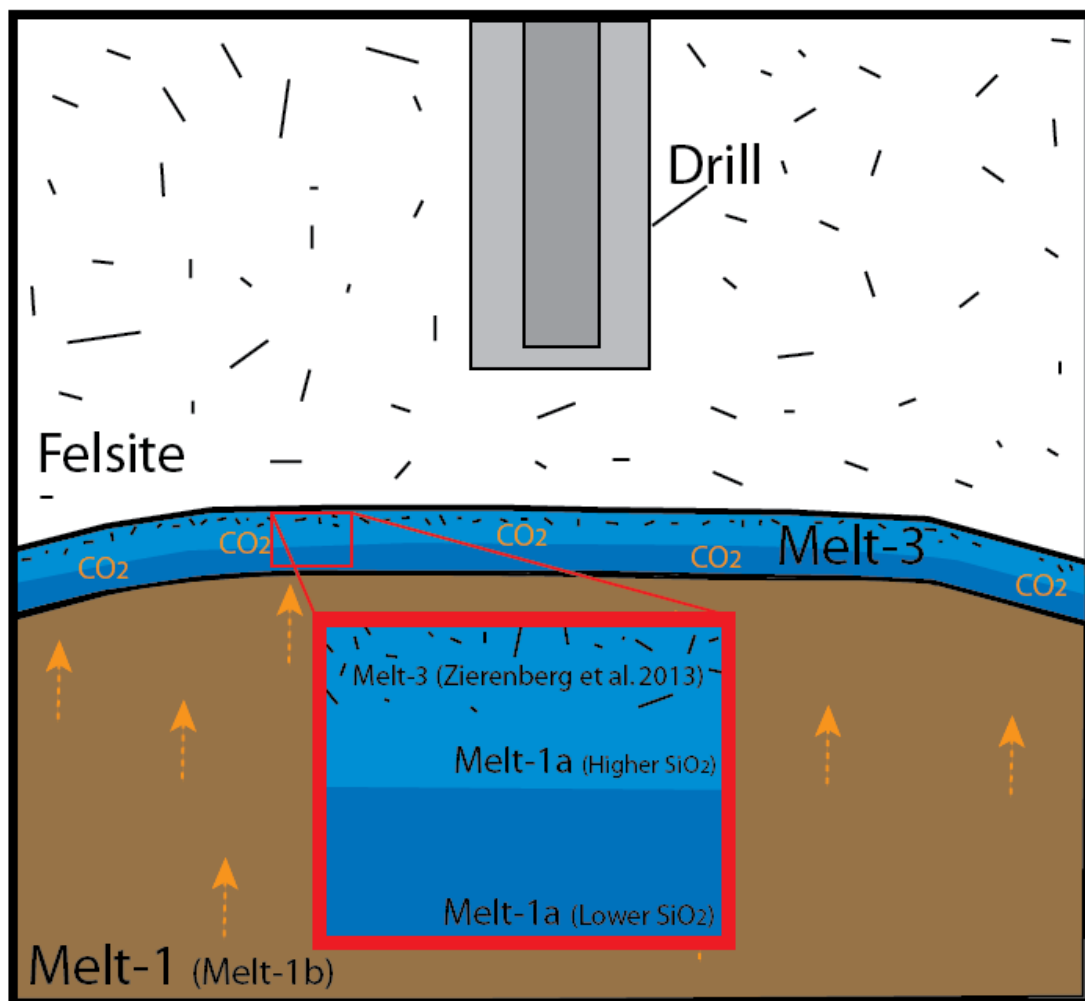
**Figure 6.4: Variation of SiO<sub>2</sub> of Melt-1b (brown glass) over time of recovery.**

## 6.2.2 Pre-drilling conditions

Fig 6.5 illustrates the conditions of the melt prior to drilling. The major element analyses have shown the distinct chemical signature of the magma body Melt-1 and Melt-3 and the surrounding felsite, which appears to have interacted with the magmatic body close to the contact. Zierenberg et al. (2013) has suggested that the crystal rich Melt-3 is a result of assimilation of the magma body (Melt-1) with the surrounding felsite. Based on the major element analysis of the crystal-poor clear glass cuttings, Melt-1a has glass chemistries that span the compositions of both Melt-1 and Melt-3. The earlier quenched clear glass have similar compositions to Melt-3, and the later quenched clear glass cuttings are similar in composition to the magma body based on Zierenberg et al. (2013) hypothesis. This supports the hypothesis of Zierenberg et al. (2013) that Melt 3 formed the boundary of the magma body, with more pristine magma being encountered as the drilling progressed.

The CO<sub>2</sub> concentrations in Melt-1a are higher ( $139.5 \pm 43$  ppm) than the CO<sub>2</sub> concentration in Melt-1b (Chapter 5, Section 5.4). The high concentration of CO<sub>2</sub> could be caused by CO<sub>2</sub> fluxing from below (Rust et al. 2004). Advancing the interpretation of

Zierenberg et al. (2013), Melt 1 partially melted the felsite to create Melts 3 and 1a. The partially melted felsite melt (Melt-1a) would be an initially anhydrous melt undersaturated in  $\text{CO}_2$  and  $\text{H}_2\text{O}$ . This allows Melt-1a to dissolve  $\text{CO}_2$  and  $\text{H}_2\text{O}$  from the deep volatiles fluxing from underlying magma (Rust et al. 2004) to create a distinct  $\text{CO}_2$  to  $\text{H}_2\text{O}$  ratio reflecting a different volatile saturation history.



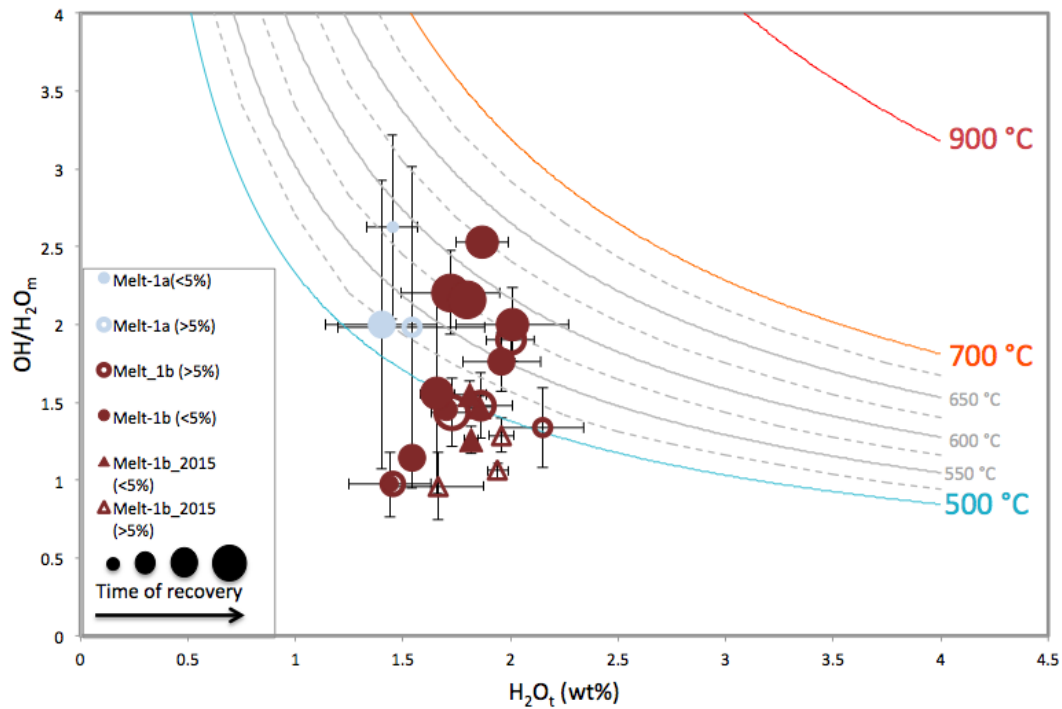
**Fig 6.5: Schematic drawing of the melt properties prior to drilling.** Melting of the felsite forms an undersaturated melt (Melt-1a) which then dissolves the  $\text{CO}_2$  from deep sources.

## 6.3 Temporal variation of temperature and pressure

The major element measurements and water content measurements were used to estimate the quenching temperatures of the melt. For this project, two methods were used to calculate the quenching temperature; the model of Nowak and Behrens (2002) and the model of Ihinger et al. (1999). The model of Nowak and Behrens (2002) was used as it considers the major element and H<sub>2</sub>O content of the glass cuttings. The model of Ihinger et al. (1999) uses the absorbance of the H<sub>2</sub>O<sub>m</sub> and OH bands in the near-IR region (chapter 4; section 4.2.4.1). This model was used in this project because it is the model Zierenberg et al. (2013) used to estimate the quenching temperature and thus will allow temperatures estimated in this study to be directly compared to the quench temperatures estimated by Zierenberg et al. (2013).

### 6.3.1 Nowak and Behrens (2001) model

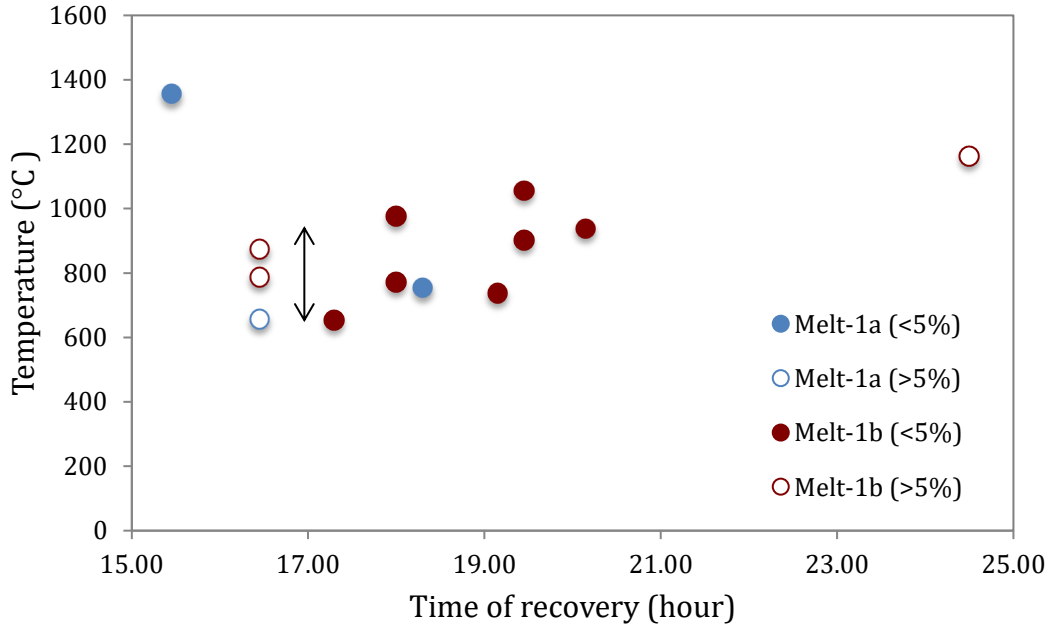
Fig 6.6 shows the equilibrium speciation temperature of the Krafla melt from the model of Nowak and Behrens (2001). To avoid uncertainties on the thickness of the glass wafers quench temperature contours were calculated by using the ration of OH/H<sub>2</sub>O<sub>m</sub>. Melt-1a shows a decrease from ~575 to ~525 °C over time. Melt-1b shows an increase in temperature over time from <500 to ~625 °C. However, at 00:50 the equilibrium speciation temperature decreases to <500 °C.



**Figure 6.6: Equilibrium speciation temperature of the glass cuttings.** Equilibrium speciation temperature contour calculated from Nowak and Behrens (2001). Lowest possible quench temperature contour is 500 °C.

### 6.3.2 Ihinger et al. (1999) model

Fig 6.7 shows the quench temperatures calculated using the Ihinger et al. (1999) model. It appears as if the quench temperature of Melt-1a is lower than the quench temperature of Melt-1b apart from the glass cutting recovered at 15:45, however, this may again reflect the limited data for Melt-1a. Both Melt-1a and Melt-1b show an increase in quench temperature over time from ~600 °C to ~1400 °C. Quench temperatures estimates from Zierenberg et al. (2013) fit within the increasing quench temperature trend.



**Figure 6.7: Quench temperature of the Krafla glass cutting over time of glass recovery.** Quench temperature calculated using a model of Ihinger et al. (1999). Arrow shows calculated quench temperature range from Zierenberg et al. (2013) calculated using the same model.

### 6.3.3 Comparisons between the two models

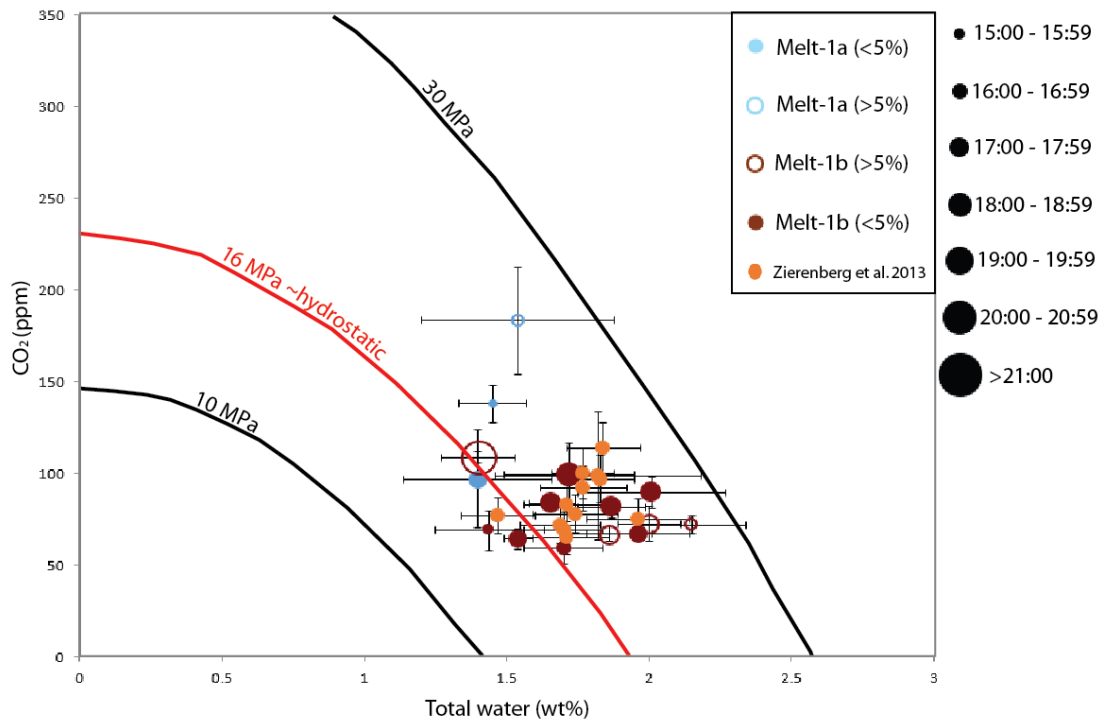
The quench temperature from the Nowak and Behrens (2001) model are much lower than the quench temperature obtained from the Ihinger et al. (1999) model. Both models indicate that the first recovered sample at 1545 has a higher quench temperature than 16:45 and that quench temperatures for Melt-1b increase with time.

## 6.4 H<sub>2</sub>O-CO<sub>2</sub> saturation pressure

Pressures at which the H<sub>2</sub>O-CO<sub>2</sub> contents of each glass sample were saturated were calculated using VolatileCalc (Newman and Lowenstern, 2002). The lowest temperature that VolatileCalc is calibrated for is 600 °C. Temperatures calculated by the Nowak and Behrens (2001) model for all but one of the samples are below this. Thus, the saturation pressures were calculated at 600 °C (Fig 6.8). Most glass cuttings have H<sub>2</sub>O-CO<sub>2</sub> saturation pressures between 15 and 25 MPa. Samples 1645B, 18:00B and 18:30C have saturation pressures below the hydrostatic pressure (~16 MPa) and the rest are above.



The H<sub>2</sub>O-CO<sub>2</sub> saturation pressure of Melt-1b increases over time. No trend is observed with H<sub>2</sub>O-CO<sub>2</sub> saturation pressure of Melt-1a over time, this may reflect the limited data points.

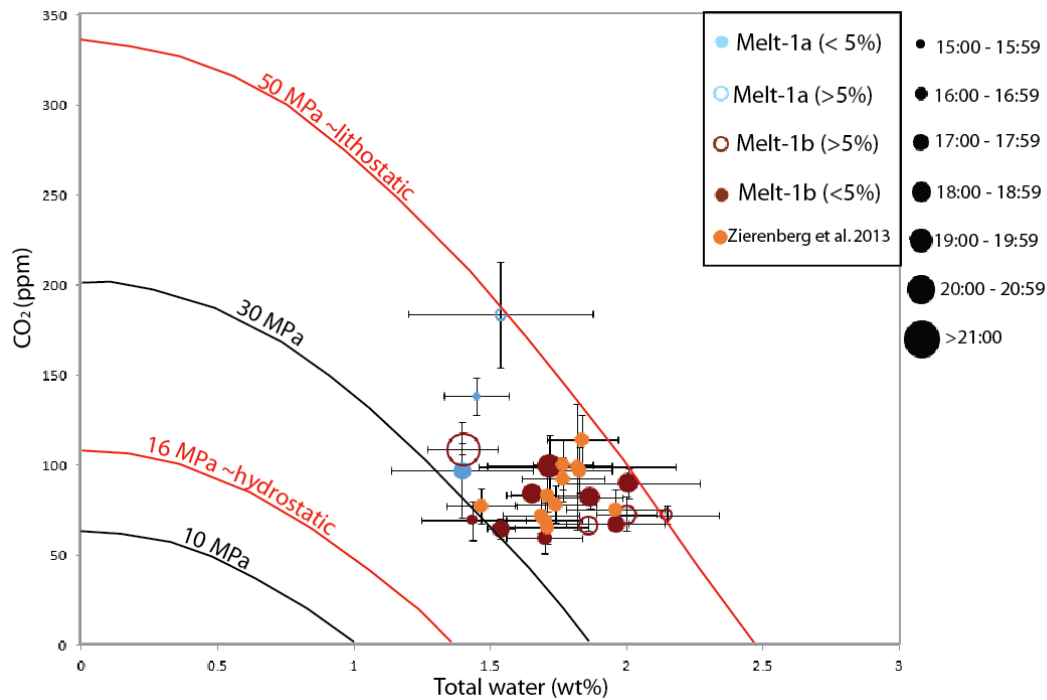


**Figure 6.8: Plot of total H<sub>2</sub>O vs CO<sub>2</sub>.** Saturation pressure contours calculated at 600 °C using VolatileCalc (Newman and Lowenstern, 2002).

Using the quench temperatures estimates using the Ihinger et al. (1999) model, H<sub>2</sub>O-CO<sub>2</sub> saturation pressures at 900 °C were calculated using VolatileCalc (Newman and Lowenstern, 2002). Quench temperature of 900 °C was used because Zierenberg et al. (2013) used this model assuming a temperature of 900 °C. H<sub>2</sub>O-CO<sub>2</sub> saturation pressures from this project can be directly compared to Zierenberg et al's (2013) H<sub>2</sub>O-CO<sub>2</sub> saturation pressures estimates. Fig 6.9 shows the saturation pressures calculated at 900 °C. The glass cuttings are mostly between ~30 and 50 MPa. These pressures are still below lithostatic pressure of ~50 MPa, but all glass cuttings are greater than the hydrostatic pressures of ~16 MPa when using a quench temperature of 900 °C. H<sub>2</sub>O-CO<sub>2</sub> saturation pressures from Zierenberg et al. (2013) ranged from 35 - 45 MPa. If correlation between H<sub>2</sub>O-CO<sub>2</sub> saturation pressures and time are expected it is expected that Zierenberg et al's (2013) points would plot in single point within the trend. This

suggests that there is no systematic correlation between H<sub>2</sub>O-CO<sub>2</sub> saturation pressures and time.

In summary, the lower temperature VolatileCalc model indicates lower pressures, and potentially that Melt-1b came from increasingly higher pressures and possibly deeper depths.



**Figure 6.9: Plot of total H<sub>2</sub>O versus CO<sub>2</sub>.** Saturation pressure contour at 900 °C calculated using VolatileCalc (Newman and Lowenstern, 2002).

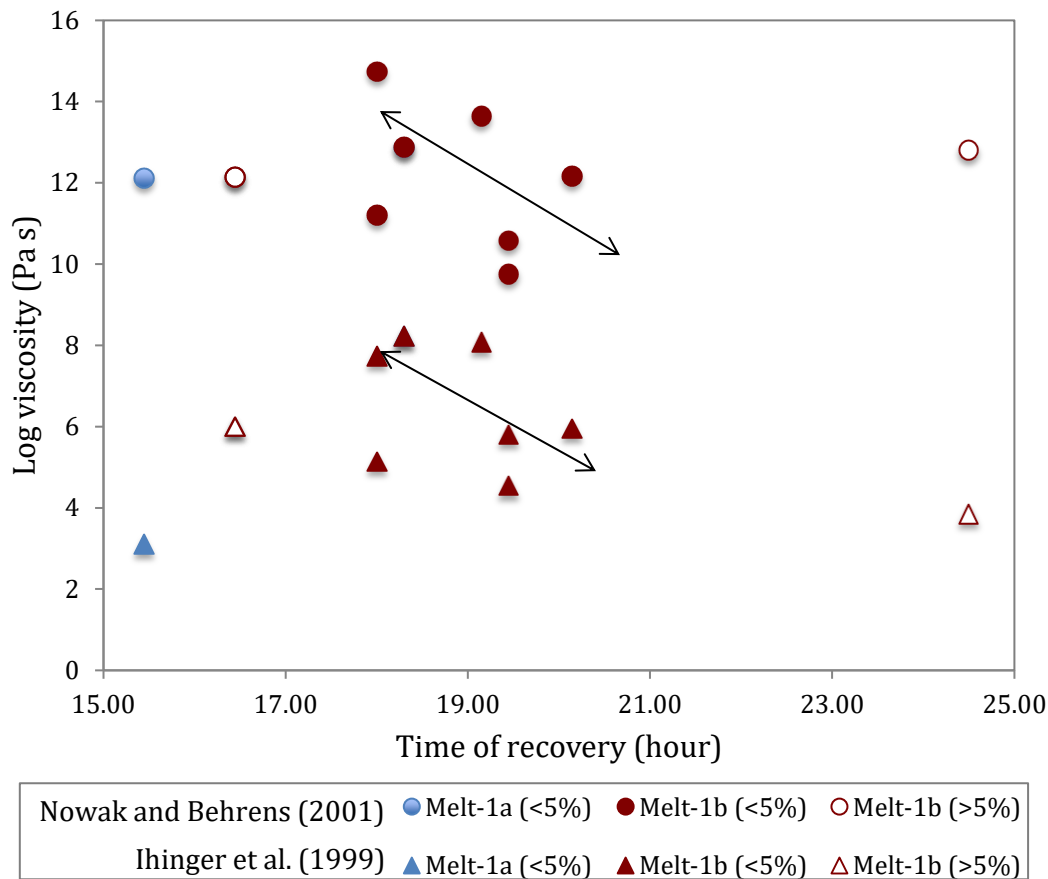
## 6.5 Viscosity and glass transition temperature

Using the temperatures calculated from the two quench temperature models viscosity was calculated using a model of Giordano et al. (2008). Fig 6.10 shows the change in melt viscosity (using parameters from both Nowak and Behrens (2001) and Ihinger et al. (1999) models) over time. The lower temperatures of Nowak and Behrens (2001) produce a much higher viscosity estimate compared to Ihinger (1999). Viscosity estimates using both Nowak and Behrens (2001) model and Ihinger et al. (1999) model shows that earlier recovered melt has relatively lower viscosity and from time intervals ~18:00 to ~20:15 the viscosity of melt-1b decreases due to the increase in temperature.

There seems to be a relationship between viscosity and vesicularity where glass cutting with high vesicularity has relatively lower viscosity than the dense glass. This lower viscosity could reflect higher diffusion rates facilitating bubble nucleation and growth, however this may reflect the limited data for this group.

Glass transition temperatures were calculated using the viscosity, assuming the glass transition temperature is the temperature at which the melt is at a viscosity of  $10^{12}$  Pa s (Chapter 3; Section 3.6.1). The average glass transition temperature of Melt-1a is  $557 \pm 2$  °C and the average glass transition temperature of Melt-1b is  $528 \pm 15$  °C. Table 6.1 summarises the glass transition temperature of the Krafla glass cuttings based on their calculated viscosity. The glass transition temperature is consistent with the quench temperature calculation from the Nowak and Behrens (2001) model and the felsite temperature estimates ( $\sim 500$  °C) from the Horner plot (Chapter 2, section 2.4.3; Morstensen et al. 2014). However, this may not reflect the actual quench temperature of the melt because of the quench effect (Chapter 3; Section 3.4.2).

In summary, the viscosity of the melts are very high, especially if the Novak and Behrens (2001) model is calculating an equilibrium temperature just prior to drilling induced quenching. However, an imperfect quench may mean that these temperatures are minimums, and these viscosities are maximums. But this would contribute to explaining why the magma did not significantly vesiculate or rise up the borehole to the surface.



**Figure 6.10: Viscosity of the Krafla melt over time.** Circle represents viscosity using specific temperatures for each sample from the model of Ihinger et al. (1999) and the triangle represents the viscosity using estimated temperatures for each samples from the model of Nowak and Behrens (2001). Both models show a decrease in viscosity with a low vesiculated Melt-1b. Viscosity calculated using a model of Giordano et al. (2008)

**Table 6.1: Average glass transition temperature of the Krafla melt based on viscosity ( $\sim 10^{12}$  Pa·s)**

	n	Tg (°C)	STDEV
Melt-1a	2	557	2
Melt-1b	13	528	15
All samples	15	532	17

## 6.6 Bubble behaviour in the Krafla magma

Volatile distribution maps from this project have identified three different groups (Chapter 5 section 5.5).

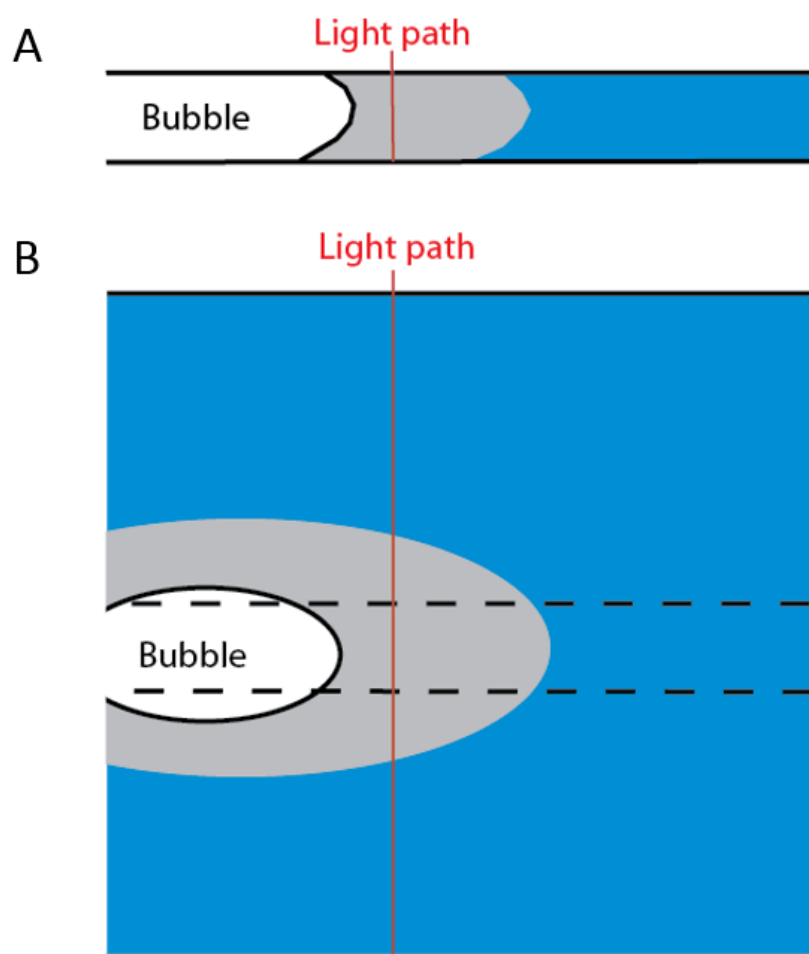
- The first group shows that the volatile distribution is constant towards a bubble. This suggests that the melt is at equilibrium (i.e. no change in pressure and/or temperature) (Chapter 3 section 3.9.3).
- The second group have shown  $H_2O_t$  and OH distribution decreases towards a bubble but the  $H_2O_m$  increases. This suggests initially the melt was depleted in OH and total water close to the bubble due to diffusion towards the bubble and bubble nucleation and growth. However the increasing  $H_2O_m$  indicates that a decrease in temperature may have driven some water resorption and enriched the melt in  $H_2O_m$ . As a result both, the original OH distribution is preserved and the resorbed  $H_2O_m$  also preserved representing the new melt condition (Chapter 3 section 3.9.3).
- The third group have shown the distribution of magmatic volatile (including  $H_2O_m$  and  $CO_2$ ) increases towards the bubble. Suggesting that the bubble is undergoing resorption, indicating that there is an increase in pressure and decrease in temperature (Chapter 3 section 3.9.3).

Different groups can also occur in the same sample and in the same bubble suggesting that the melt is preserved in a state of locally variable disequilibrium and bubbles were behaving heterogeneously. All of the glass cuttings showed higher concentration of OH than  $H_2O_m$  this suggests that throughout time the melt quenches relatively rapidly (McIntosh et al. 2014) freezing in the state of disequilibrium. This relatively rapid quench is also supported by the preservation of highly elongate shaped bubbles, which preserves shearing in the magma present at the time of quenching. However, we emphasize that the spatial variation in volatiles is very small and is close to the error of the measurements.

Possible reasons why maybe the spatial variation in volatile distribution is relatively small could be:

1. In thick samples the Far-field (average concentration of volatile distribution with no influence on bubbles and crystals) concentration overwhelms the area affected by the bubbles (Fig 6.11). Thicker samples also means that the beam light is more likely to go through areas which may have effected area

2. Bubbles may also be too small. Small bubbles (find average bubble size) effect small area than the larger bubbles and perhaps wasn't detectable.



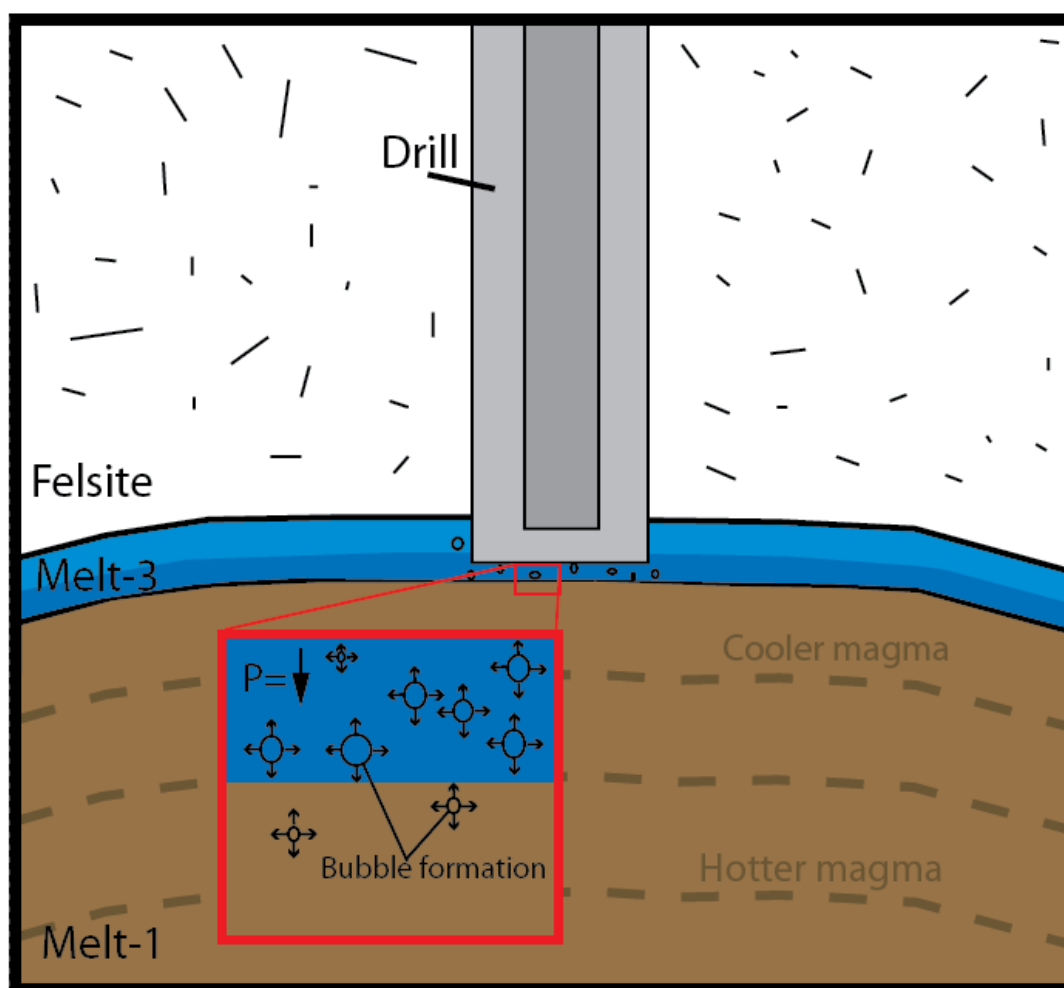
**Figure 6.11: Schematic diagram showing how sample thickness can influence recorded volatile measurements.** In thin samples (A) where the light sensor can directly measure near the edges of bubbles (grey region) there is little influence from the far-field (blue region) bulk volatile concentration. In thicker samples where the bubble and near regions are fully encased by the far-field (B), measured volatile concentrations can be skewed by the surrounding far-field bulk concentration.

## 6.7 Magmatic response

### 6.7.1 Initial response (15:00 – 17:00)

Fig 6.12 illustrates the initial magmatic response from the intercepted drill. Evidence of depressurisation is present in the earlier recovered Melt-1a glass cuttings. These include high abundance of bubbles in Melt-1a (Zierenberg et al. 2013; Saubin et al.

2017). However, this magmatic response was short in duration as at ~17:00 the abundance of bubble rich Melt-1a decreased and the recovered glass cuttings became dense (Saubin et al. 2017). This suggests that depressurisation from the drilling only affected the magma for a short interval before returning back into equilibrium. The short depressurisation period suggest that the drill only affected a small area of the magma and thus the depressurisation of the melt was not great enough or long enough in induce an eruption. Another reason why eruption did not occur is that the melt viscosity was too high to allow significant magma chamber scale bubble nucleation and growth sufficient enough to induce an eruption. It is also worth considering that the pressure in drill hole is variable (Sabin et al. 2017), but also controlled by the drillers who intentionally try to minimise large pressurisation or depressurisation events. This additional control by the drillers may have prevented an eruption if magmas had been suddenly decompressed to atmospheric conditions.



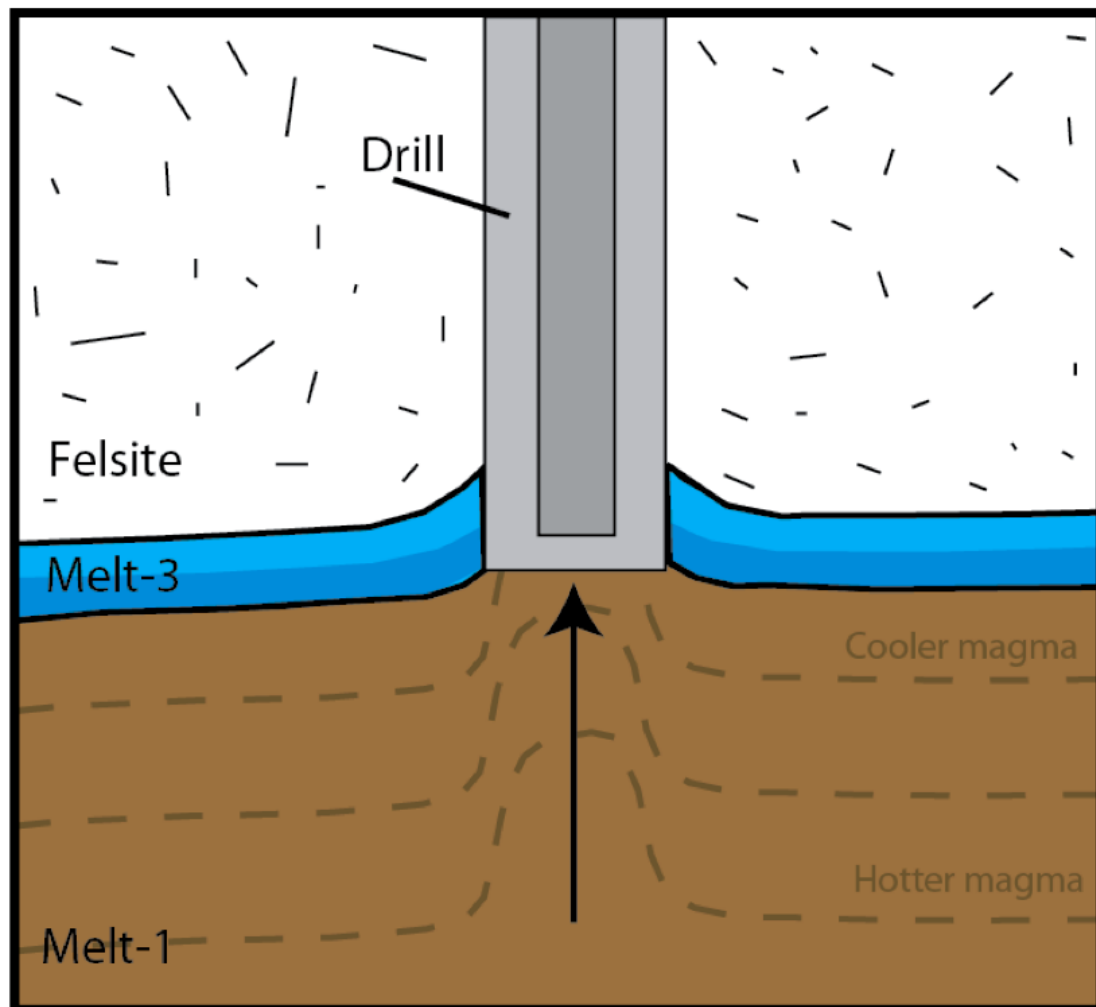
**Figure 6.12: Schematic illustration of the initial response of the magma interception.** As a response to depressurisation from the drill, bubble formation occurred in Melt-1a (Melt-3)

### 6.7.2 Later response (>17:00)

Figure 6.13 illustrates the later magmatic response. An increase in the quenching temperature (Fig 6.6 and Fig. 6.7) and corresponding decrease in viscosity (Fig 3.10) from both models from Nowak and Behren (2001) and Ihinger et al. (1999) indicated hotter magma rising up the borehole. The H<sub>2</sub>O-CO<sub>2</sub> saturation pressures models also show a small increase in pressure (difference of ~15 MPa assuming temperature is 600 °C or difference of ~20 MPa assuming temperature is 900 °C) indicating that either the drill didn't penetrate deep within the magma or that the magma did not rise far to meet the drill (a few metres or tens of meters). The lack of bubbles in these Melt-1b samples



indicate that either this melt rapidly quenched before vesiculation was possible or that the melt was understaturated.



**Figure 6.13: Schematic drawing of relatively hot magma rising in the later response to the intersection of the magma body.**

## 6.8 Summary and future research directions

The major element results from this project have indicated that Melt-1a is caused by mixing of the magma body (similar to Melt-1 defined by Zierenberg et al. 2013) and the melt created in the surrounding felsite (similar to Melt-3 defined by Zierenberg et al. 2013). The bubbles in the first returned melt indicate that decompression induced by drilling was relatively localised due to the quick transition back to equilibrium (ie no bubble formation). The increase in temperature and corresponding decrease in viscosity between 16:45 to 20:39, indicates that a hotter magma (Melt-1b) was tapped.

Saturation pressures do not vary significantly suggesting that the melt did not come from very different depths. The low temperatures and viscosities and decompression magnitude described here explain why there was no eruption initiated with this event. The results of this project has indicated an on-going need to investigate further the magmatic response of sudden depressurisation, e.g. if failure of the valves in the drill hole allowed sudden depressurization of the melt. The following outlines potential future work that is required on the Krafla glass cuttings or studies of similar nature:

- Collecting and processing more data especially on the clear glass (boundary melt between the magma body and the surrounding felsite).
- Investigate whether the recovery time of the glass cuttings accurately represents that of the glass at the time of quenching. Due to the nature of the extraction process, it is hard to determine the true timing of glass quenching to the time at which it was recovered from the drill. This variable is important to constrain for similar studies as it can provide further insight into volatile migration during rapid depressurisation events.
- Investigation on how, or if, the previous two intersections of the magma by the drilling have affected the melt prior to the third intersection, it is subsequent to this intersection that the time-series of samples were collected.
- Further studies on whether different bottom hole assemblies (BHA) and size of the drill bit can affect how the intersected magma body responds to the intersection.
- More work on mapping the magmatic volatile distribution around bubbles. From this project the volatile distribution transects have shown both resorption and growth behaviour, however, there are no systematic trends with quench time and sample texture. This could be due to the limited data. A future study of volatile distribution around the bubbles would include looking at different bubble shapes (stretched and round bubbles). As performed in this project, FTIR can be used to measure volatile distribution study, but in addition, a secondary ion mass spectrometry (SIMS) calibrated backscatter scanning electron microscopy (BSEM) (Humphreys et al. 2008; McIntosh et al. 2014) could also be used to investigate H<sub>2</sub>O distribution around bubbles.

- Investigation on what causes the doublet peak ( $\text{CO}_2$ ) observed in the thin sample from SR-FTIR analysis (Chapter 4 section 4.2.4.1.1). Understanding how the doublet peak may or may not be related to measurement methods will provide a more accurate measurement of  $\text{CO}_2$  and therefore, more precise estimates on the  $\text{H}_2\text{O}$ - $\text{CO}_2$  saturation pressure.

## Chapter 7: Conclusion

The unexpected magma drilling of the IDDP-1 drill project has provided research opportunities for investigating subsurface magma and future geothermal studies. The aim of this project was to understand how the Krafla melt responded to being intercepted by the IDDP-1 drilling project. Major element data from EPMA analysis have shown that prior to drilling there was a complex melt boundary around the magma body which was a result of the melting and mixing of the surrounding felsite with the magma body. Magmatic volatile analyses from FTIR have shown that each melt type has a distinct volatile signature. Melt 1a (interpreted to represent the magma) is typically slightly higher in  $H_2O_t$  and  $H_2O_m$ , while Melt 1b (interpreted to represent the assimilation of the felsite and magma) is higher in  $CO_2$ . Volatiles in Melt 1b changes over time, with slight increases in the concentration of total  $H_2O$ , and a significant increase in  $CO_2$  and  $OH/H_2O_m$ . High  $OH/H_2O_m$  indicates that the melt has rapidly quickly. Bubbles in the first returned melt (Melt-1a) indicate that decompression induced by drilling was relatively slow and local. The increase in quench temperature and decrease of viscosity indicates that hotter magma (Melt-1b) either was progressively tapped or rose into the borehole. The low variation of  $H_2O$ - $CO_2$  saturation pressure over time suggests that either the drill did not penetrate far into the magma or the magma did not rise far into the borehole. Table 7.1 summarises the objectives and their answers for this project.

Table 7.1: Summary of project objectives and the associated answers

Objective	Answer	
1. Determine any changes in magmatic volatile concentration through time	Overall the $H_2O_t$ concentration for the clear glass cuttings (mixing of the felsite and magma) are lower than those of the brown glass cuttings (magma). No significant changes in the concentration of $H_2O_t$ over time. In all glass cutting the concentration of OH is higher than the concentration of $H_2O_m$ . OH/ $H_2O_m$ of the brown glass will increase.  CO <sub>2</sub> concentrations in the brown glass increases over time. No systematic trend in the clear glass.	Chapter 5
2. Map the distribution of H <sub>2</sub> O species and CO <sub>2</sub> around bubbles in the samples IDDP-1 glass cuttings	Volatiles distributions around bubbles have shown three different groups were observed. 1. No change in volatile concentration towards the bubble. Indicating the melt was in equilibrium 2. $H_2O_t$ and OH concentrations decrease towards the bubble and $H_2O_m$ increases over time. Indicating initial bubble growth then the melt experienced a decrease in temperature causing the bubble to resorb. 3. H <sub>2</sub> O speciation and CO <sub>2</sub> concentration increase towards the bubble suggesting bubble resorption  There were no systematic changes between time, glass textures and glass thickness. This indicates that the melt condition was heterogeneous	
3. Determine the magmatic degassing history of the Krafla magma body in response to a sudden loss of pressure as a result of interception by the drilling	Evidence of depressurisation in the earlier clear glass cuttings includes a high abundance of bubbles. However, this magmatic response was short lived. This suggests that depressurisation from the drilling only affected the magma for a short interval.	Chapter 6
4. Use the degassing history to assess quench timing and changes in magmatic viscosity.	The increase in OH/ $H_2O_m$ and CO <sub>2</sub> over time suggests that the quench temperature increased with time and viscosity decreased with time. This indicates hotter magma rising. The H <sub>2</sub> O-CO <sub>2</sub> saturation pressure model also shows a small increase in pressure, suggesting that either the drill didn't penetrate deep within the magma or that the magma did not rise far.	
5. why didn't the Krafla magma erupt while it was being drilled	Overall, the effect from the drilling is small. The short depressurisation period suggests that the drill only affected a small area of the magma and therefore depressurisation of the melt was not great enough to induce an eruption.	

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# Appendix 1: Standards testing for EPMA

**Table A1.1: Table of expected values for VG-568 standard and the obtained values from the EPMA for this project.**

<b>VG-568</b>	<b>Expected value</b>	<b>STDEV</b>	<b>n</b>	<b>EPMA value</b>	<b>STDEV</b>	<b>n</b>
<b>Na<sub>2</sub>O</b>	3.74	0.34	50	3.36	0.12	10
<b>MgO</b>	0.04	0.02	50	0.03	0.01	10
<b>Al<sub>2</sub>O<sub>3</sub></b>	12.14	0.17	50	12.28	0.1	10
<b>SiO<sub>2</sub></b>	76.77	0.42	50	77.35	0.2	10
<b>K<sub>2</sub>O</b>	4.97	0.09	50	4.84	0.13	10
<b>CaO</b>	0.42	0.03	50	0.46	0.02	10
<b>TiO<sub>2</sub></b>	0.08	0.05	50	0.06	0.01	10
<b>MnO</b>	0.03	0.02	50	0.02	0.02	10
<b>FeO</b>	1.11	0.09	50	1.16	0.07	10

**Table A1.2: Table of expected values for VGA-99 standard and the obtained values from the EPMA for this project.**

<b>VG-A99</b>	<b>Expected value</b>	<b>STDEV</b>	<b>n</b>	<b>EPMA value</b>	<b>STDEV</b>	<b>n</b>
<b>Na<sub>2</sub>O</b>	2.70	0.04	13	2.45	0.07	10
<b>MgO</b>	5.10	0.08	13	4.92	0.07	10
<b>Al<sub>2</sub>O<sub>3</sub></b>	12.49	0.13	13	12.45	0.09	10
<b>SiO<sub>2</sub></b>	51.00	0.27	13	51.01	0.1	10
<b>K<sub>2</sub>O</b>	0.86	0.05	13	0.85	0.02	10
<b>CaO</b>	9.21	0.11	13	9.1	0.12	10
<b>TiO<sub>2</sub></b>	4.10	0.03	13	3.99	0.03	10
<b>MnO</b>	0.20	0.02	13	0.2	0.02	10
<b>FeO</b>	13.24	0.1	13	13.56	0.15	10

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## Appendix 2: Volatile distribution transects

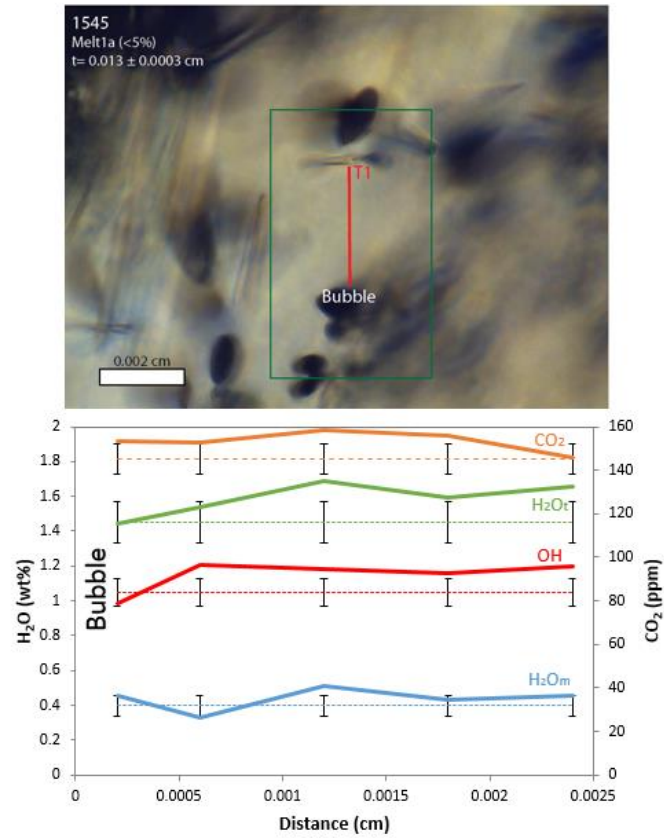
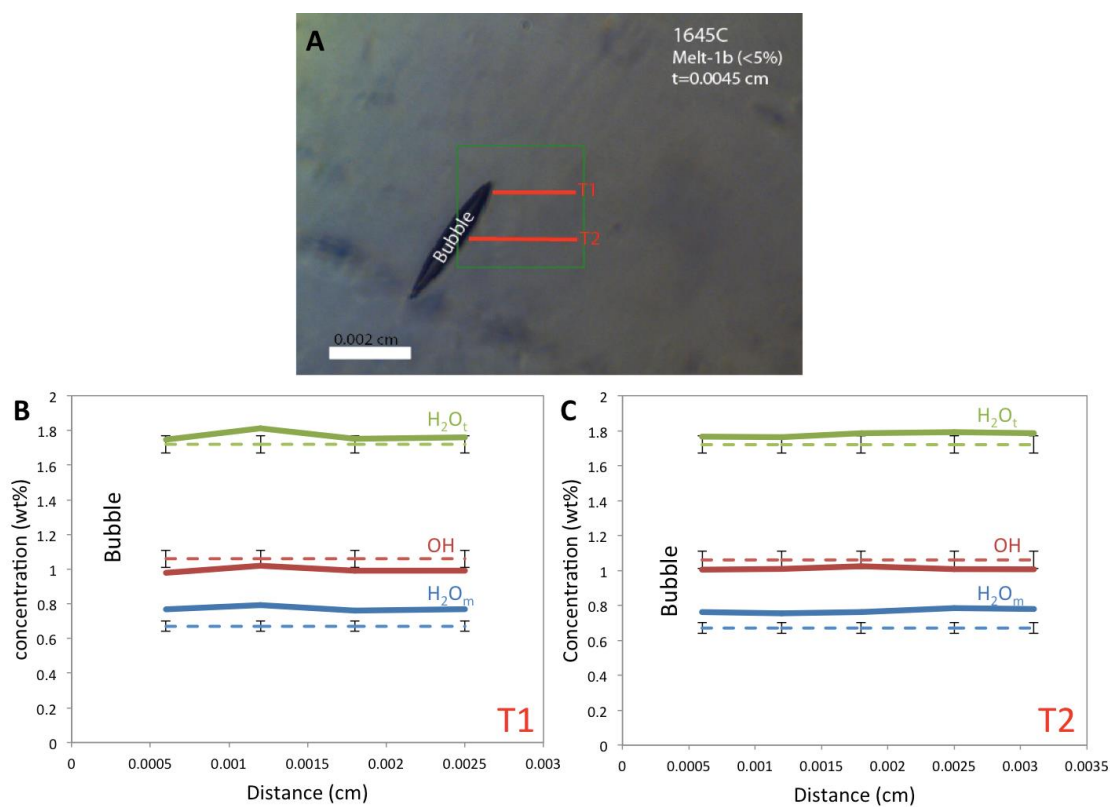
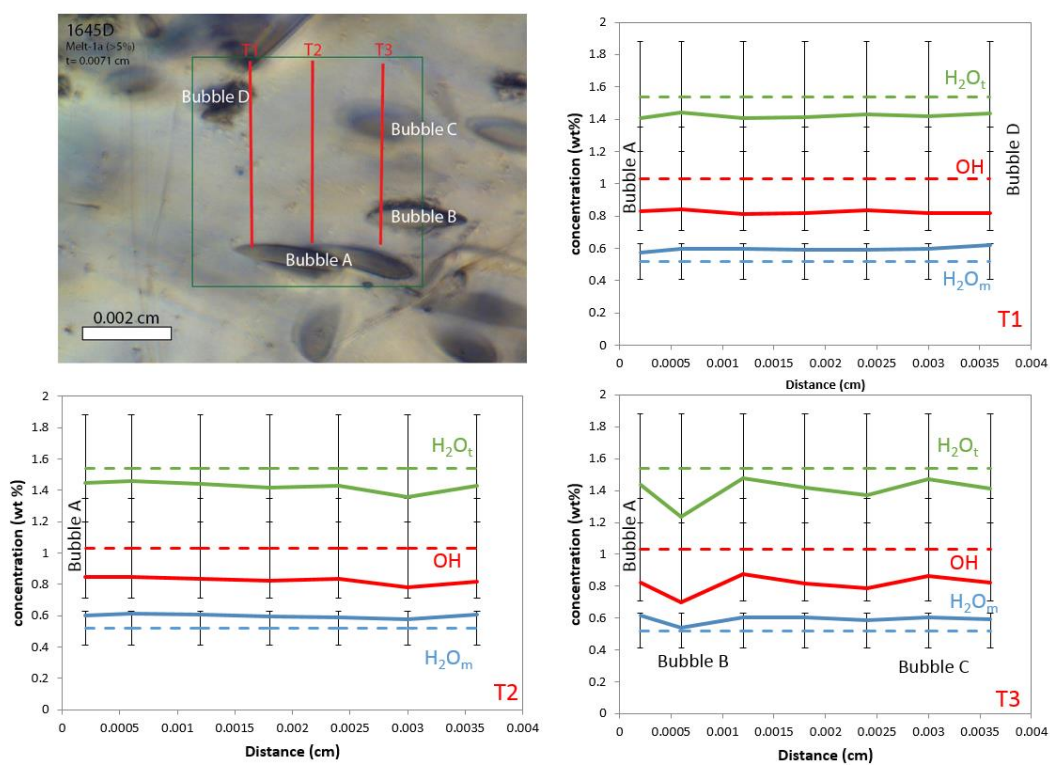


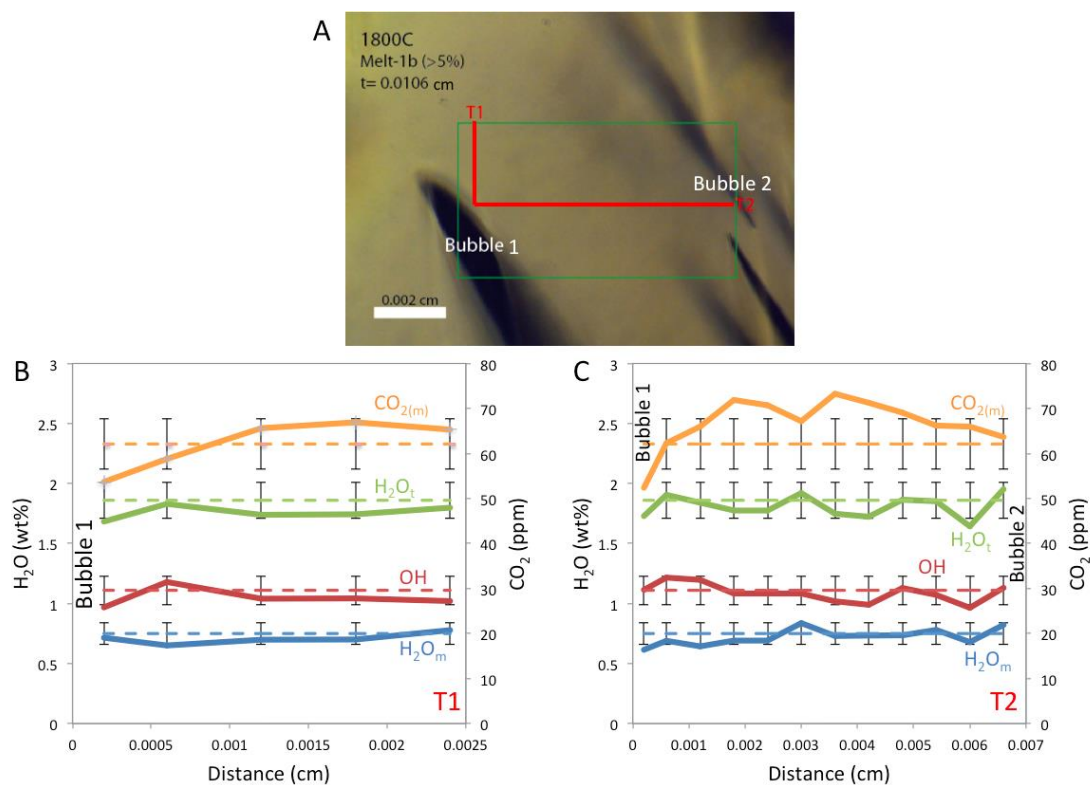
Figure A2.2:  $H_2O$  and  $CO_2$  transects in sample 1545



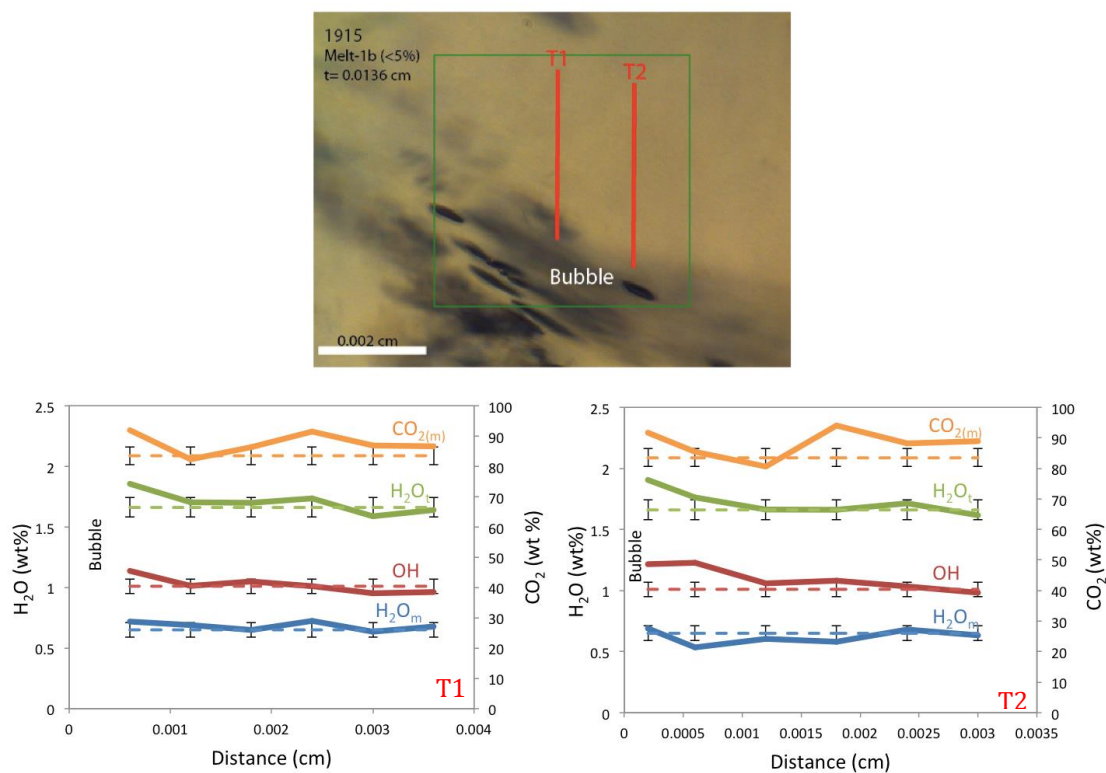
**Figure A2.2: H<sub>2</sub>O transects in sample 1645C**



**Figure A2.3: H<sub>2</sub>O transects in sample 1645D**

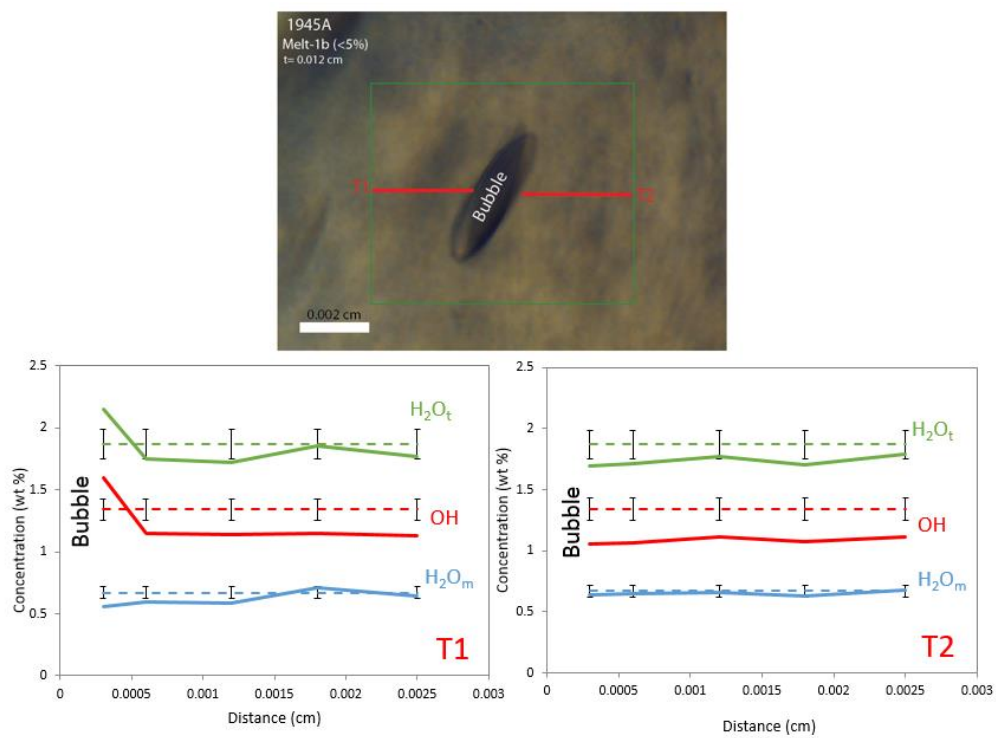


**Figure A2.4:  $\text{H}_2\text{O}$  and  $\text{CO}_2$  transects in sample 1800C**



**Figure A2.5:  $\text{H}_2\text{O}$  and  $\text{CO}_2$  transects in sample 1915**





**Figure A2.6: H<sub>2</sub>O transects in sample 1945A**